

LSR definition

$$(\Pi, \Theta, Z) = (0, V_{circ}, 0)$$

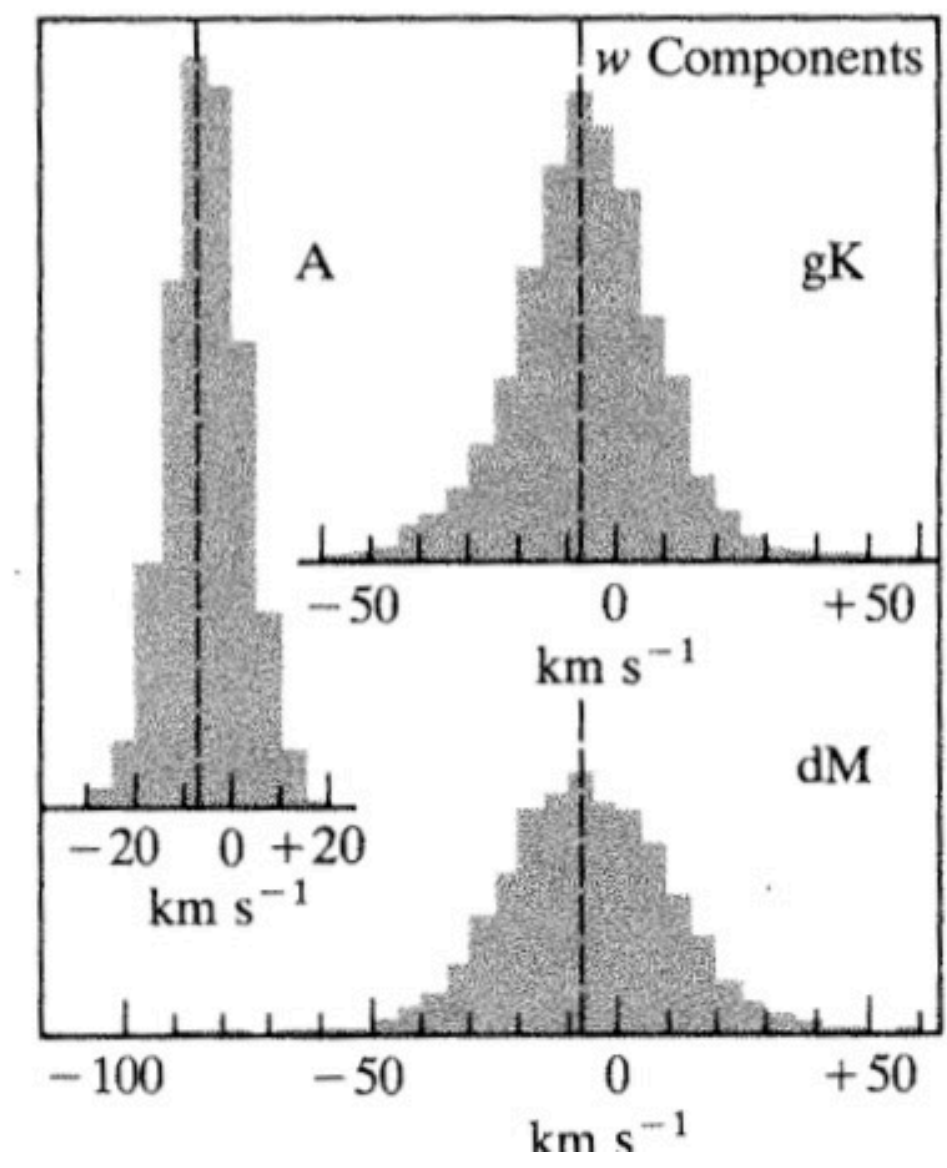
