

DARK MATTER

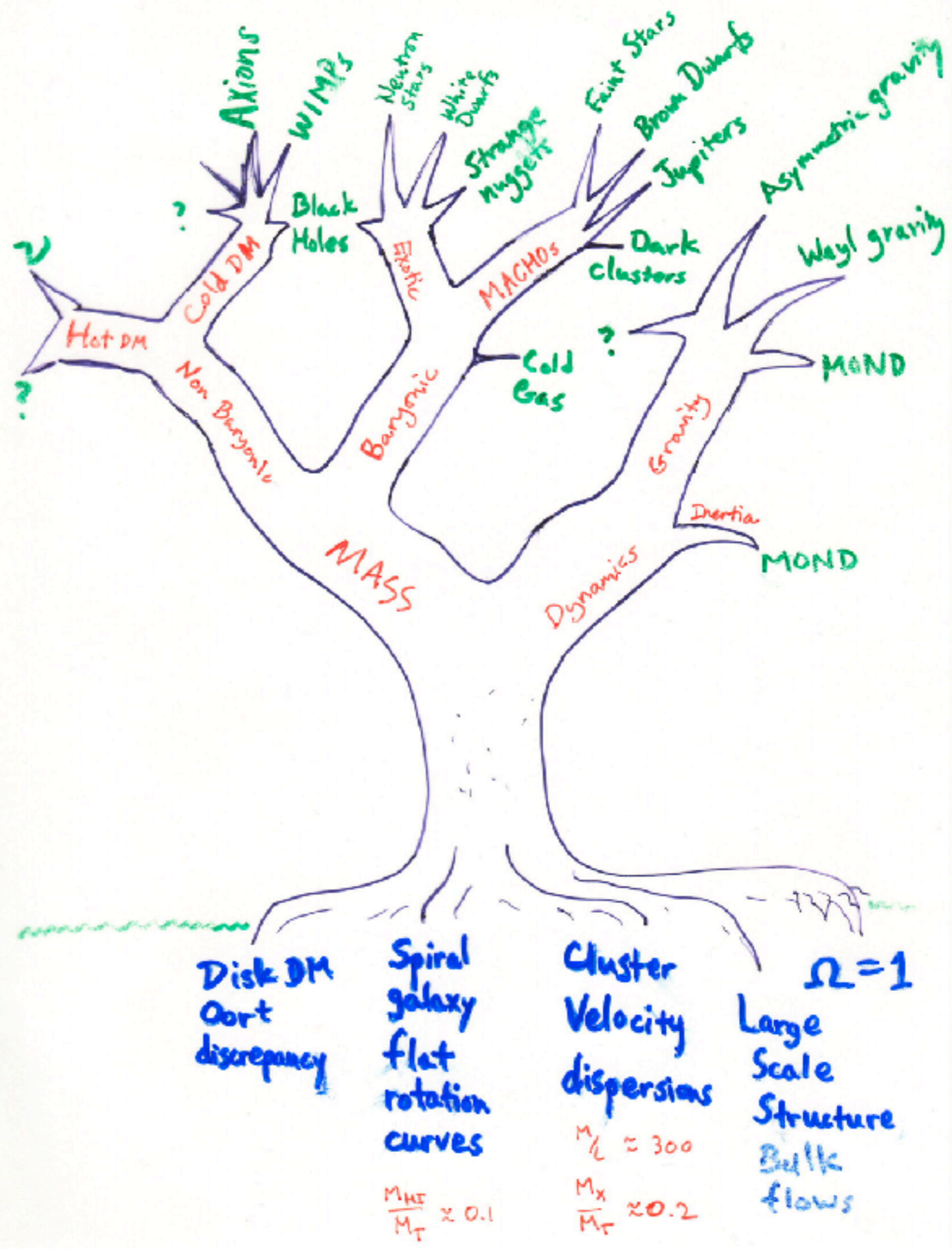
ASTR 333/433

TODAY

FUNDAMENTAL PLANE
ORBITAL ANISOTROPY

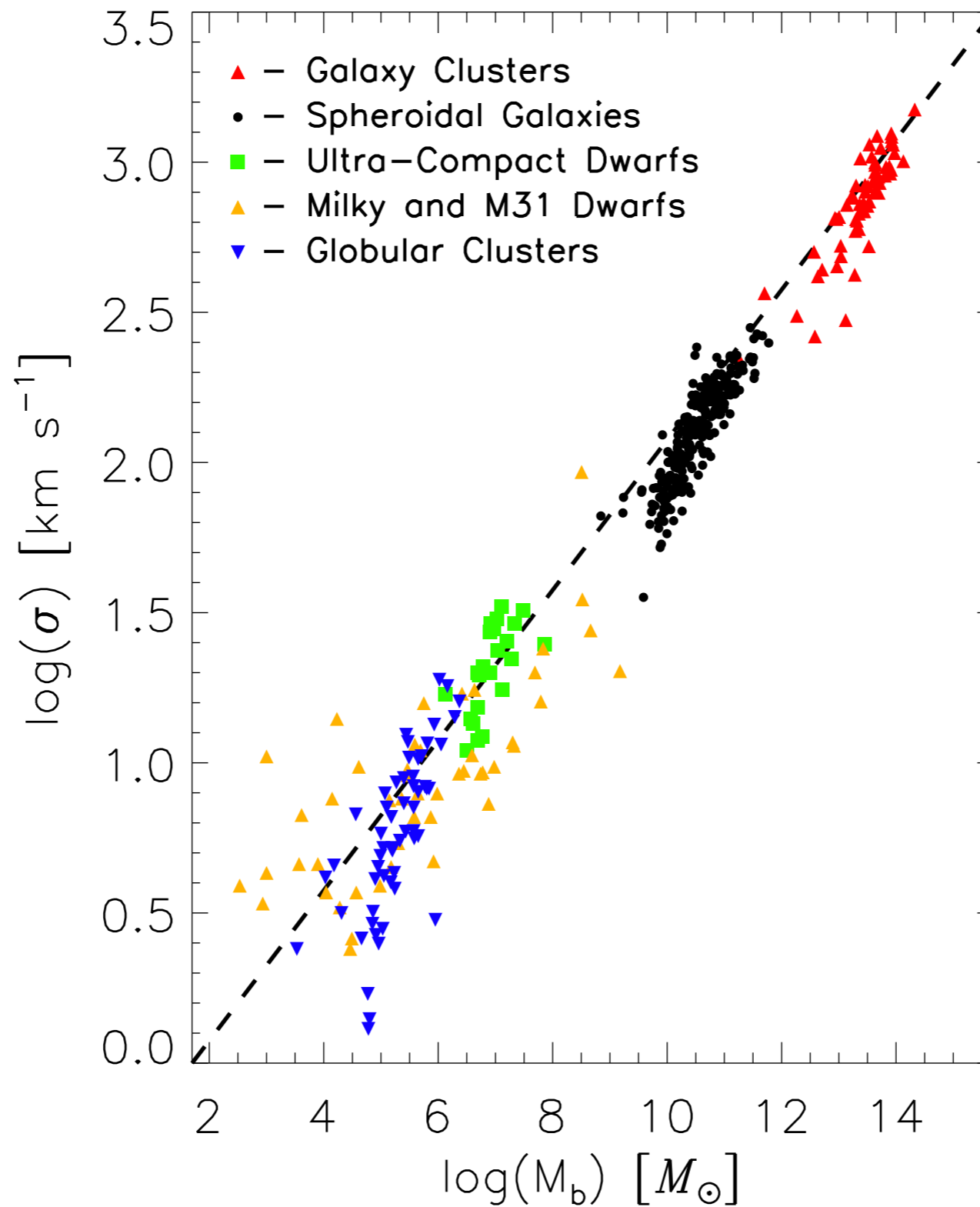
FEEDBACK
BARYON CONTENT

Homework 3 due April 5



Faber-Jackson (pressure supported)

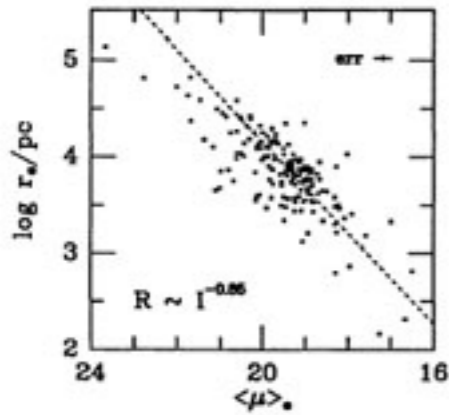
Tully-Fisher for Ellipticals



Fundamental Plane (pressure supported)

