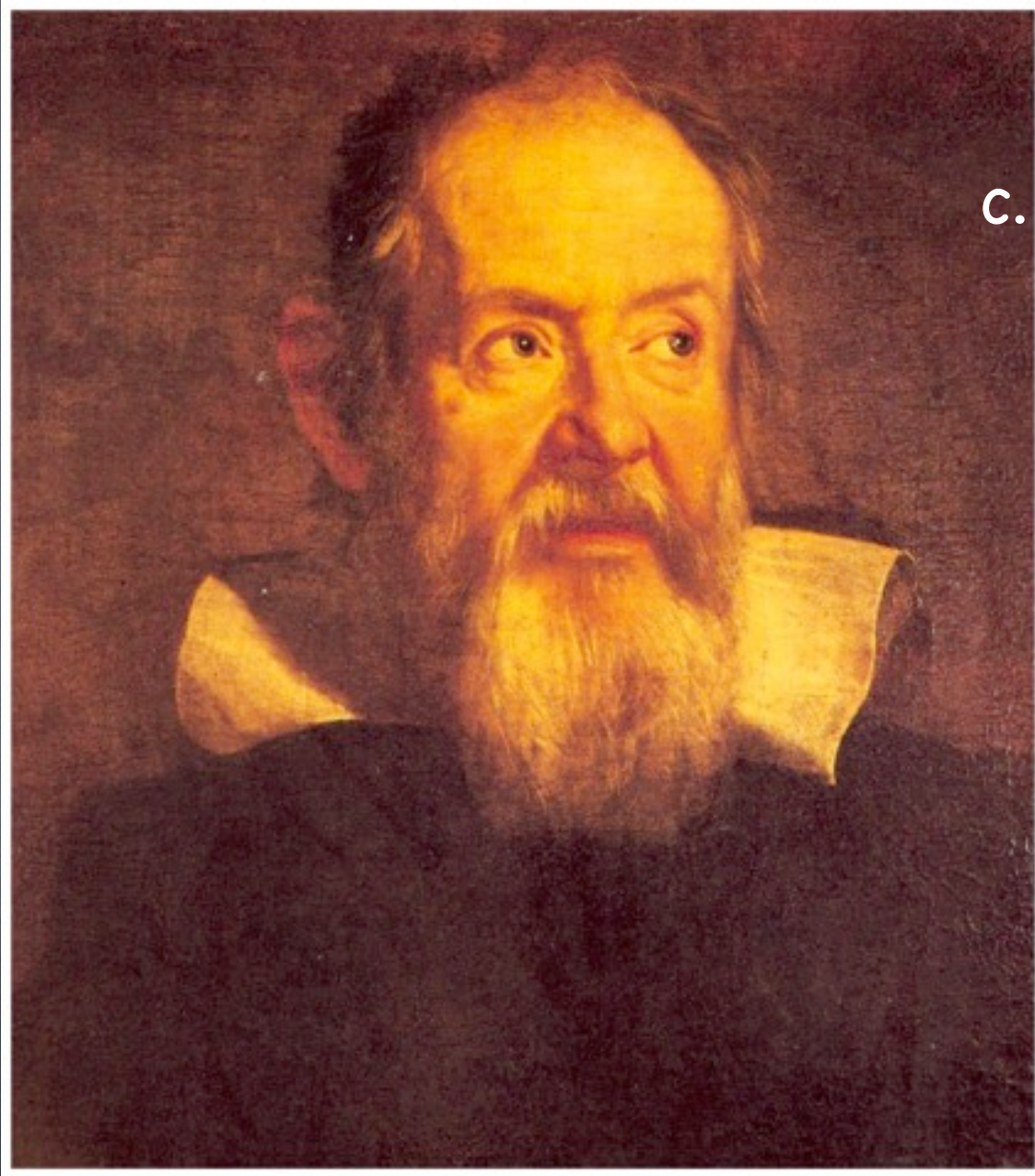


TODAY

- THE COPERNICAN REVOLUTION
- GALILEO
- PLANETARY MOTION
 - TYCHO BRAHE'S OBSERVATIONS
 - KEPLER'S LAWS

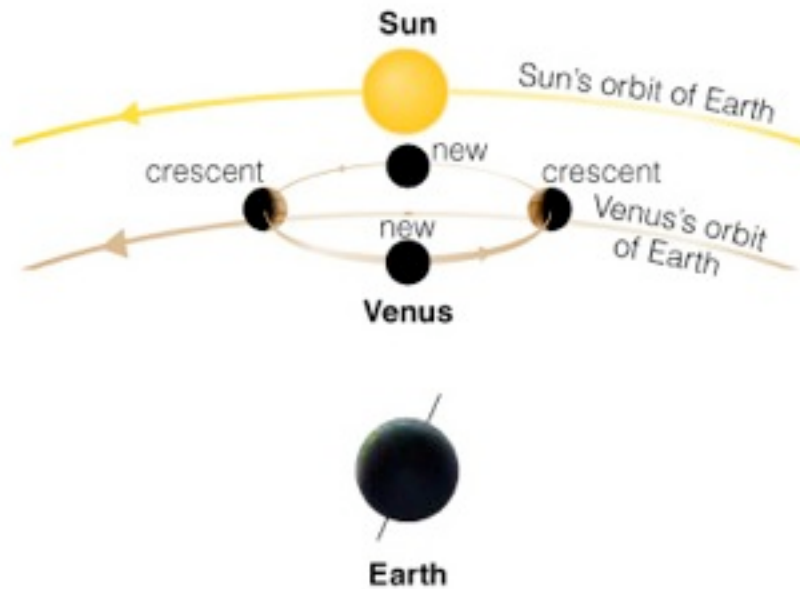
Galileo

c. 1564-1640

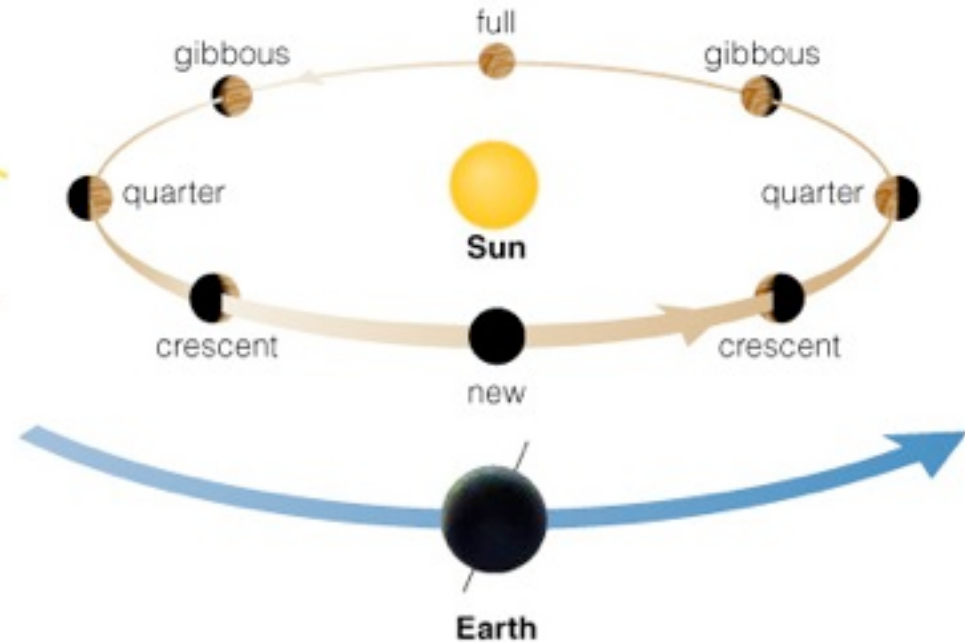