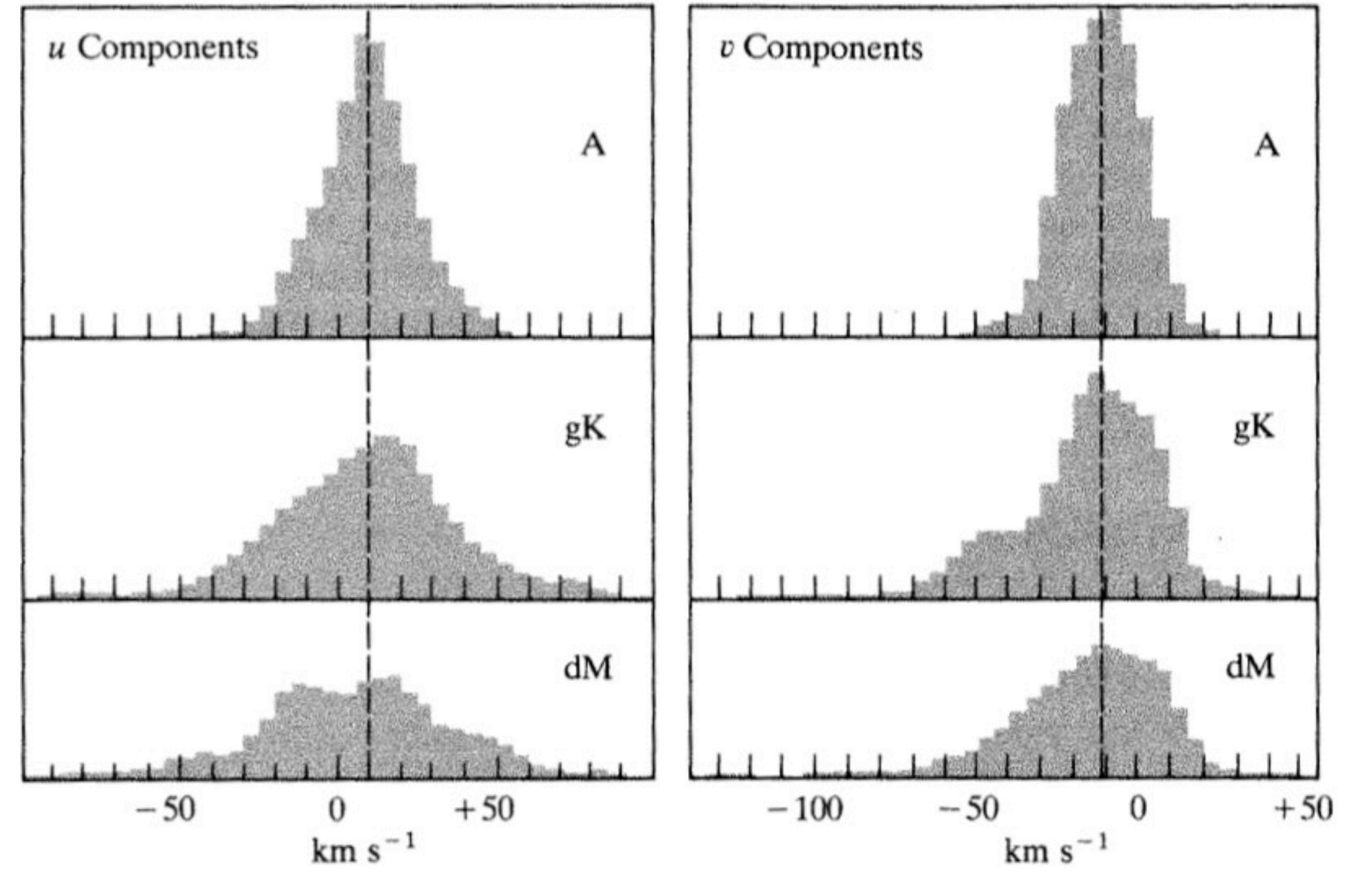
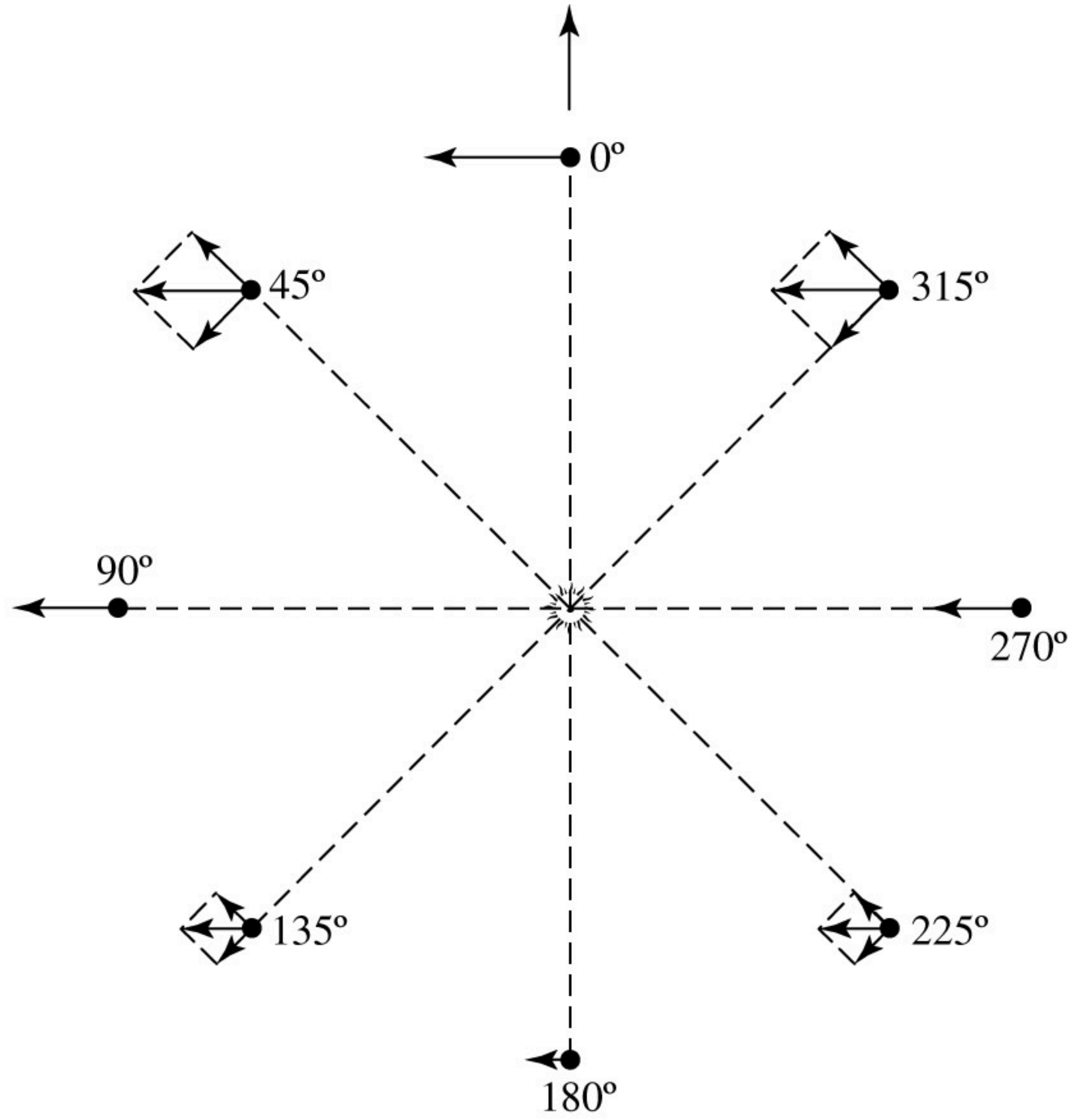
What we really see...