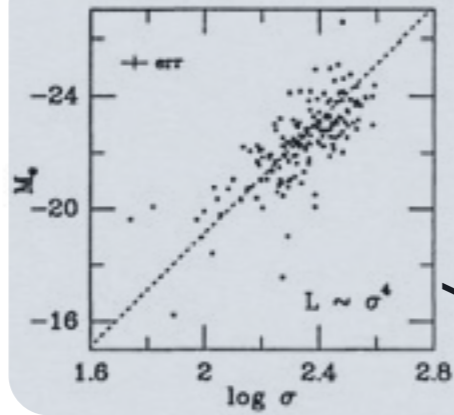
surface
brightness

size



velocity
dispersion

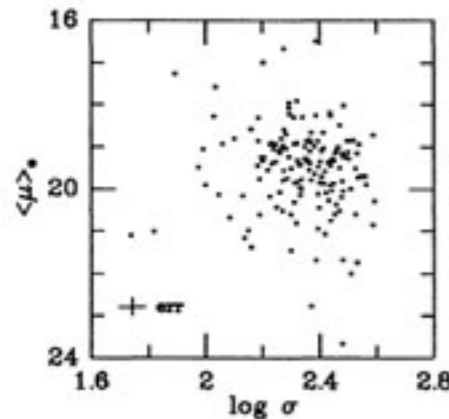
Faber-Jackson



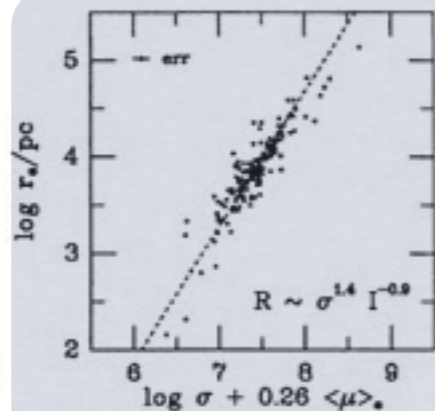
Luminosity

surface
brightness

size



velocity
dispersion



size

v-disp &
surf bright.

Fundamental Plane

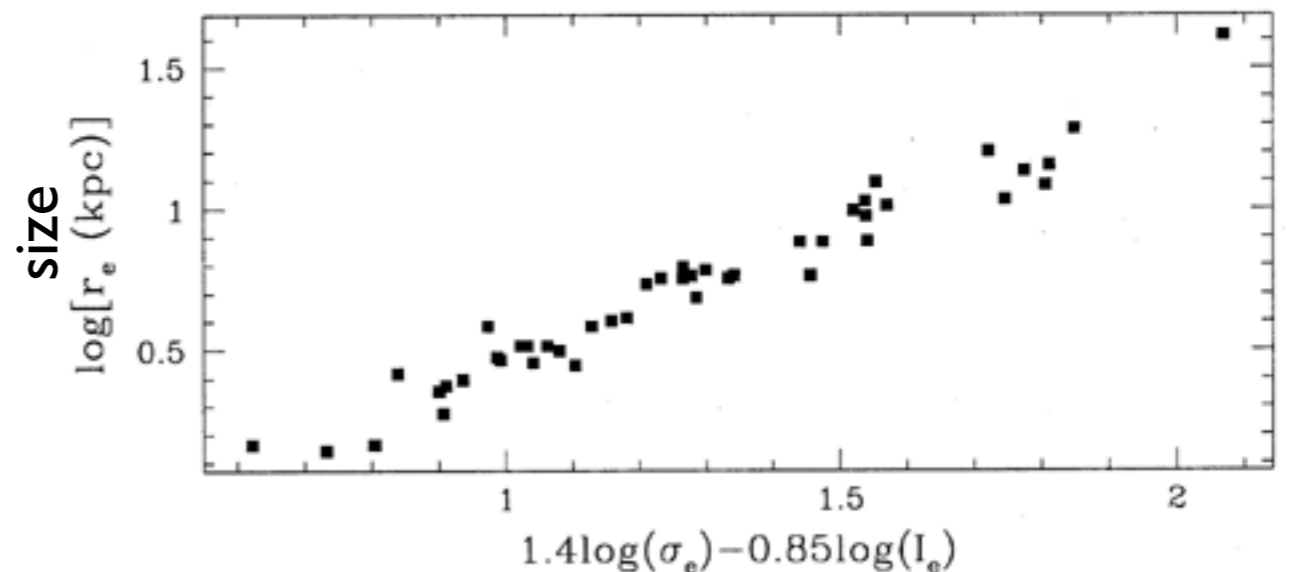
“Viral” fundamental plane

$$R_e \propto \sigma^2 I_e^{-1}$$

observed fundamental plane
“tilted” wrt virial expectation:

$$R_e \propto \sigma^{1.4} I_e^{-0.85}$$

velocity dispersion & surface brightness



$$M \propto \sigma^2 R_e$$

virial theorem

$$L \propto I_e R_e^2$$

luminosity,
surface brightness
size

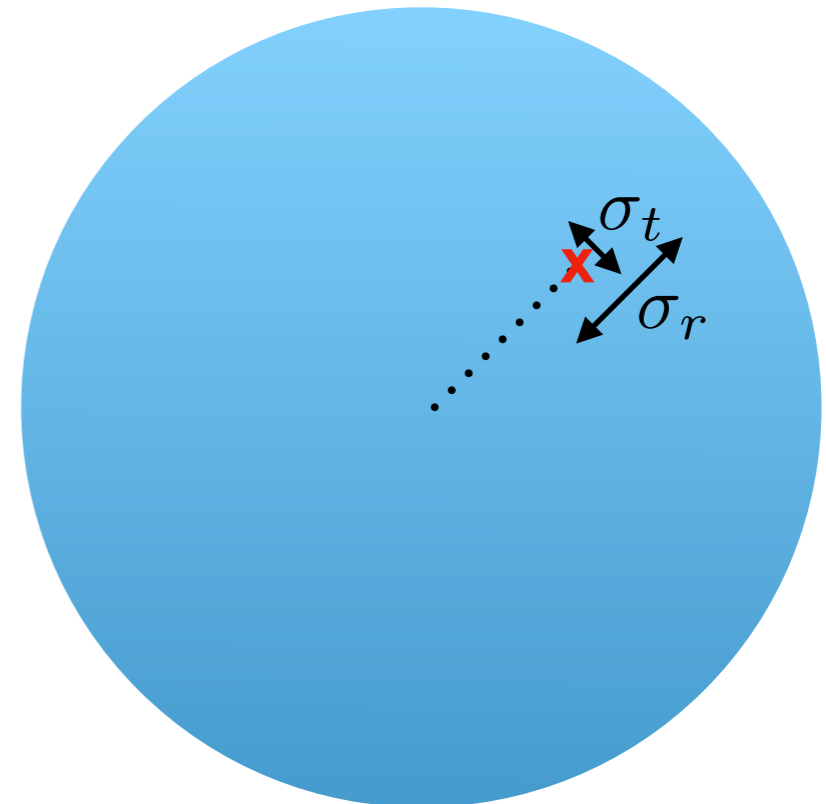
Orbital Anisotropy

The anisotropy parameter measures how radial or circular orbits are

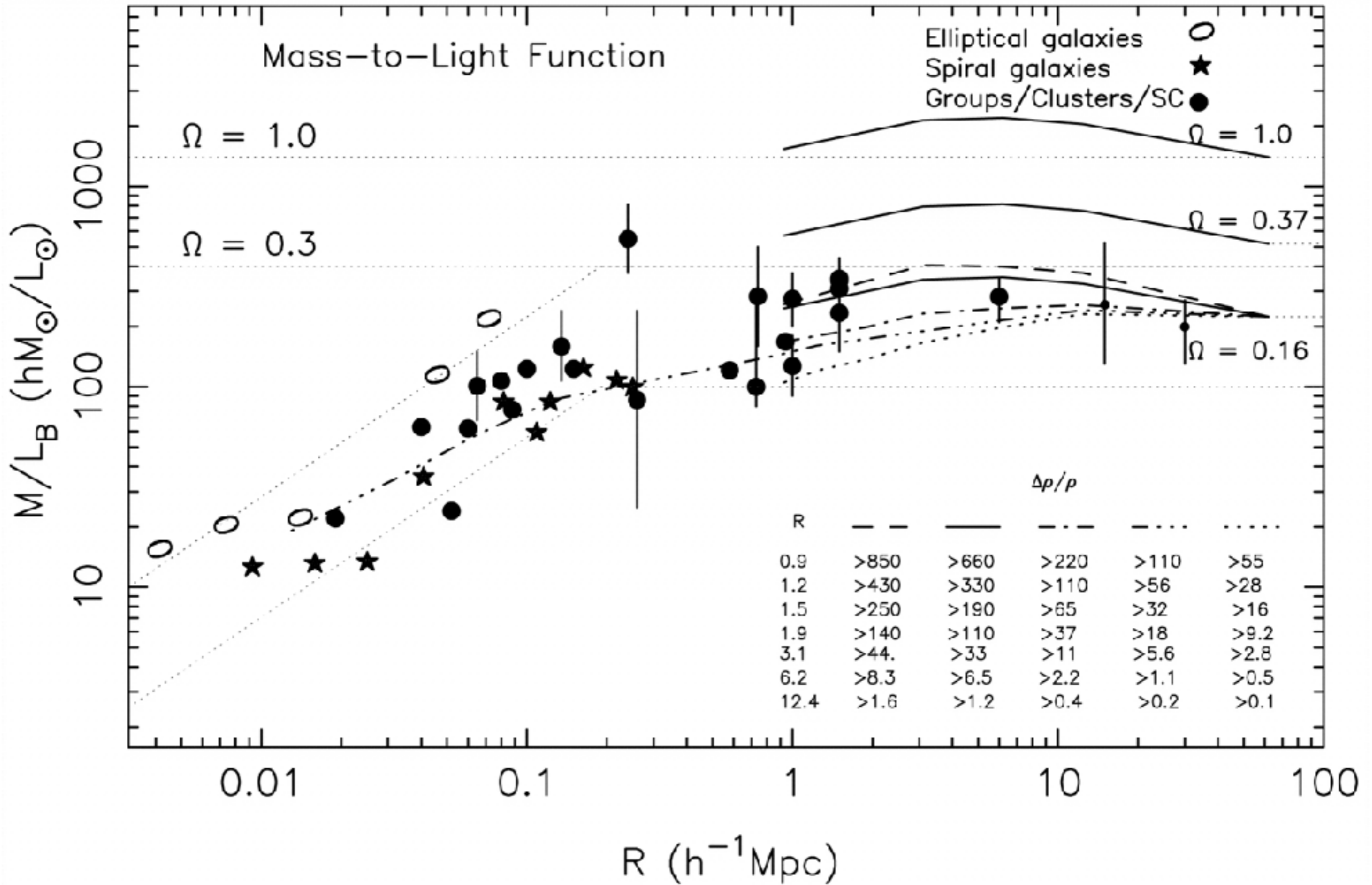
$$\beta = 1 - \frac{\sigma_t^2}{\sigma_r^2}$$