First telescopic astronomical observations

Ptolemaic View of Venus



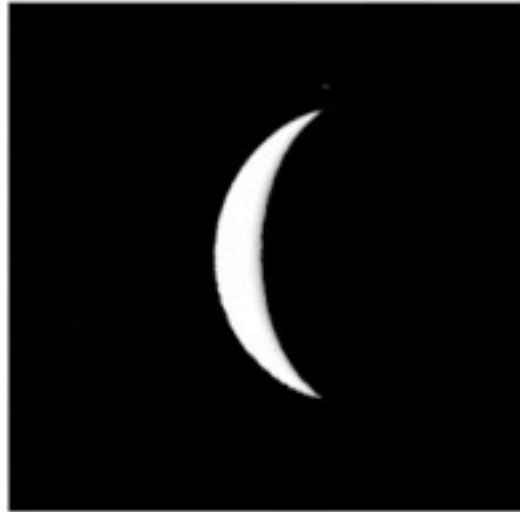
Copernican View of Venus



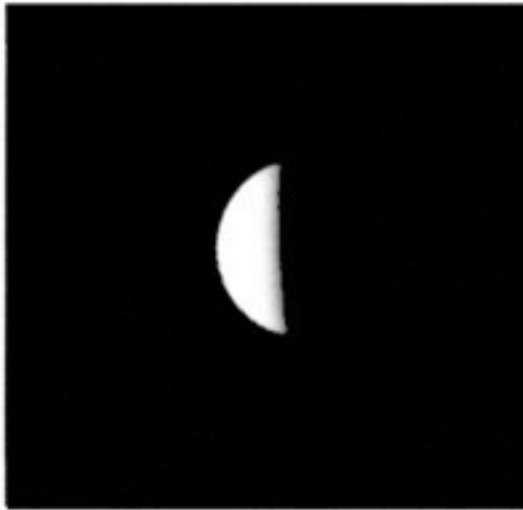
Galileo's observations of phases of Venus proved that it orbits the Sun and not Earth.



