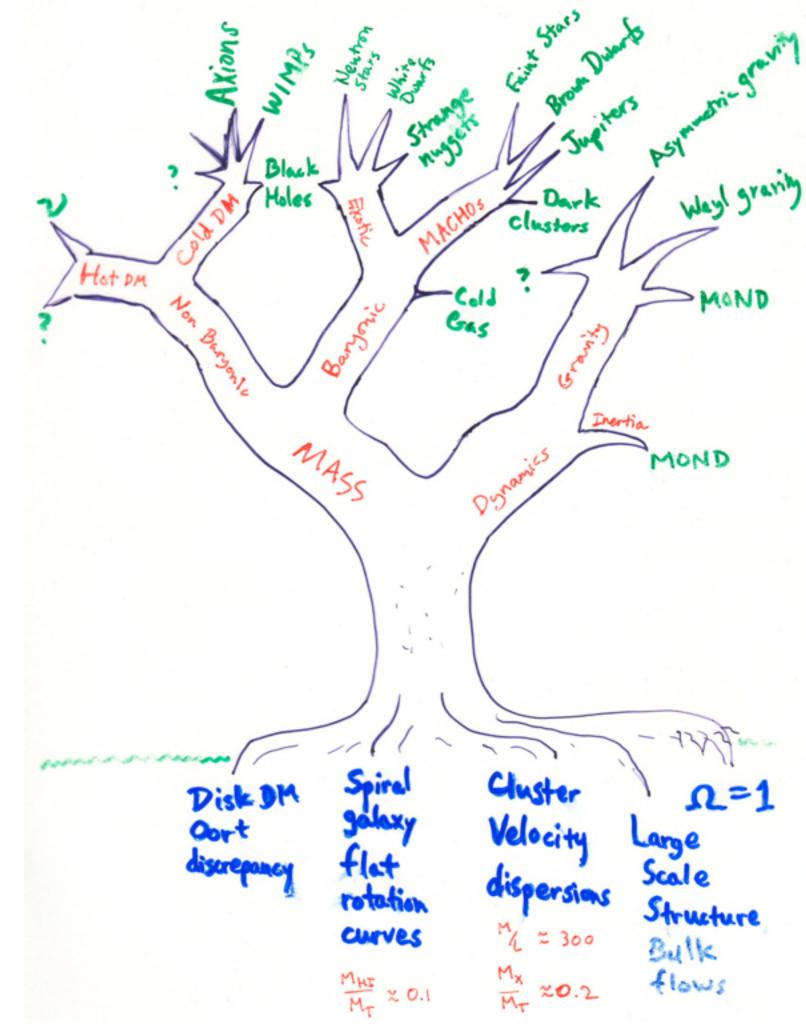
## DARK MATTER ASTR 333/433

DARK MATTER HALO ASSEMBLY & SHAPE BARYONIC EFFECTS

Homework 2 Due Feb. 25 Midterm March 1





NFW shape

NFW halos triaxial. More massive halos less round

perhaps because they are still building up hierarchically ?