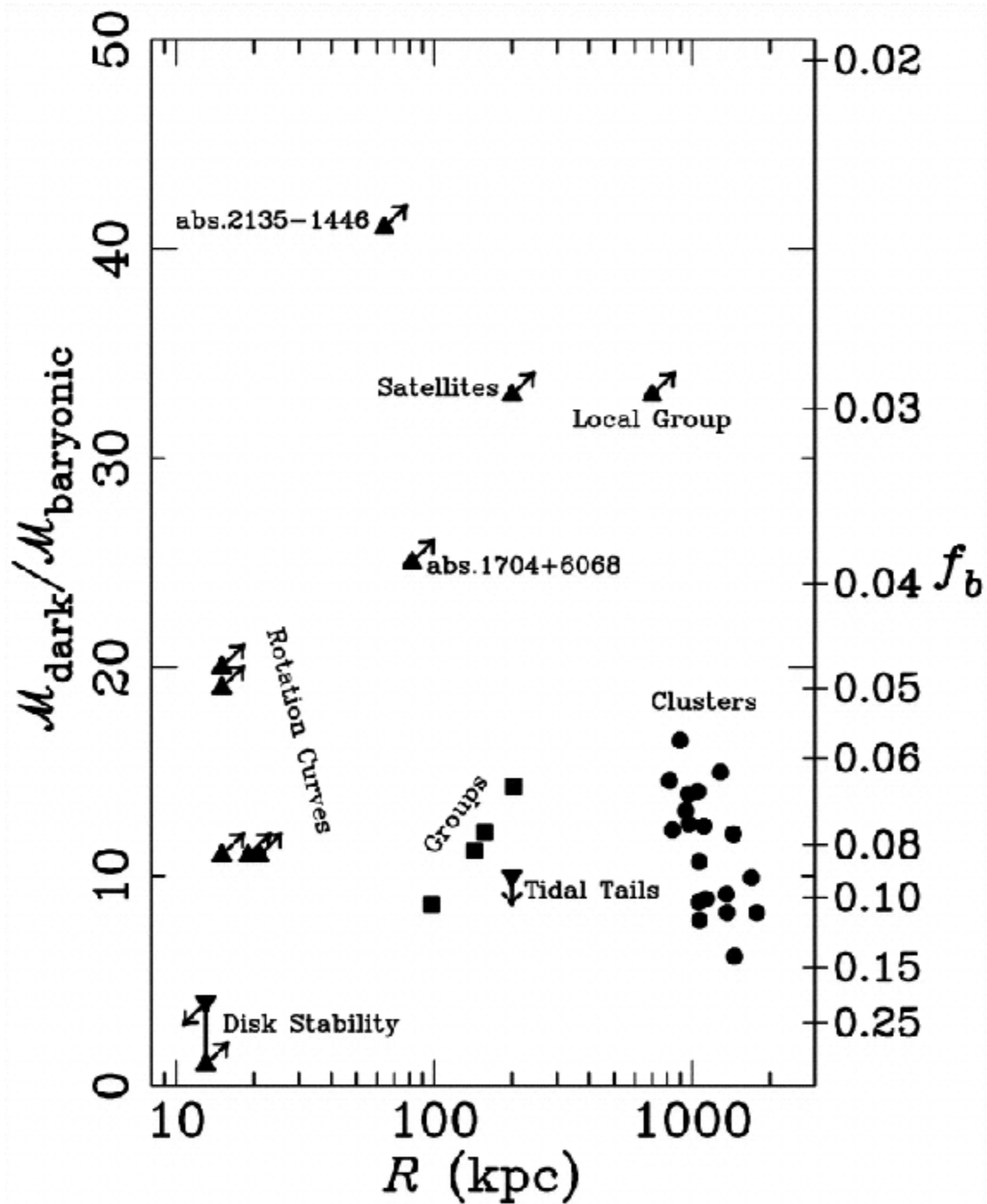
← tangential velocity dispersion
← radial velocity dispersion