$$(u, v, w) = (\Pi, \Theta - V_{circ}, Z)$$

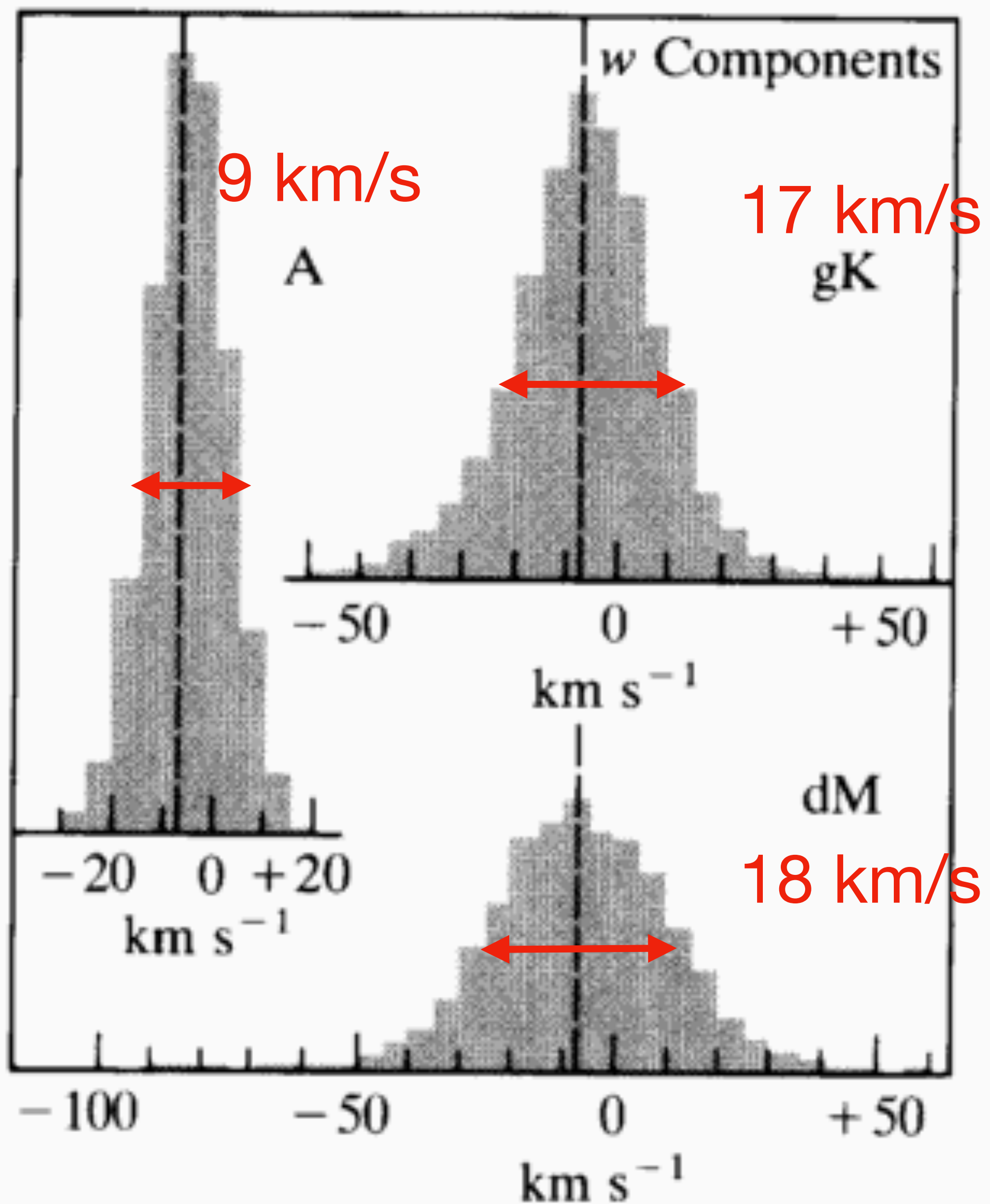
- A: star at apogalacticon, $v < 0$, “lag”
- B: star at perigalacticon, $v > 0$, “lead”
- C: star on circular orbit = LSR, $v = 0$

