DARK MATTER

ASTR 333/433
SPRING 2024
TR 11:30AM-12:45PM
SEARS 552

http://astroweb.case.edu/ssm/ASTR333/

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CASE WESTERN RESERVE
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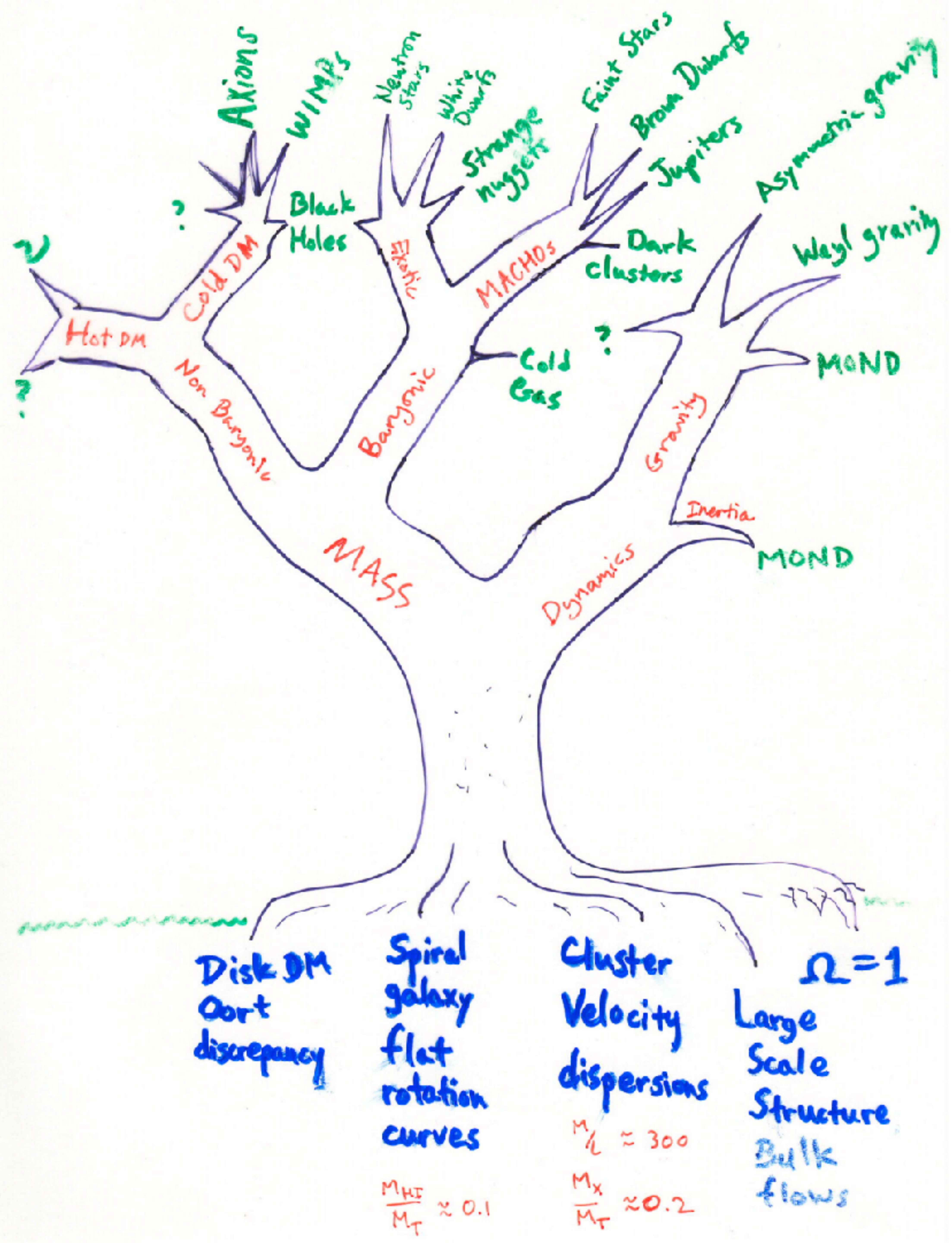
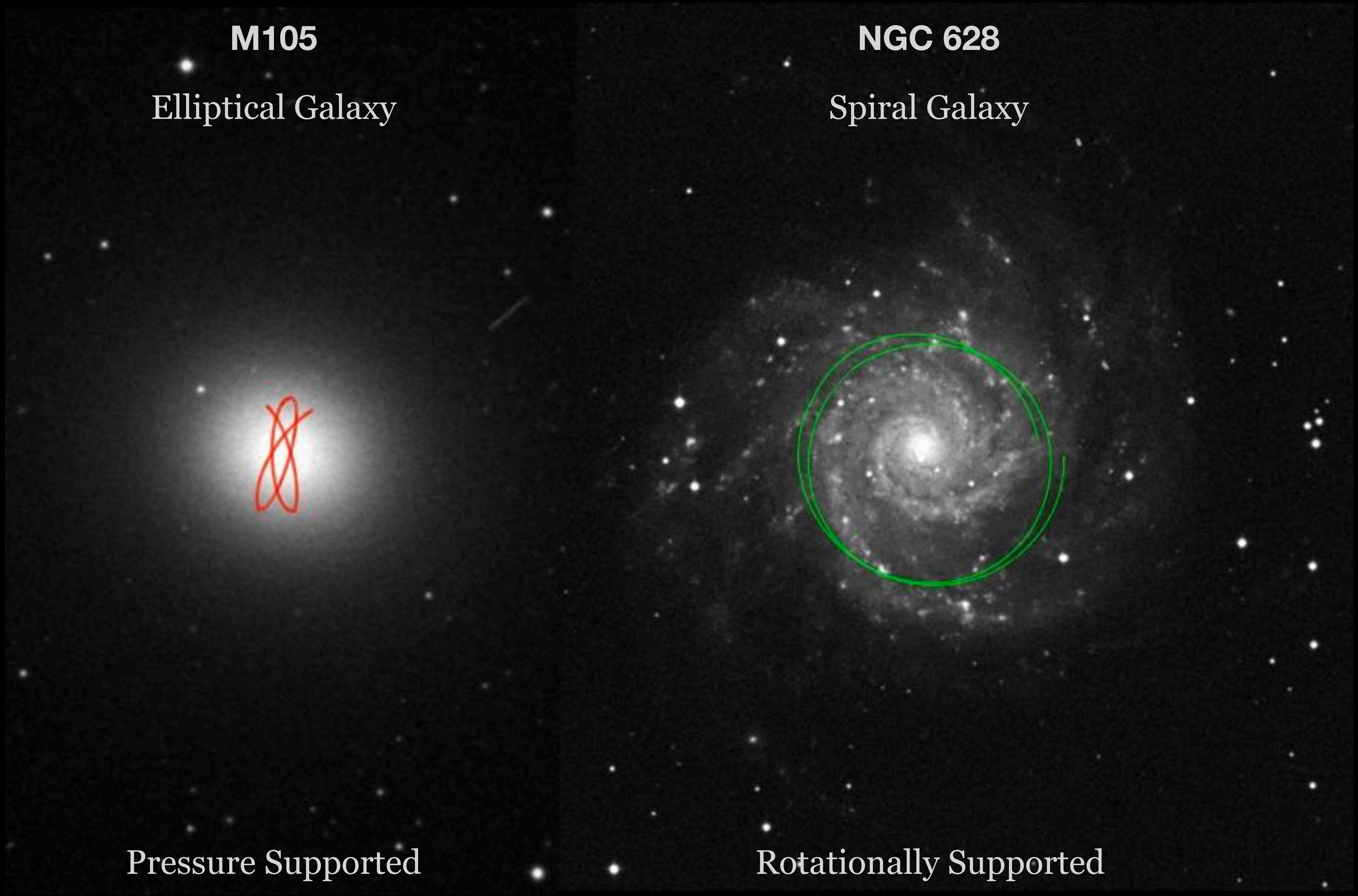