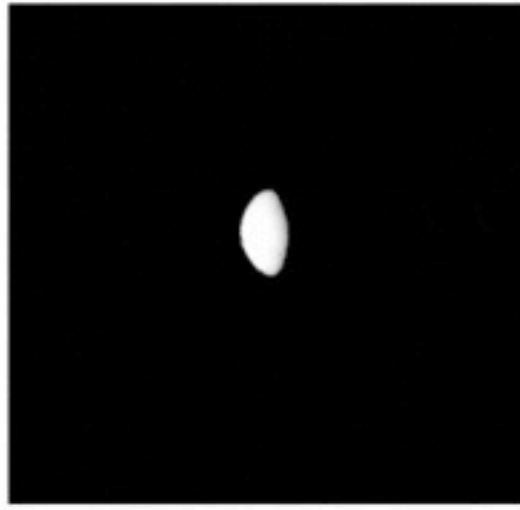
$\alpha = 58^\circ$



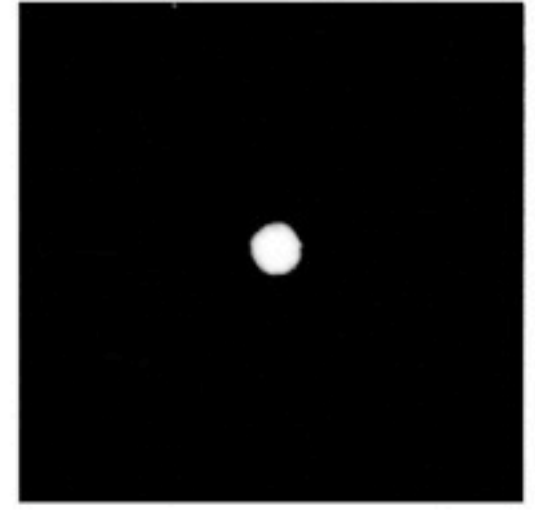
$\alpha = 42^\circ$



$\alpha = 24^\circ$



$\alpha = 15^\circ$



$\alpha = 10^\circ$

Phase and angular size of Venus depend on elongation

Competing Cosmologies

Geocentric

Ptolemaic

Earth at center

Heliocentric

Copernican

Sun at center

The sun is the source of light in both models

Explains

- **Motion of Sun**
- **Motion of Moon**
- **Solar and Lunar Eclipses**
- **Phases of Moon**

Explains

- **Motion of Sun**
- **Motion of Moon**
- **Solar and Lunar Eclipses**
- **Phases of Moon**

Retrograde Motion

Needs epicycles

Consequence of Lapping **nicer**

Inferiority of Mercury & Venus

Must tie to sun

Interior to Earth's Orbit **nicer**

Predicts

- No parallax ✓
- Venus: crescent phase only X

- Parallax X
- Venus: all phases ✓

Heliocentric Cosmology

- Provides better explanation for
 - Retrograde motion
 - proximity of Mercury and Venus to the Sun
- Provides only explanation for
 - Phases of Venus
 - Angular size variation of Venus
- What about parallax?
 - Undetectable if stars VERY distant
 - Finally detected in 1839