Sun : $(u, v, w) = (-10, 5, 7)$ km/s

Wait, how do we measure this??



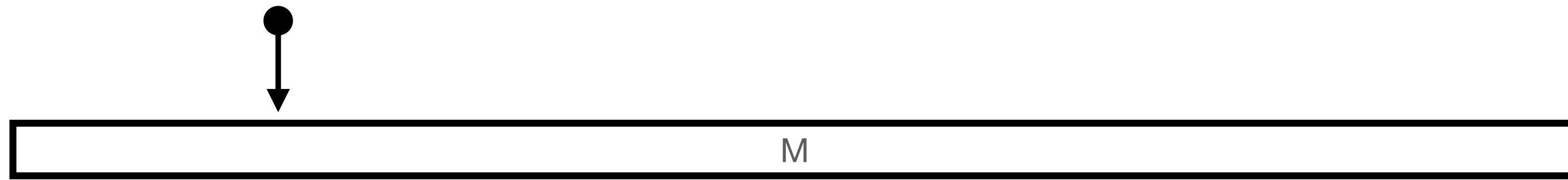
Look specifically at Z motion (w velocity)



width of distributions = “velocity dispersion” σ

Spectral Type	Dispersion (km/s)	Scale Height (pc)
B	6	60
A	9	120
K giant	17	270
M dwarf	18	350
white dwarf	25	500

Why do σ and z increase with spectral type?