Stellar orbits in galaxies



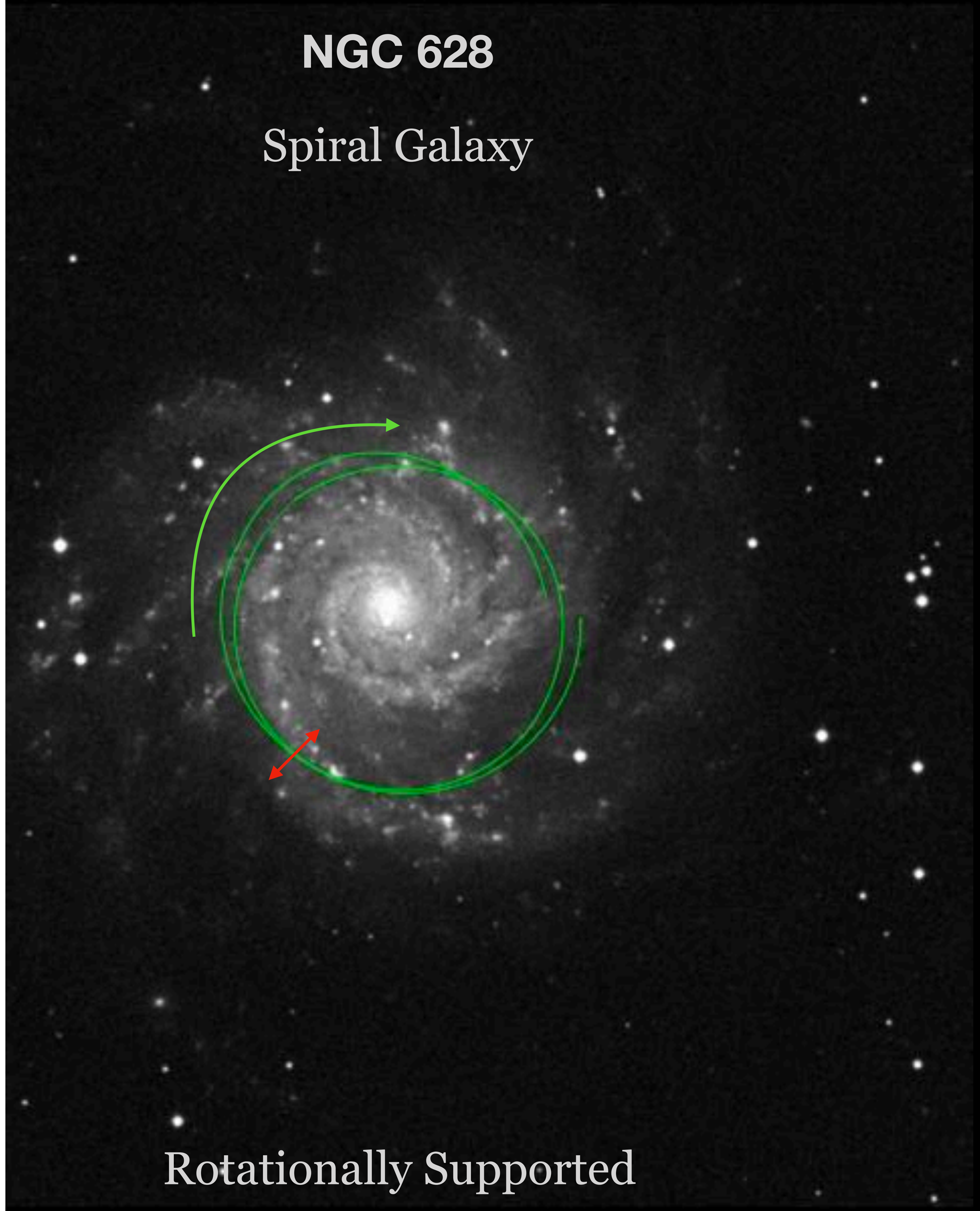
Eccentric radial orbits
Random orientations