Galileo's telescopic discoveries

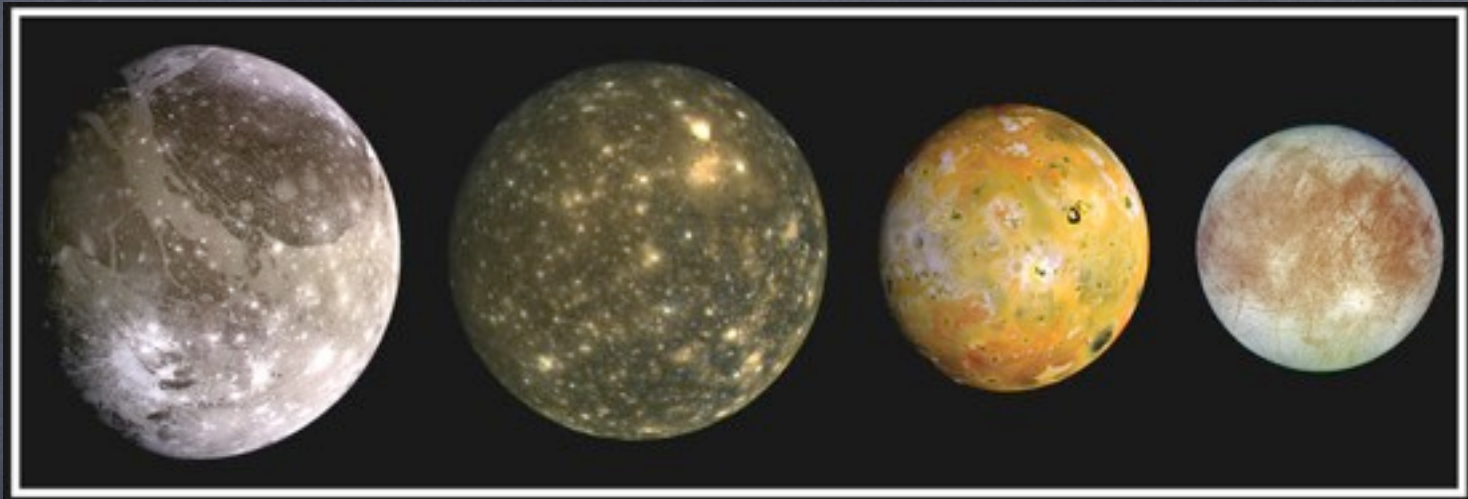
- Stars in the Milky Way
- Mountains on the Moon
- Sun spots (celestial spheres NOT perfect)
- Rings of Saturn (barely resolved)
- Moons of Jupiter ("Medicean stars")
 - Earth NOT center of all revolution
- Phases of Venus
 - Good test of geocentric hypothesis

Jupiter and moons



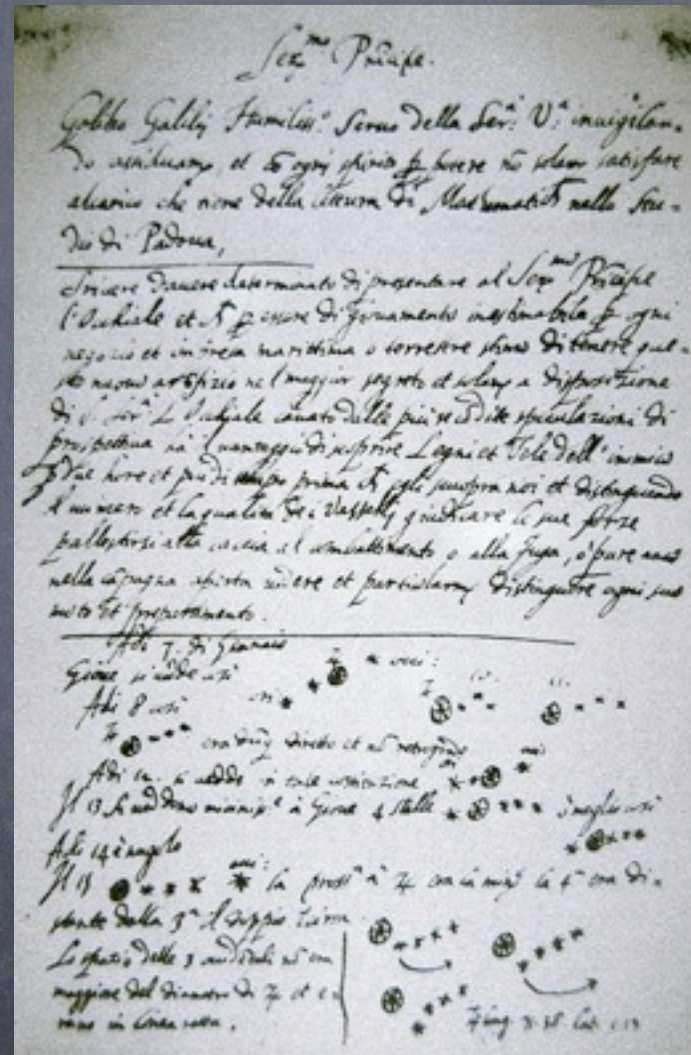
Fig. 4.16

Galilean moons (from Galileo spacecraft!)



NASA

Letter from Galileo
to Prince of Venice
reporting the discovery
of Jupiter's moons...