balance KE with PE for a small mass orbiting a big mass

$$\frac{1}{2}mv^2 \sim \frac{GMm}{r}$$

$$v^2 \sim \frac{2GM}{r}$$

the big mass M is a disk and has a radius r and surface density Σ_0

$$M \sim \Sigma_0 \pi r^2 \quad \Sigma_0 = (M_\odot/\text{pc}^2)$$

$$v^2 \sim 2\pi G \Sigma_0 r$$

extended to a group of stars

$$\sigma_z^2 \sim 2\pi G \Sigma_0 z_0$$

find surface density from scale heights and velocity dispersions

$$\Sigma_0 \sim 2\pi G \sigma_z^2 z_0$$

Oort Limit
= 75 M_\odot/pc^2

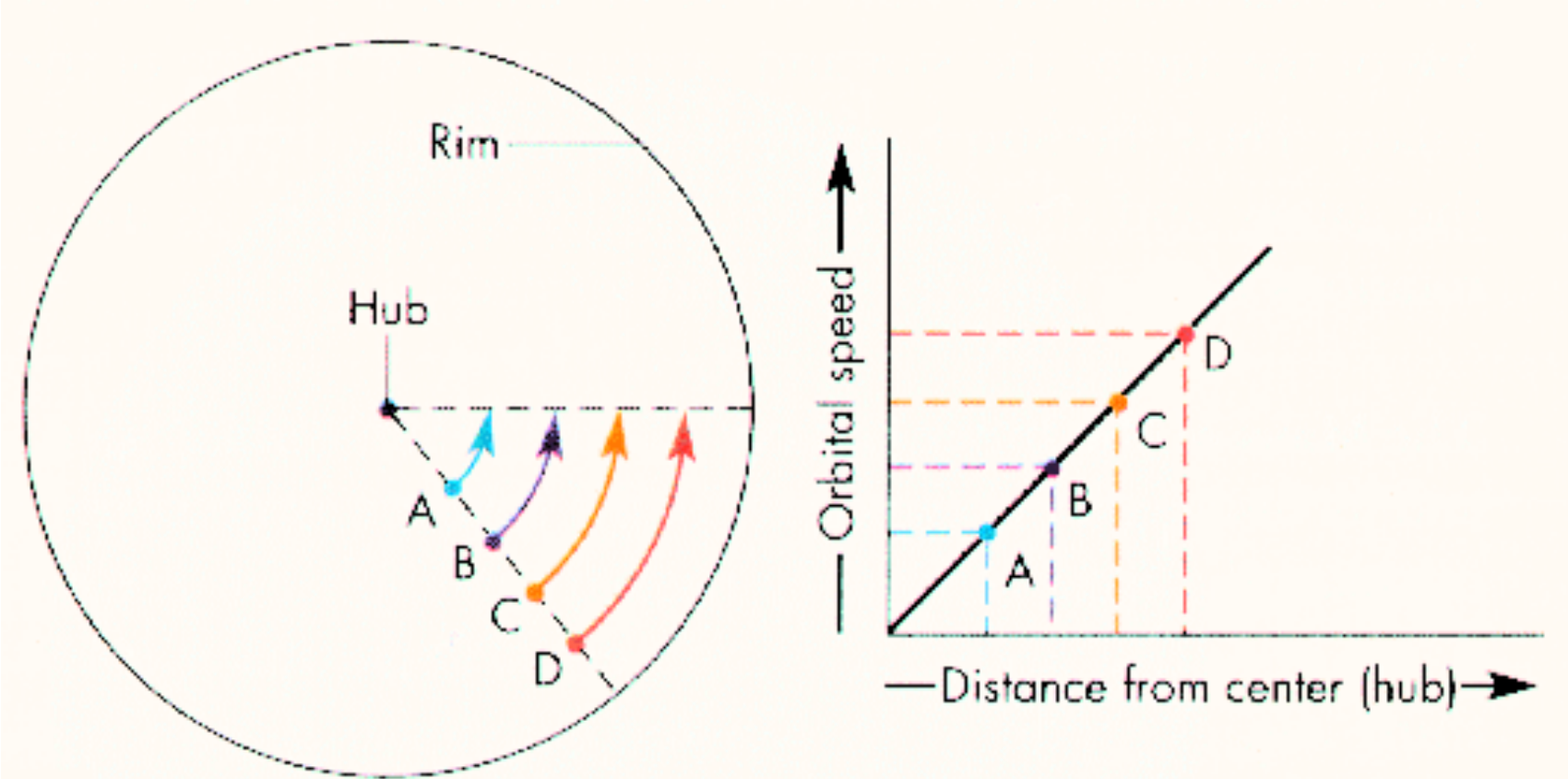
Stars	30 M_{\odot}/pc^2
White dwarfs, neutron stars, black holes	5 M_{\odot}/pc^2
Brown Dwarfs	2 M_{\odot}/pc^2
Gas (H _I + H ₂ + He + molecules)	13 M_{\odot}/pc^2
Total	50 M_{\odot}/pc^2

???