Nearly circular orbits
Same direction, same plane

Stellar orbits in galaxies

orbital frequency Ω
round & round

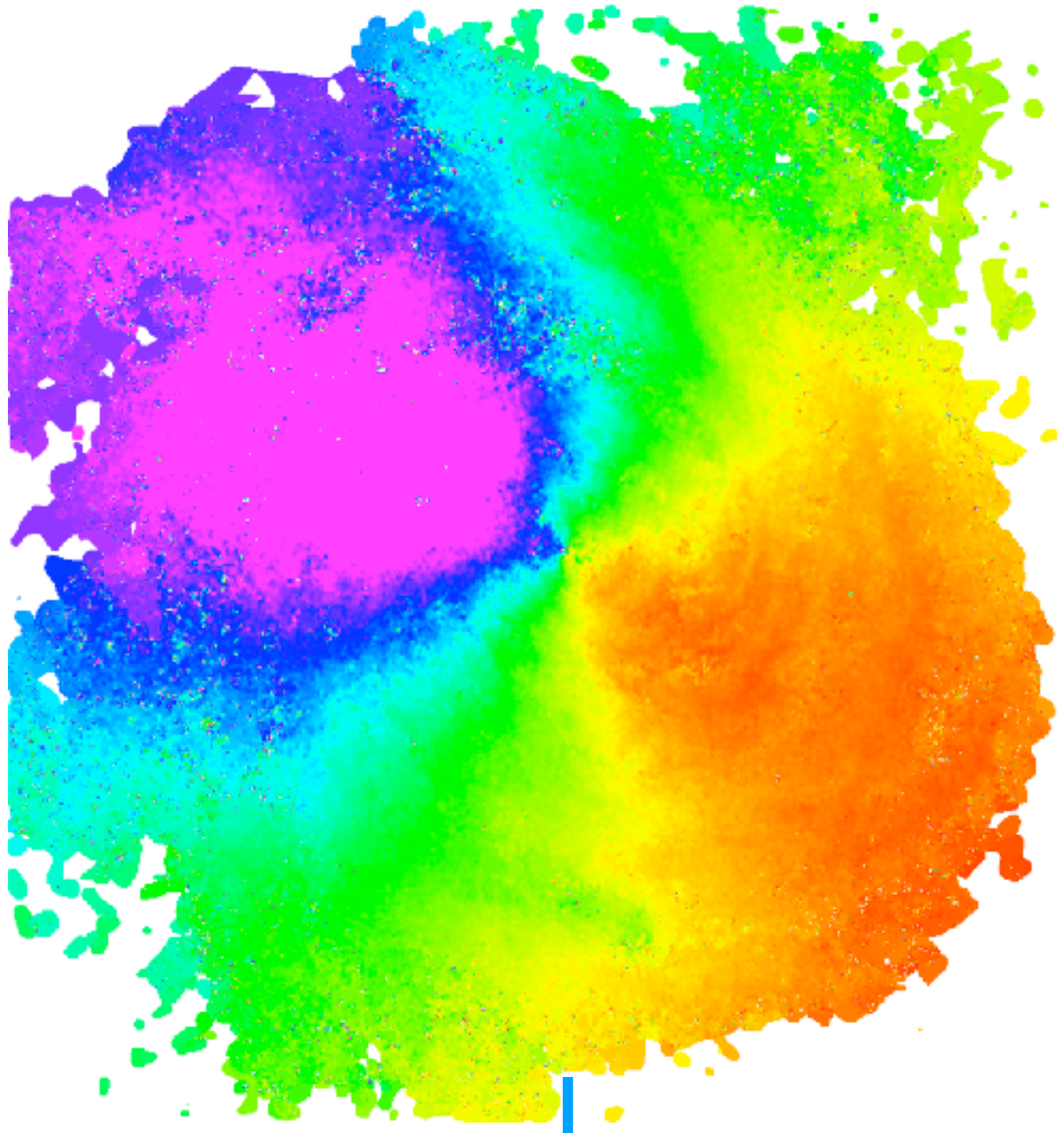
epicyclic frequency κ
in & out



Nearly circular orbits
Same direction, same plane

