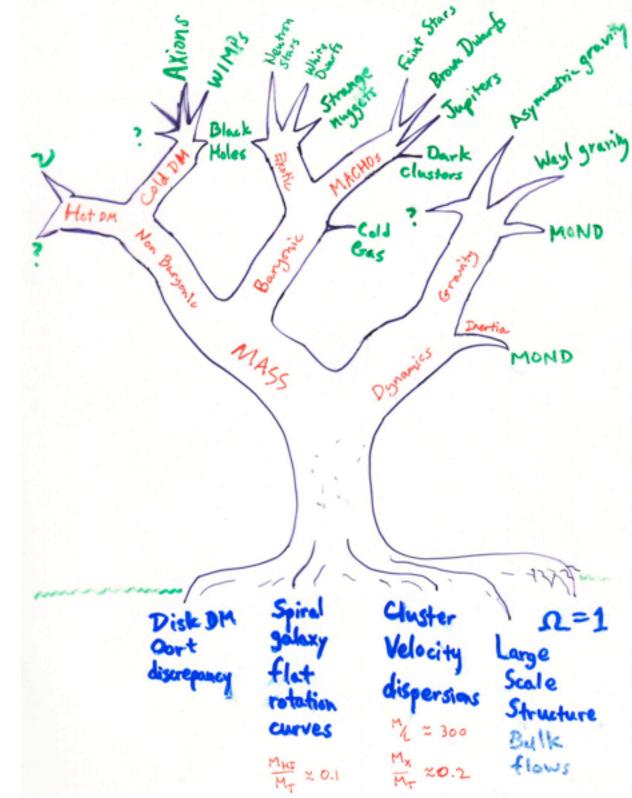
DARK MATTER

ASTR 333/433 FALL 2013 M T 4:00-5:15PM SEARS 552

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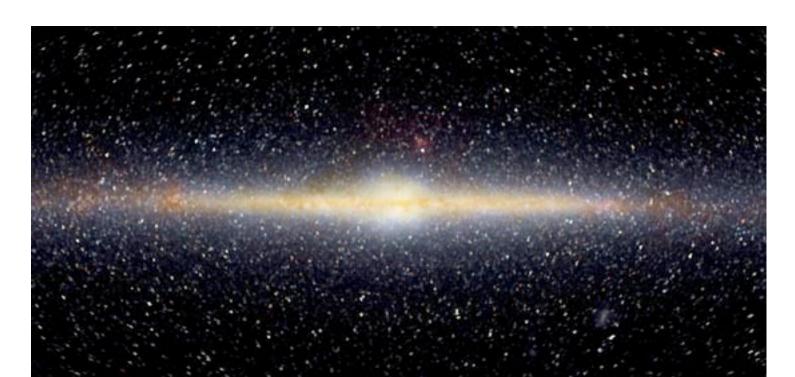
Galactic Kinematics

Galactic constants

$$R_0 \Theta_0 A B$$

$$\Omega < \kappa < \nu_z$$

Epicycle approximation



Galactic Coordinates

from solar system

from Galactic Center

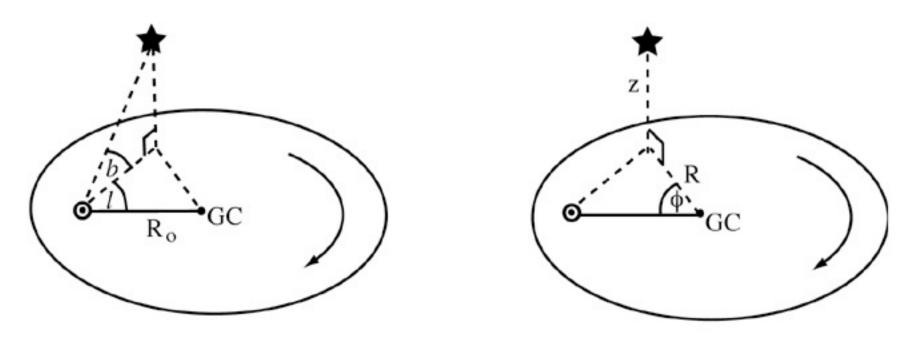


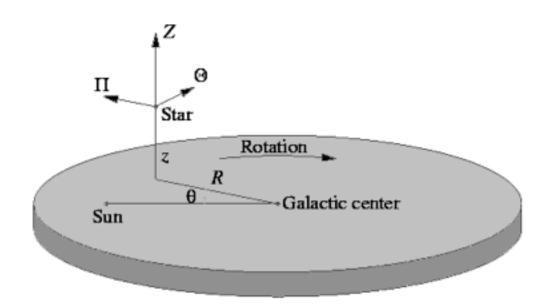
Fig 1.10 'Galaxies in the Universe' Sparke/Gallagher CUP 2007

longitude & latitude ℓ,b

 R, ϕ, z

The Local Standard of Rest

Let's define a coordinate system:



- R = galactocentric distance
- theta = azimuthal coordinate
- z = height above/below the plane

Position : (R, θ, z) Velocity : (Π, Θ, Z)

- Pi = velocity in/out from center
- Theta = tangential velocity
- Z = velocity up and down

LSR - local standard of rest

Define a point in space that is moving on a perfectly circular orbit around the center of the galaxy at the Sun's galactocentric distance. We measure all velocities of stars relative to this point, which is known as the Local Standard of Rest.

 $\Pi_{LSR} = 0$

 $\Theta_{LSR} = \Theta_0$

The velocity of the Local Standard of Rest (LSR) is then given by

$$Z_{LSR} = 0$$

More generally, if the Galactic potential is not axis-symmetric (e.g., because of the Galactic bar), then the LSR orbit is oval.

 R_0 distance to Galactic Center

 Θ_0 orbital velocity of LSR

 $\Omega_0 = rac{\Theta_0}{R_0}$ angular velocity of LSR

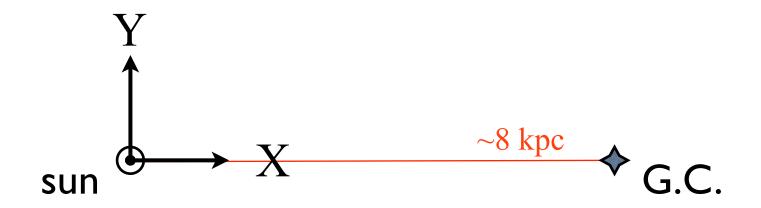
A Oort constant A

B Oort constant B

Local Galactic Coordinates

Cartesian coordinates centered on solar system

As opposed to Galactic Center. Beware sign conventions



X, Y, Z:

X points towards the Galactic Center

Y points in direction of the sun's orbital motion

Z is perpendicular to the Galactic Plane

U, V, W are velocities in these directions

Solar Motion

The residual solar motion wrt the average of local stars is

radial

$$U_{\odot} = 10 \; \mathrm{km} \, \mathrm{s}^{-1}$$

azimuthal

$$V_{\odot} = 12 \; \mathrm{km} \, \mathrm{s}^{-1}$$

Some say V = 5 km/s, some say 15 km/s!

vertical

$$W_{\odot} = 7 \; \rm km \, s^{-1}$$

The Sun is moving

- a bit towards the galactic center
- · faster than the LSR
- northward out of the galactic plane

Currently we are near the mid-plane

(Remember this doesn't account for the rotation of the disk!)