Rotation Curve

a plot of circular (tangential) velocity v. radius

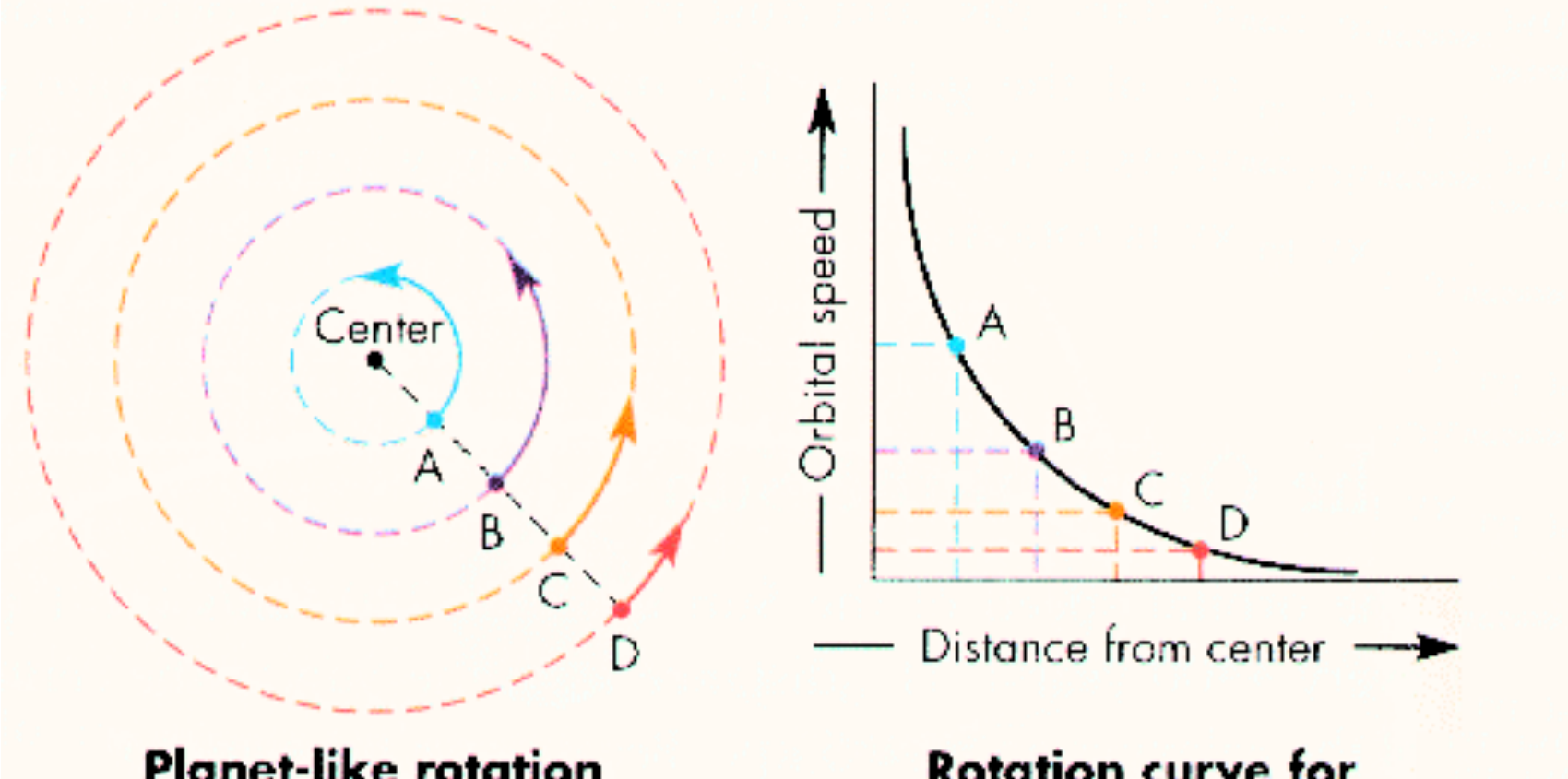
What does the rotation curve of a solid disk look like?