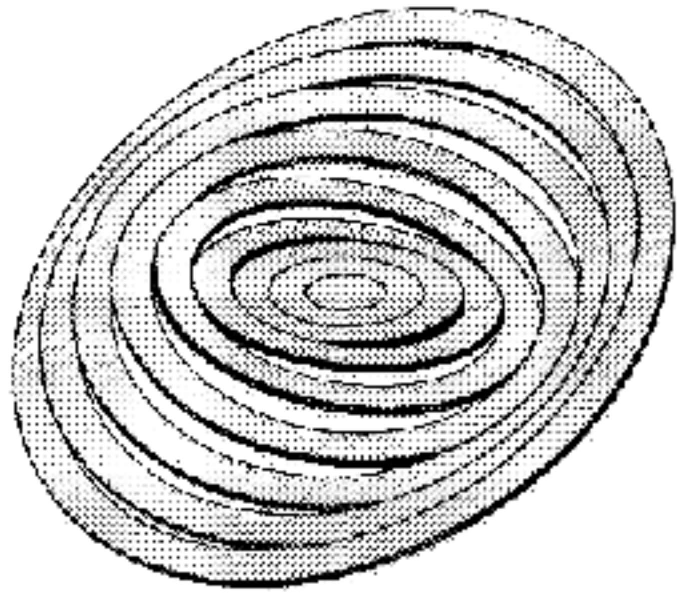
21cm interferometric observations give atomic gas distributions and velocity fields

NGC 6946



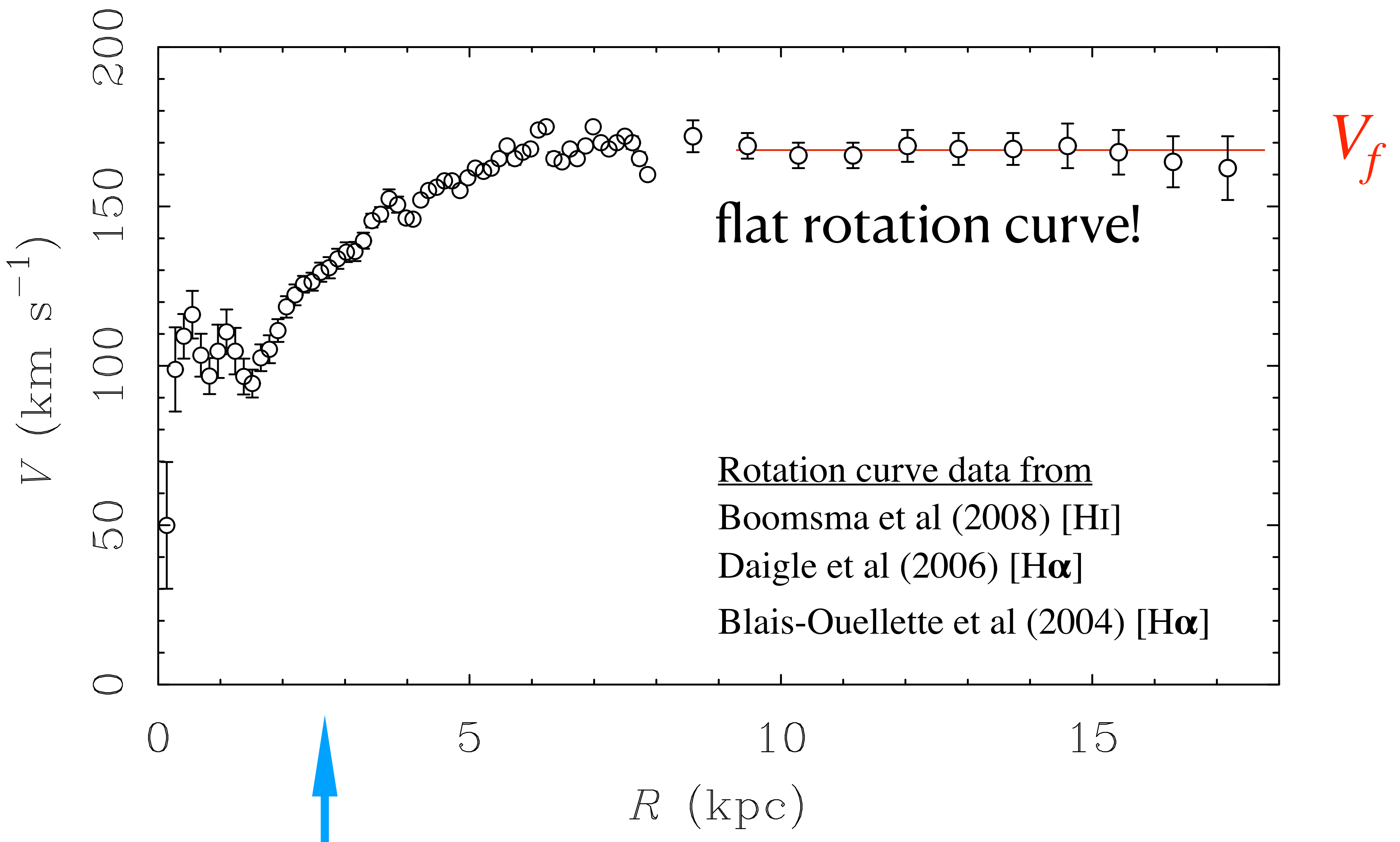
THINGS (Walter et al. 2008; de Blok et al. 2008)