Maccio et al (2007) *Concentration, spin and shape of dark haloes* 63

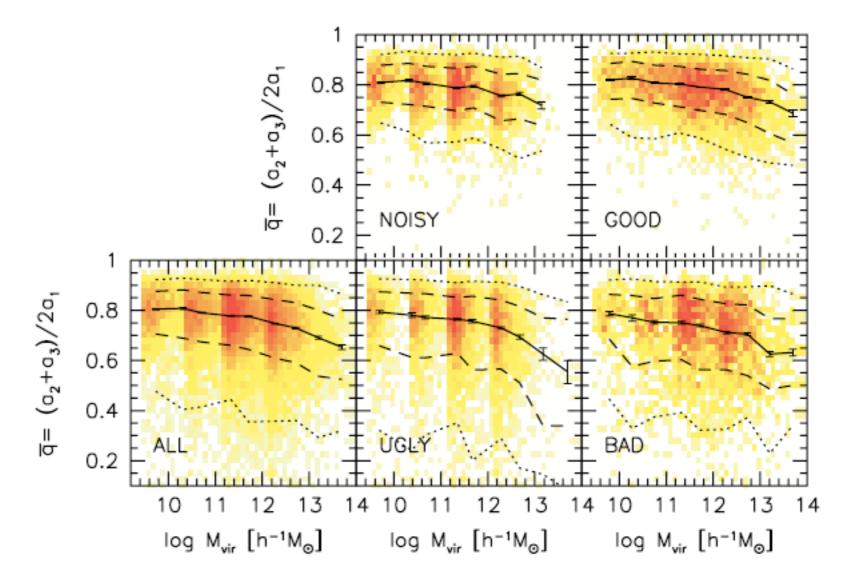
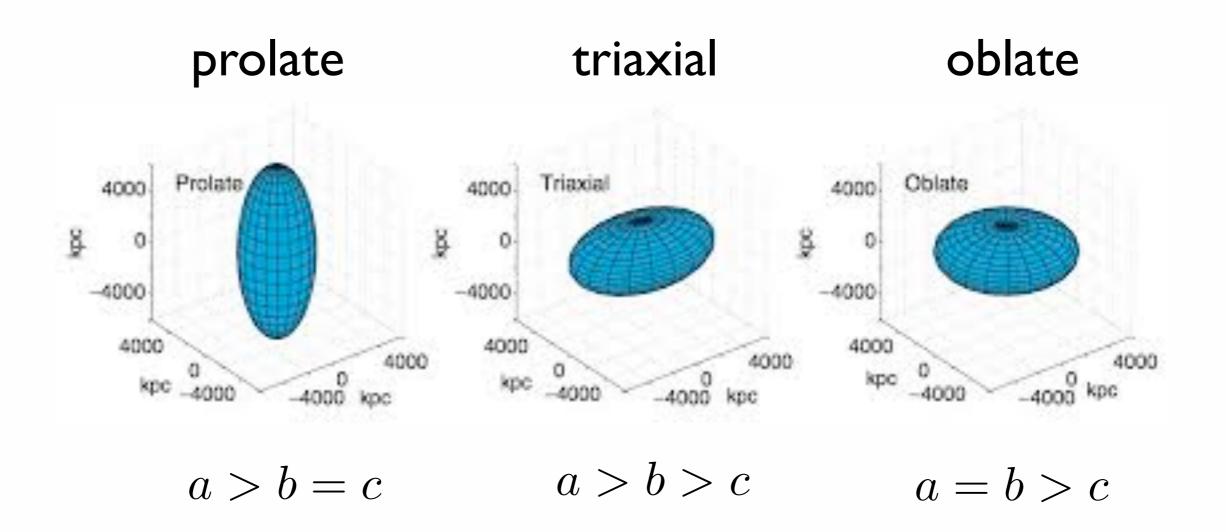


Figure 6. Relation between  $\bar{q}$  and  $M_{vir}$  for different subsamples of haloes. The solid lines show the 50th percentile, dashed lines show the 16th and 84th percentiles, and the dotted lines show the 2.5th and 97.5th percentiles. The error bar gives the Poisson error on the median.