Обновление Десятка
1620

2. Д. 9bris. марта H. 12	○ * *
30. марч	* * ○ *
2. Xbris.	○ * * *
3. марч	○ * *
3. Ho. 5.	* ○ *
4. марч.	* ○ * *
6. марч	* * ○ *
8. марч H. 13.	* * * ○
10. марч.	* * * ○ *
11.	* * ○ *
12. H. 4 уек:	* ○ *
13. марч	* * ○ *
14. курс.	* * * ○ *

"Medician stars"

Fig. 4.17

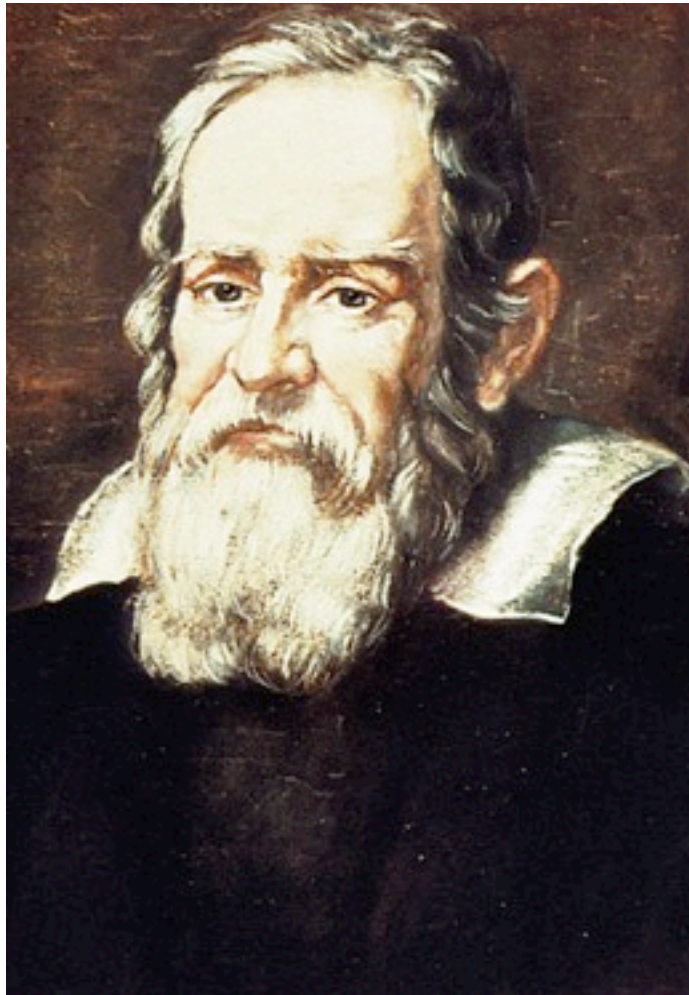
Heavenly
spheres
NOT
perfect



The moon has
geographical
features like
mountains & craters

Even the
sun has
spots!





Galileo Galilei

In 1633 the Catholic Church ordered Galileo to recant his claim that Earth orbits the Sun.

His book on the subject was removed from the Church's index of banned books in 1824.



Galileo was formally vindicated by the Church in 1992.

Tycho Brahe

the last great
naked-eye
observer



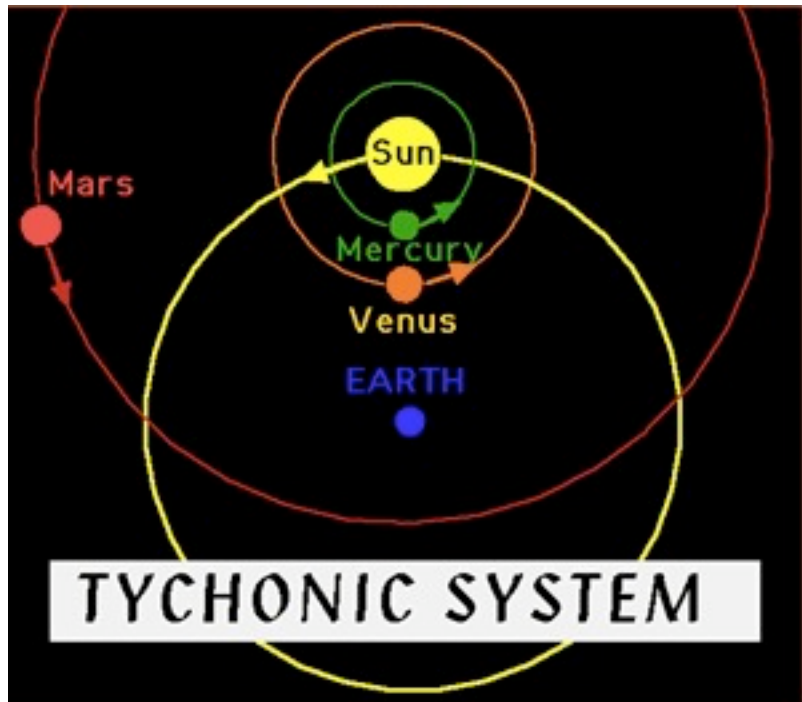
1546–1601

Tycho Brahe



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- Brahe compiled the most accurate (one arcminute) naked eye measurements ever made of planetary positions.
- He still could not detect stellar parallax, and thus still thought Earth must be at the center of the solar system (but recognized that other planets go around Sun).