tilted ring model



to which we make tilted ring fits

Rotation curve



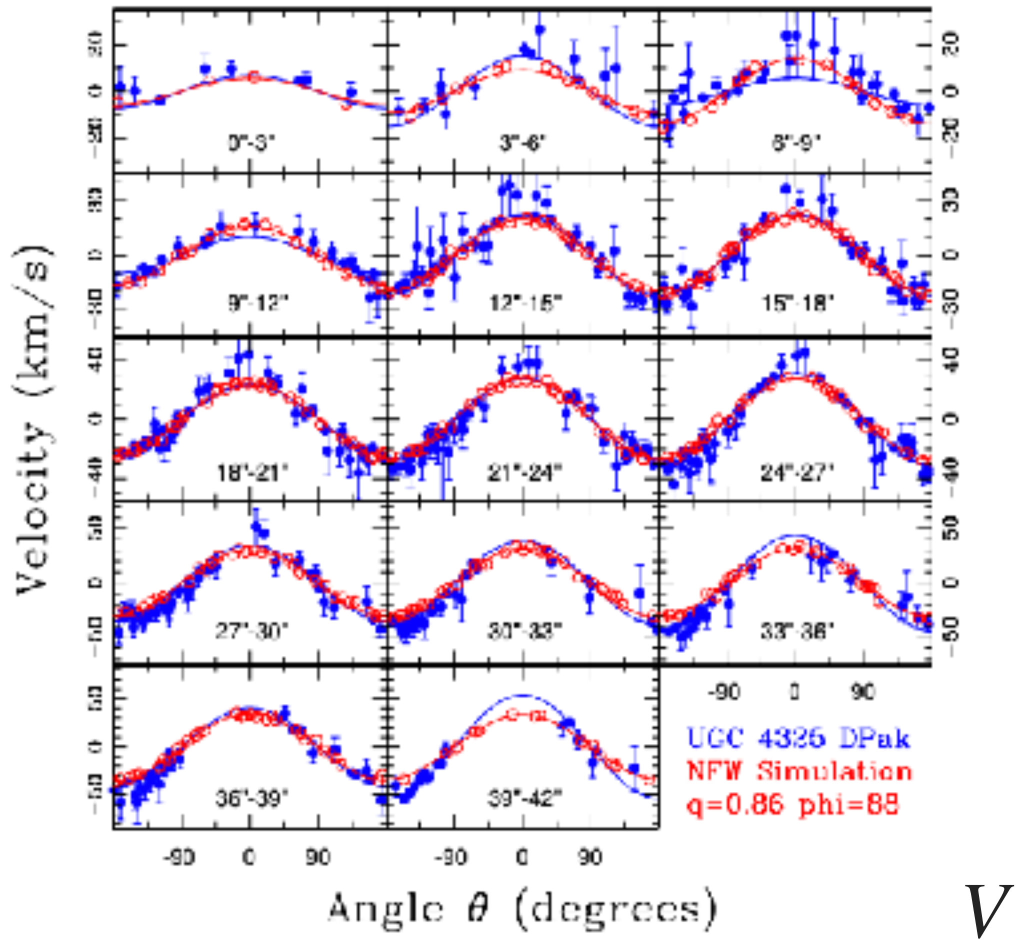
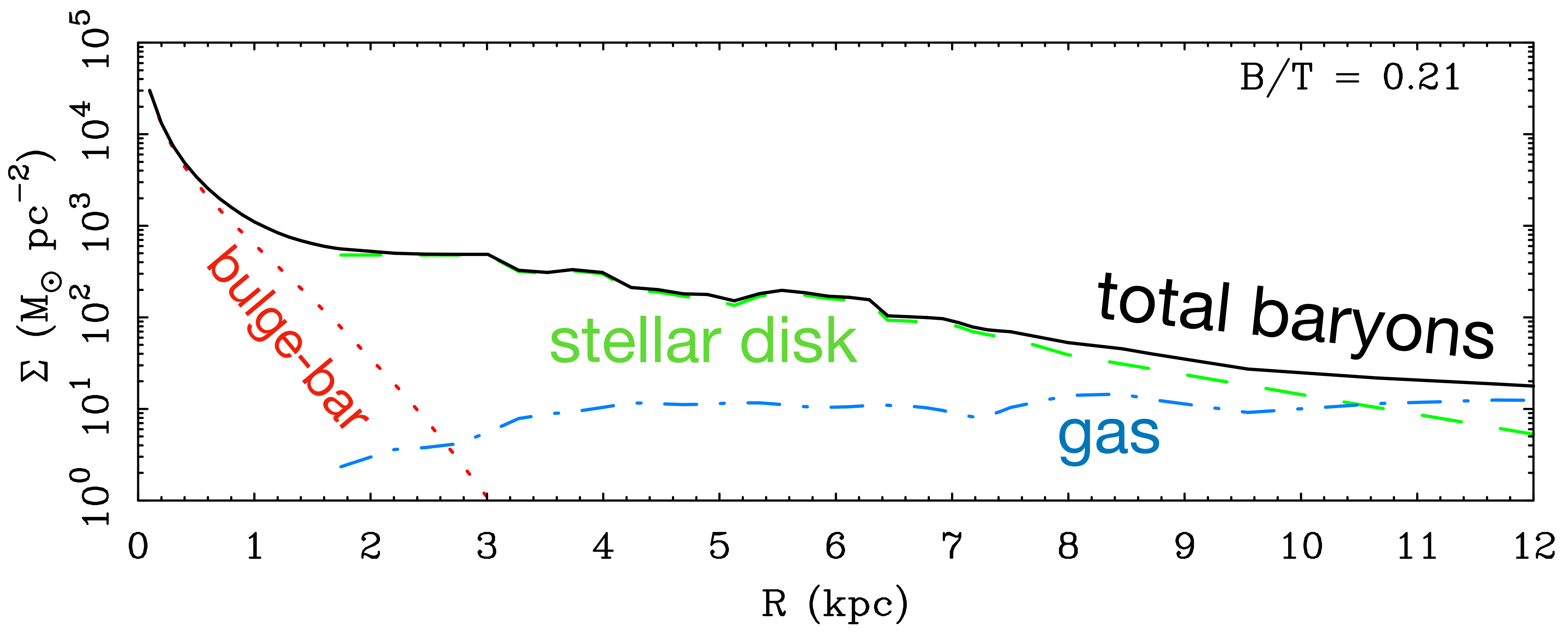


Figure 5.6: The (0.86, 88°) simulation results (red) over-plotted with the observed UGC 4325 data (blue). The simulation and data match well between ~ 12″ - 30″.