#### shape

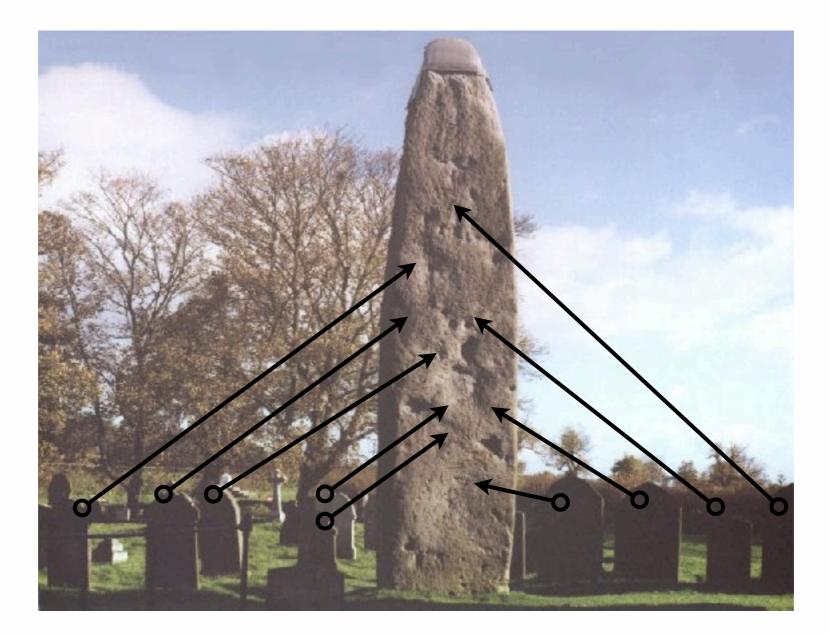


Simulations blobby and even more complicated

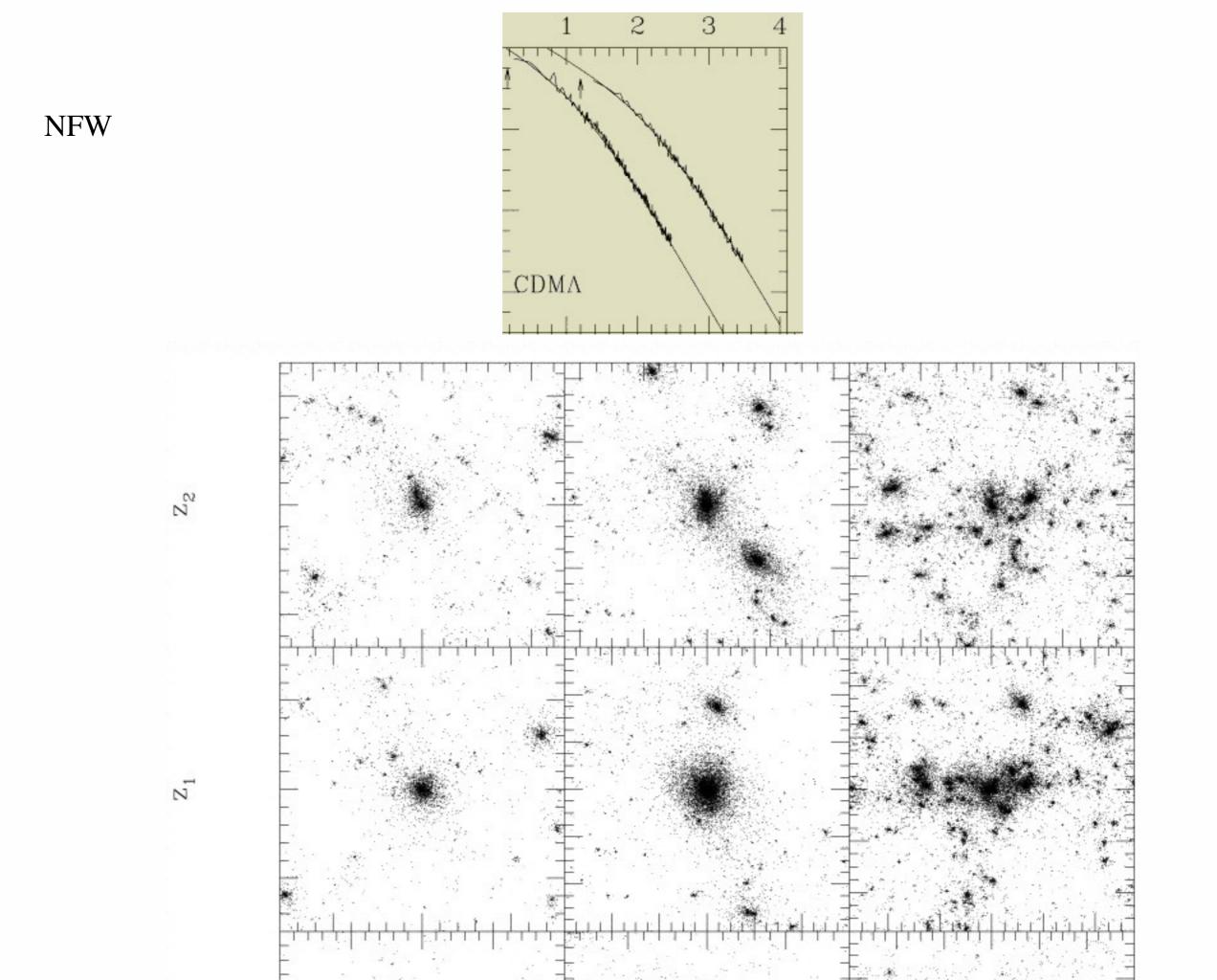
# Monolithic Galaxy Formation



## Hierarchical Galaxy Formation



many little dark matter halos merge to form one big dark matter halo



Hierarchical galaxy formation (*not* monolithic)

Small objects conglomerate to make big ones

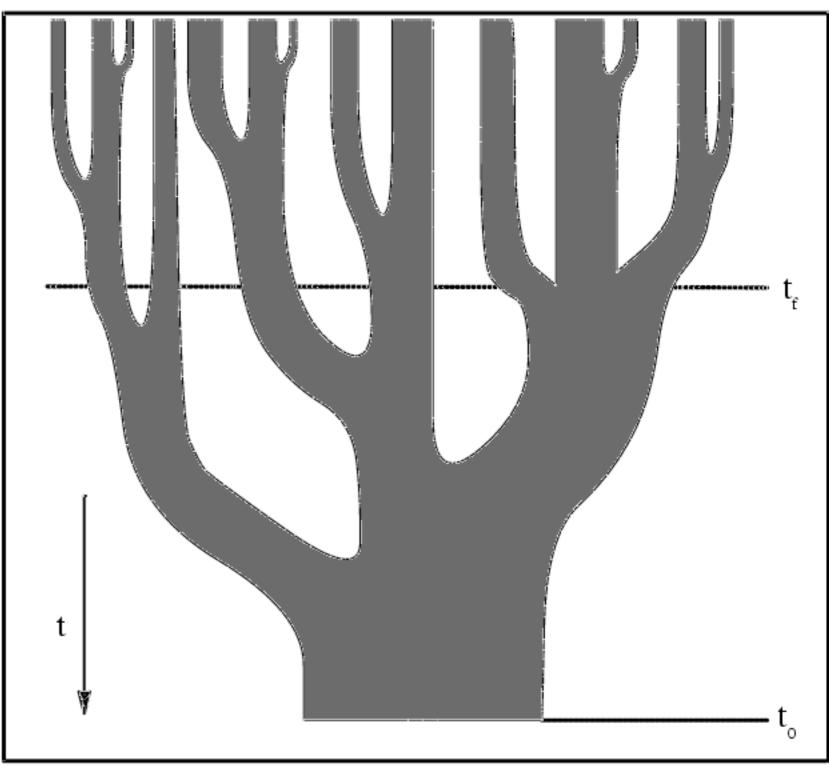


Figure 6. A schematic representation of a "merger tree" depicting the growth of a halo as the result of a series of mergers. Time increases from top to bottom in this figure and the widths of the branches of the tree represent the masses of the individual parent halos. Slicing through the tree horizontally gives the distribution of masses in the parent halos at a given time. The present time  $t_0$  and the formation time  $t_f$  are marked by horizontal lines, where the formation time is defined as the time at which a parent halo containing in excess of half of the mass of the final halo was first created.