Johannes Kepler
(1571–1630)

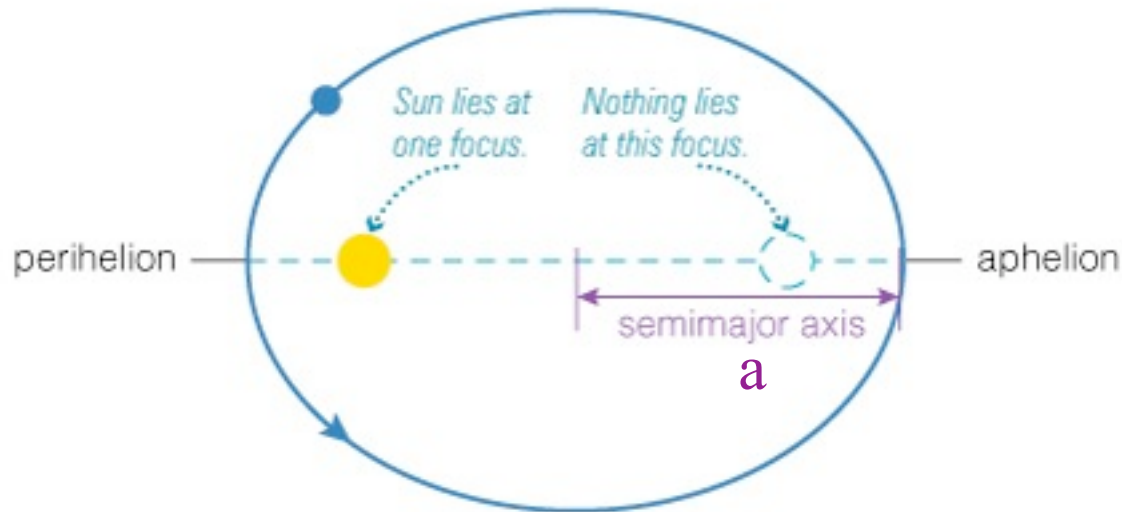


- Kepler analyzed Brahe's data
- Kepler first tried to match Tycho's observations with circular orbits.
- But an 8-arcminute discrepancy led him eventually to ellipses.

“If I had believed that we could ignore these eight minutes [of arc], I would have patched up my hypothesis accordingly. But, since it was not permissible to ignore, those eight minutes pointed the road to a complete reformation in astronomy.”