Wheel-like rotation

Rotation curve for wheel-like rotation

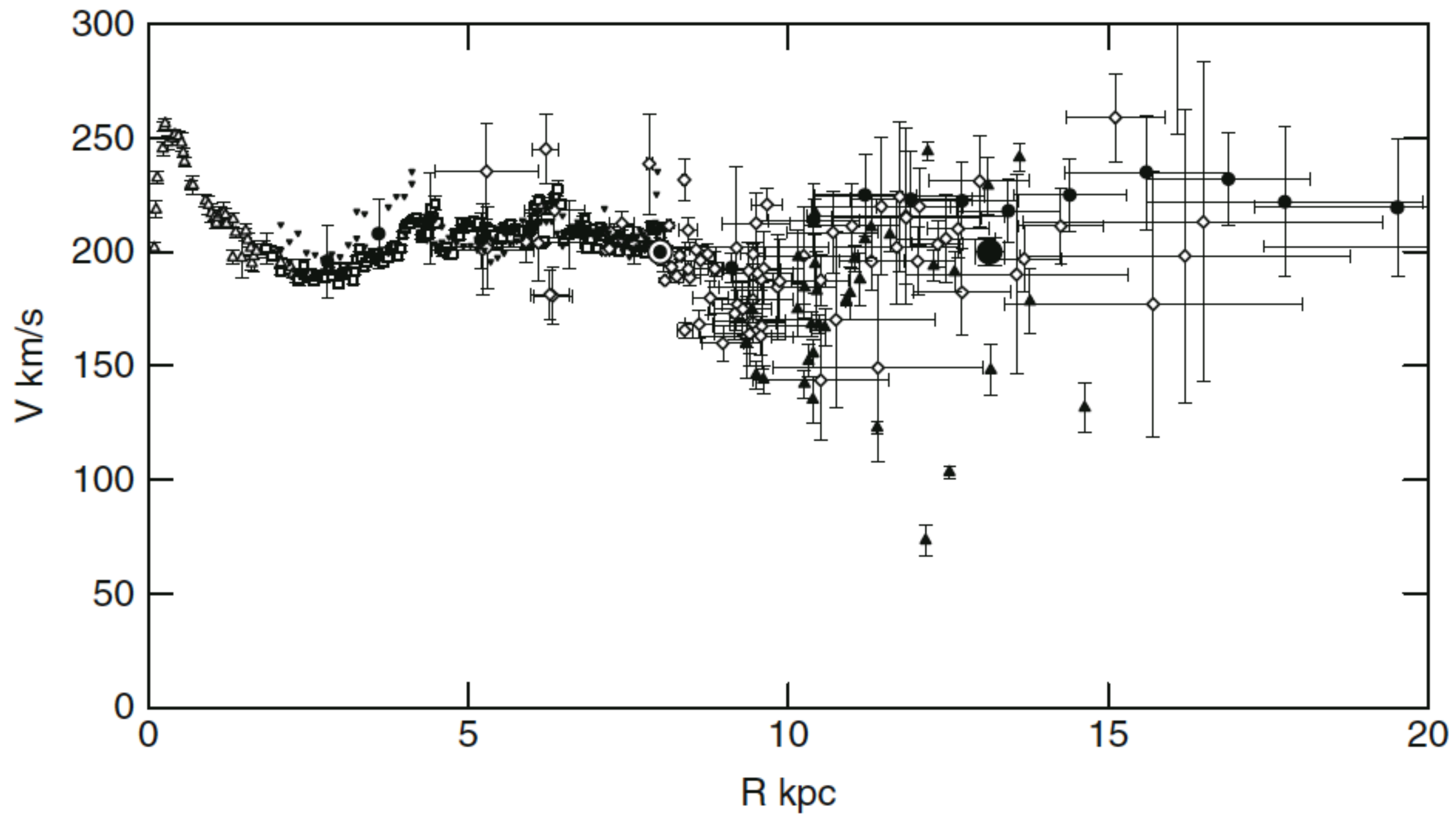
What does the rotation curve of the Solar System look like?