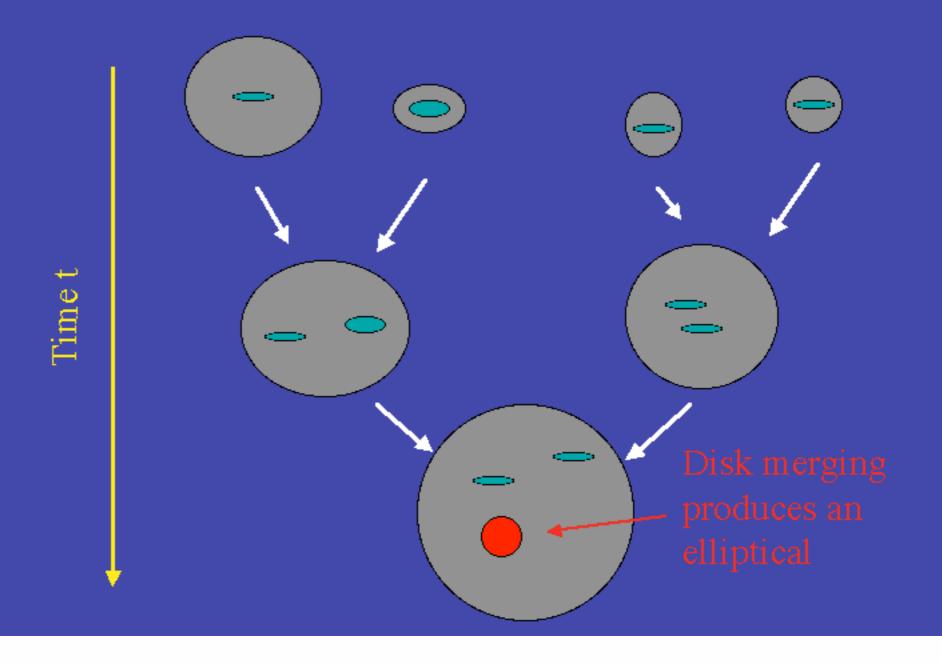
Gray: dark matter halos

Blue: gas rich disks

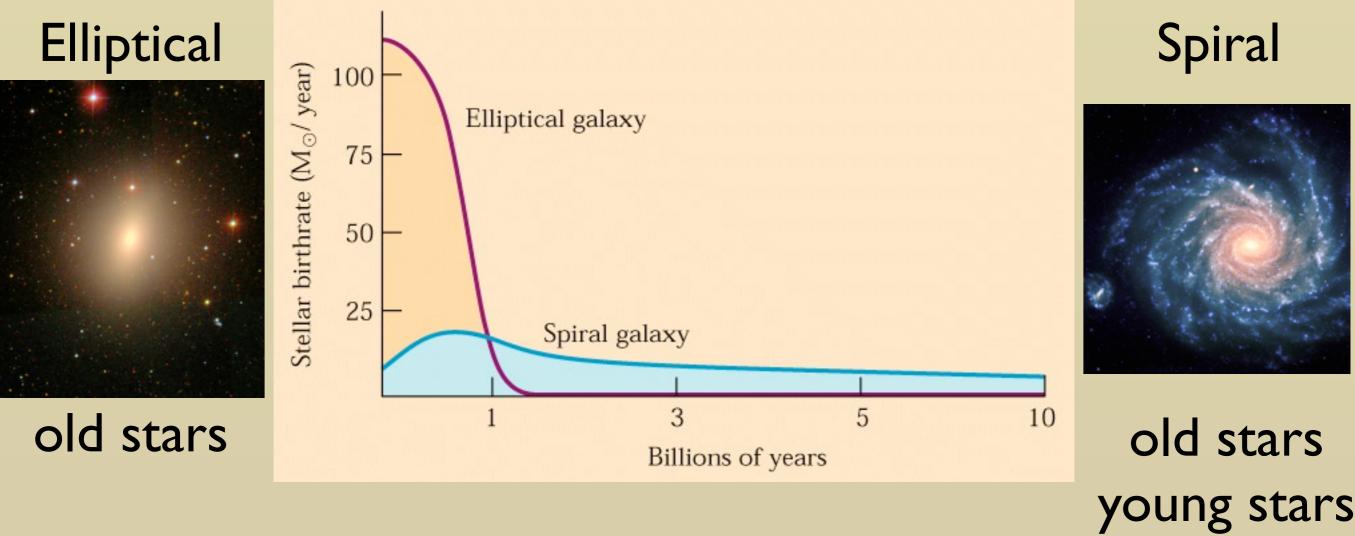
Red: elliptical merger remnant

sometimes it is imagined that a disk re-forms around an elliptical to form a bulge+disk system like and Sa galaxy

#### **Hierarchical Galaxy Formation**



### **Generic Star Formation History**



cold gas