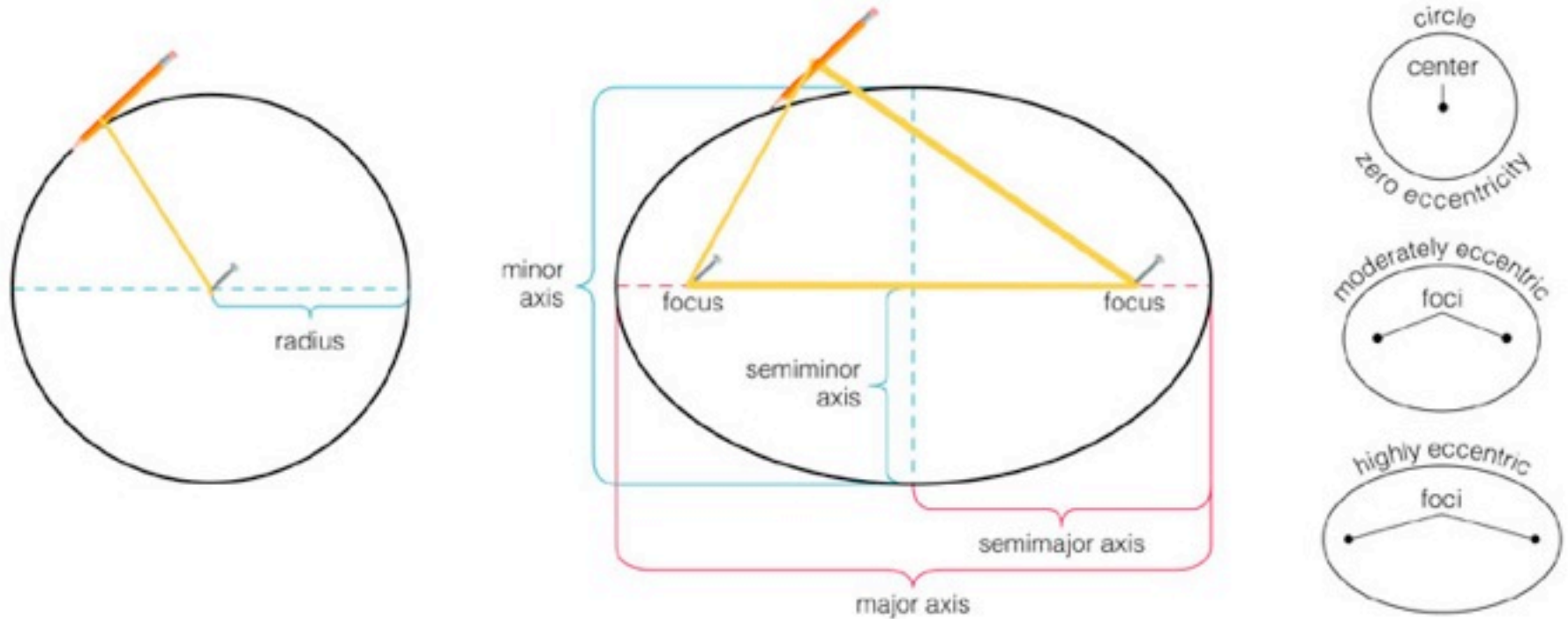
Kepler's Laws of planetary motion

Kepler's First Law: The orbit of each planet around the Sun is an *ellipse* with the Sun at one focus.



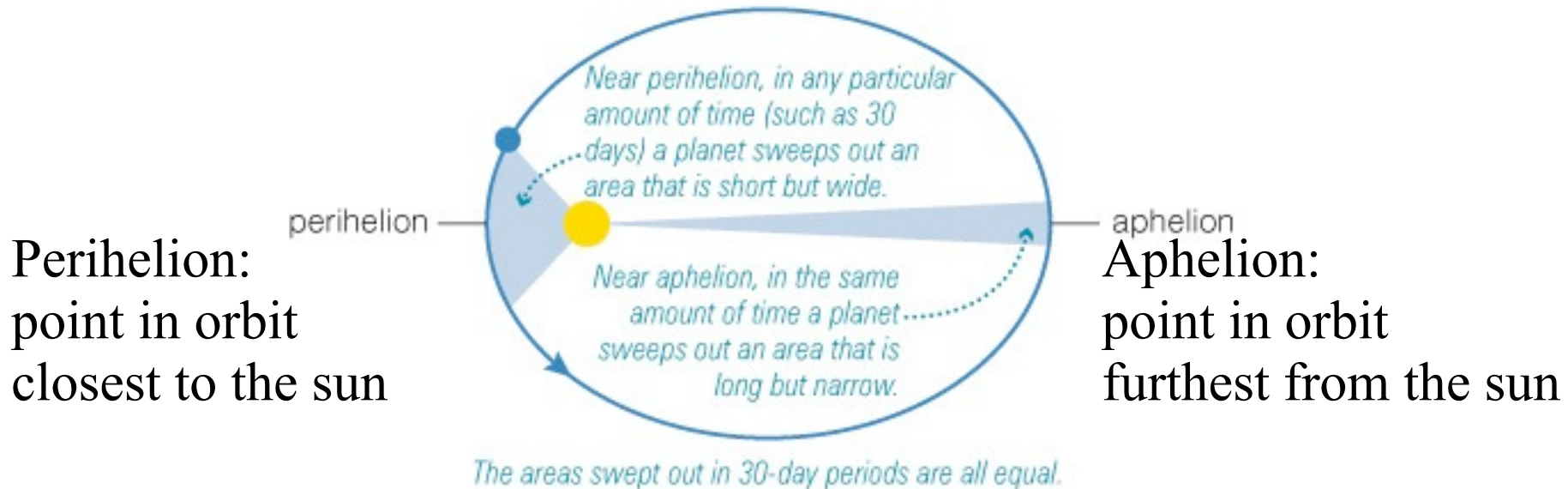
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An ellipse is the shape that is equidistant from two **foci**. The **eccentricity** of an ellipse depends on the ratio of the long and short axes. Half of the long axis is the **semimajor axis, a** .



An ellipse looks like an elongated circle. Indeed, a circle is a special case of an ellipse where the two foci overlap.

Kepler's Second Law: As a planet moves around its orbit, it sweeps out equal areas in equal times.



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This means that a planet travels faster when it is nearer to the Sun and slower when it is farther from the Sun.