$$M(r) = G^{-1} r \sigma_r^2 \left(-\frac{d \ln n_*}{d \ln r} - \frac{d \ln \sigma_r^2}{d \ln r} - 2\beta \right)$$

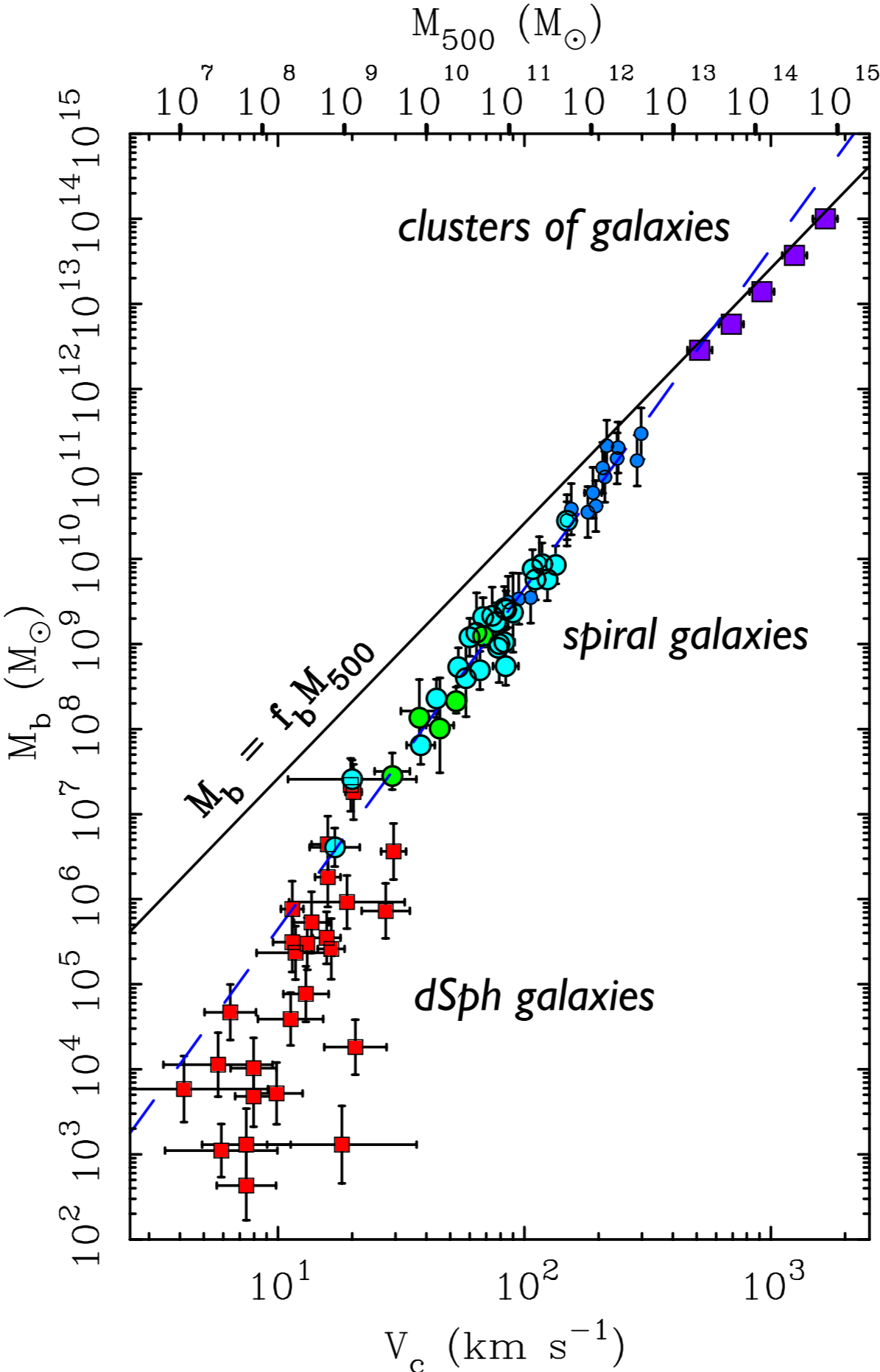


Neta Bahcall



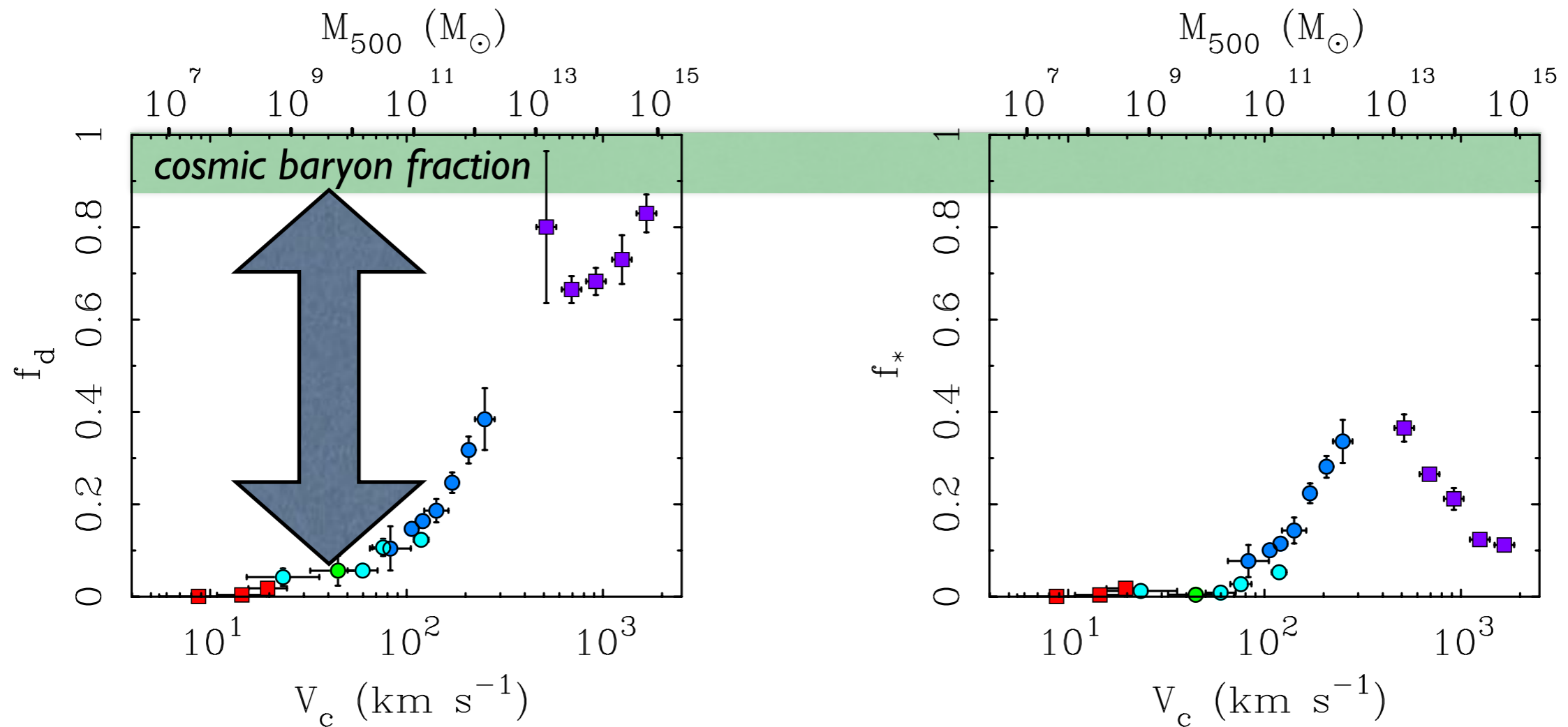


Extended TF



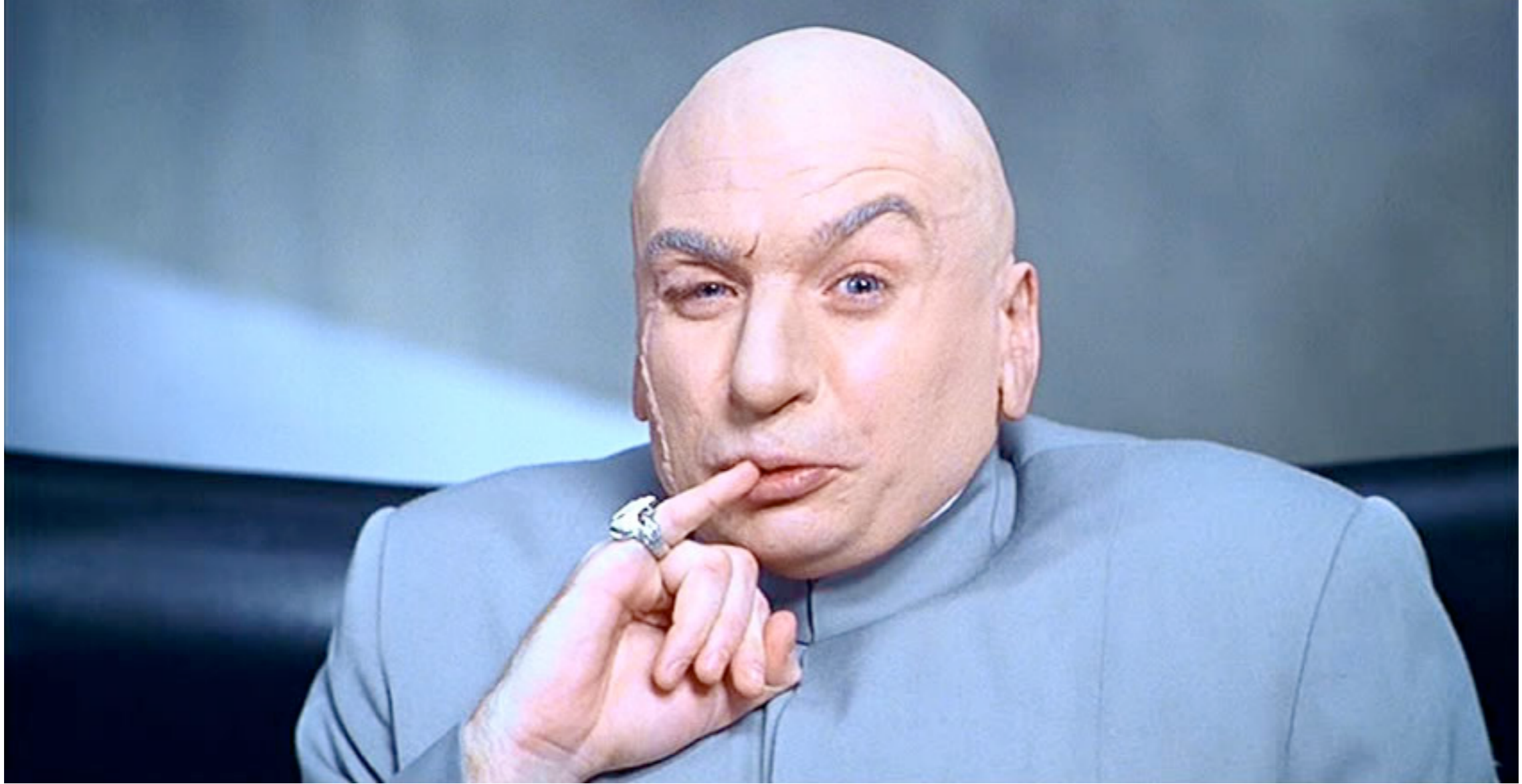
$$f_d = \frac{M_b}{f_b M_{500}}$$

$$f_* = \frac{M_*}{f_b M_{500}}$$



Halo by halo missing baryon problem

2 missing mass problems: baryonic AND non-baryonic DM



Theorist: I have an even better idea. I'm going to place model galaxies in easily escapable dark matter halos by invoking overly elaborate and exotic feedback schemes.

Observer: Wait, aren't you even going to test that? It might not work!

Theorist: No no no, I'm going to leave it alone and not actually witness it happening; I'm just gonna assume it all went to plan. What?