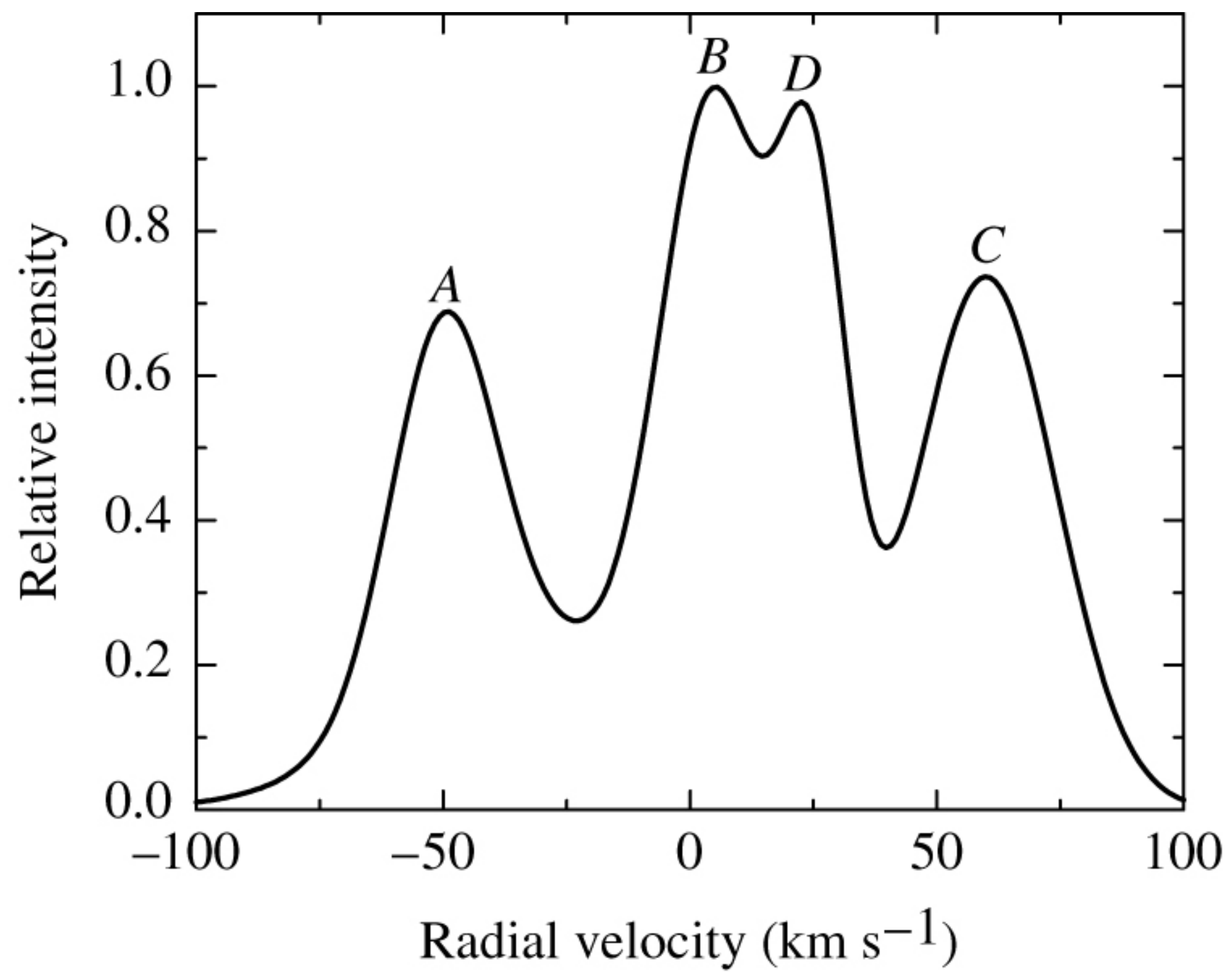


Planet-like rotation

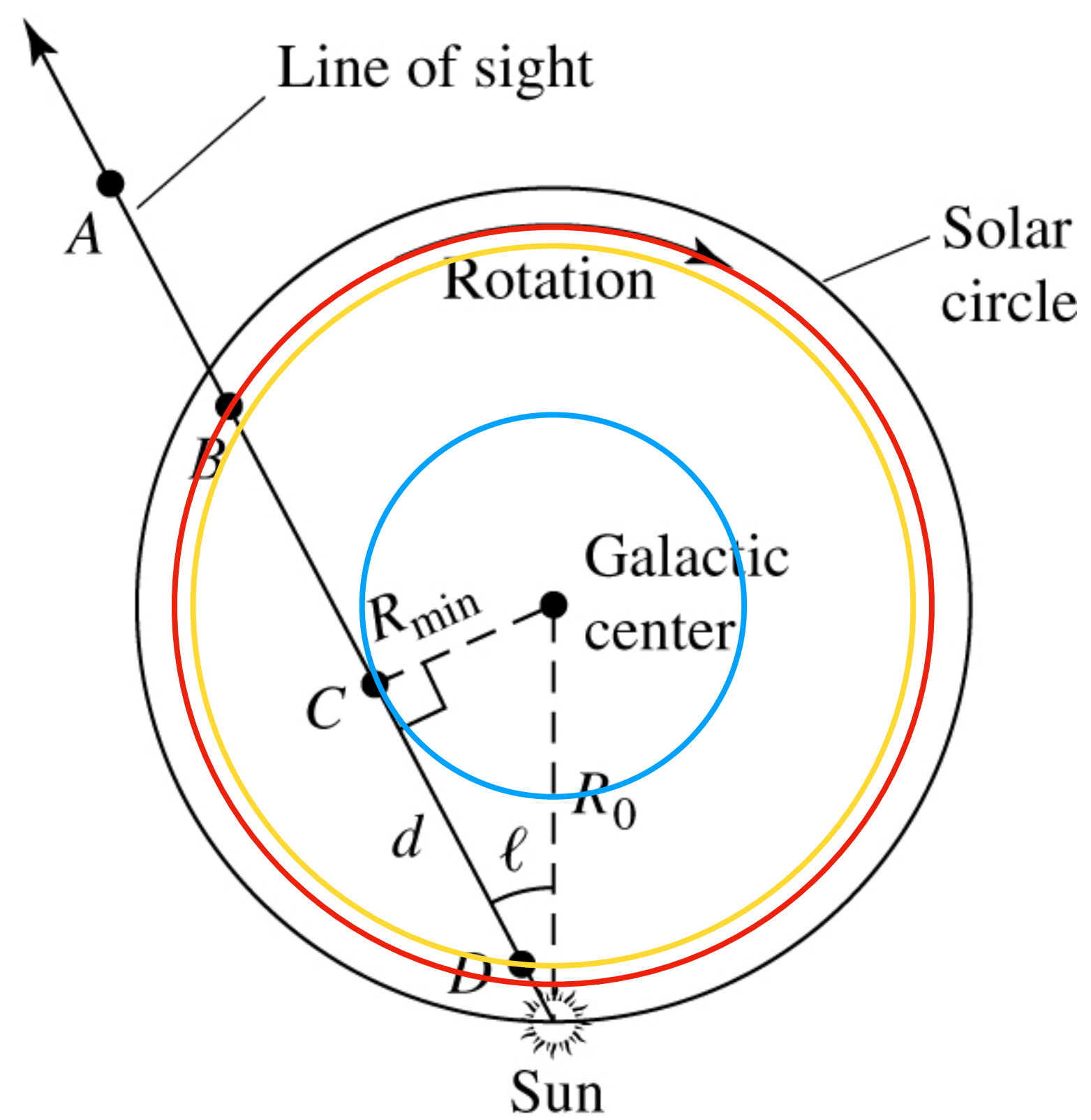
Rotation curve for planet-like rotation

What do you expect the rotation curve of the Milky Way to look like?





(a)



(b)