Kepler's Third Law

More distant planets orbit the Sun at slower average speeds, obeying the relationship

$$P^2 = a^3$$

p = orbital period in years

a = distance from Sun in AU

(semi-major axis of orbit's ellipse)

Earth: $P = 1$ year, $a = 1$ AU

Kepler's Third Law

- A worked example: Mercury: $P = 0.24$ year

$$P^2 = a^3$$

$$a = P^{2/3}$$

$$a = (0.24)^{2/3}$$

$$a = 0.39 \text{ AU}$$

Kepler's Third Law

- A worked example: Jupiter: $a = 5.2$ AU

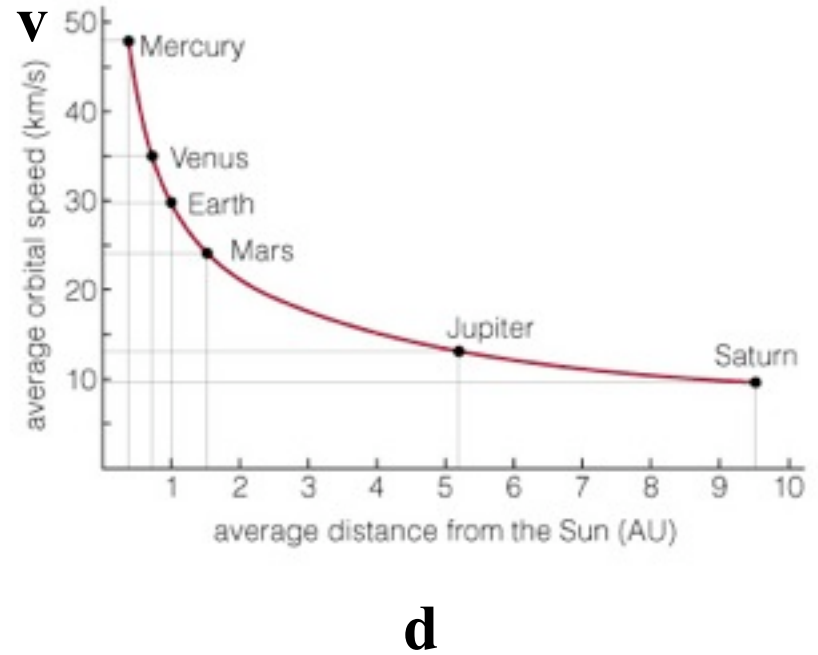
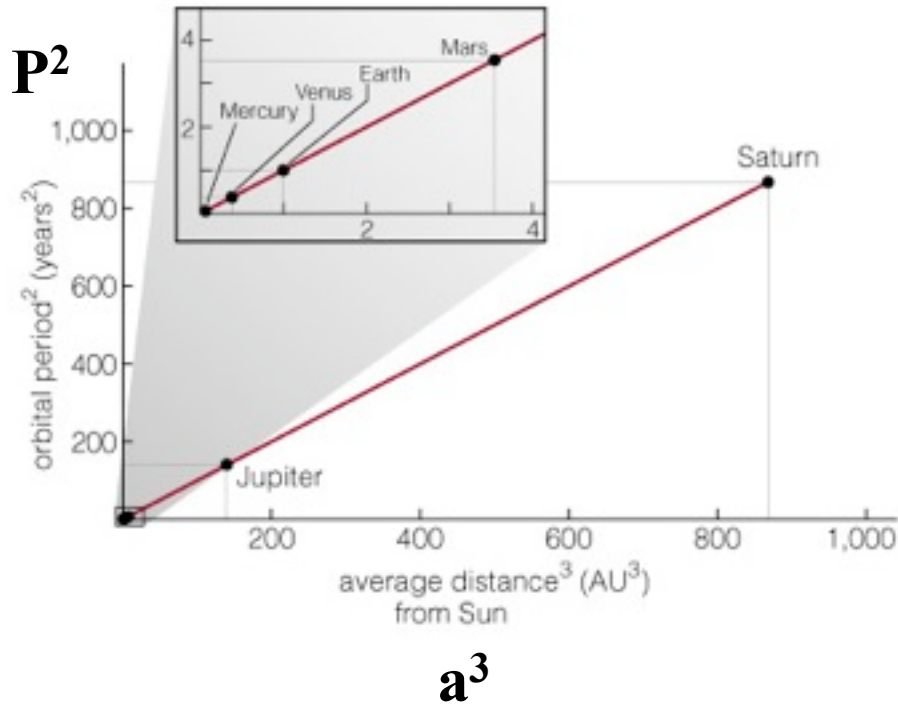
$$P^2 = a^3$$

$$P = a^{3/2}$$

$$P = \sqrt{(5.2)^3}$$

$$P = 11.9 \text{ years}$$

Graphical version of Kepler's Third Law



- Kepler's Laws:

1. The orbit of each planet is an ellipse with the Sun at one focus.
2. As a planet moves around its orbit it sweeps out equal areas in equal times.
3. More distant planets orbit the Sun at slower average speeds: $p^2 = a^3$.

Major Figures in Copernican Revolution

- Copernicus: [Re]Invented Heliocentric model.
- Tycho Brahe: Collected large, accurate database on planetary positions & motions. Failed to detect parallax; suggested hybrid.
- Kepler: Used Brahe's data to discern his three Laws of planetary motion.
- Galileo: telescopic observations and physical arguments in favor of heliocentric model.