The sinusoidal variation of velocity in each ring measures the position angle, inclination, and rotation curve $V_c(R)$.

$$V \sin i = V_{sys} + V_c \cos\theta + V_r \sin\theta$$

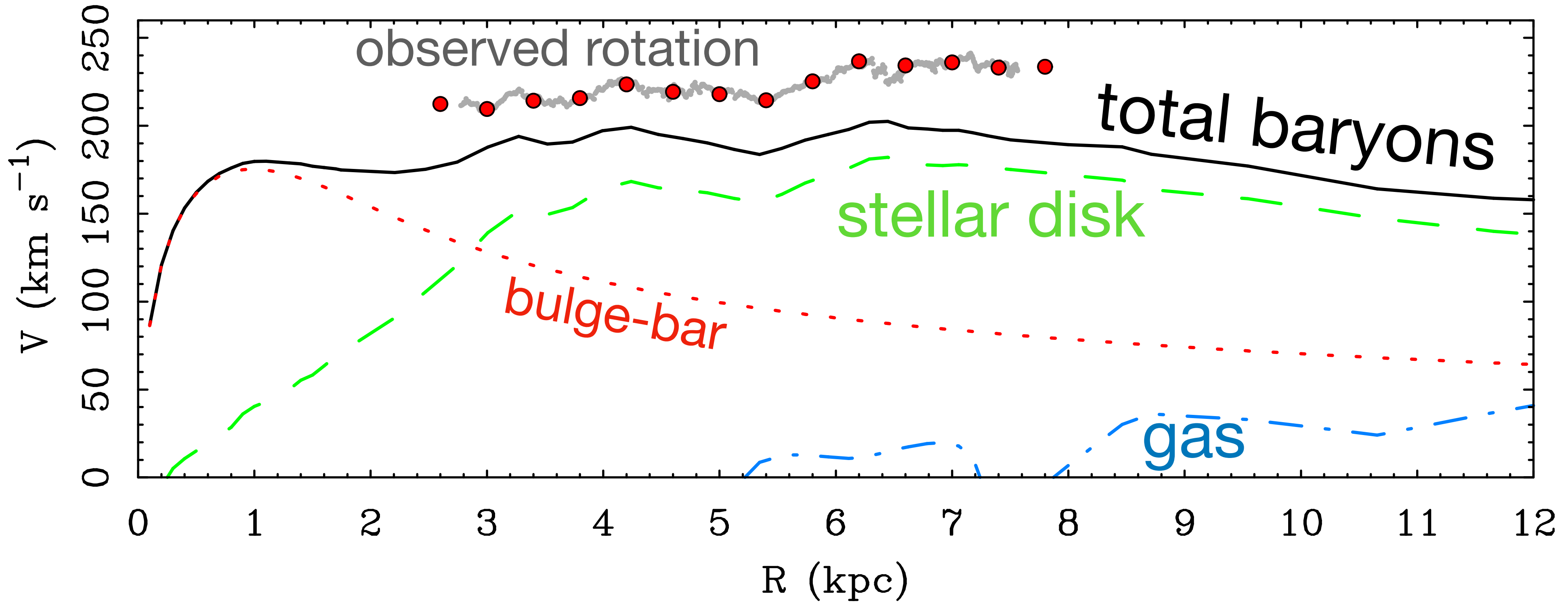
Galactic mass model



Surface density profile

$\Sigma(R)$ *from observed surface brightness profile*

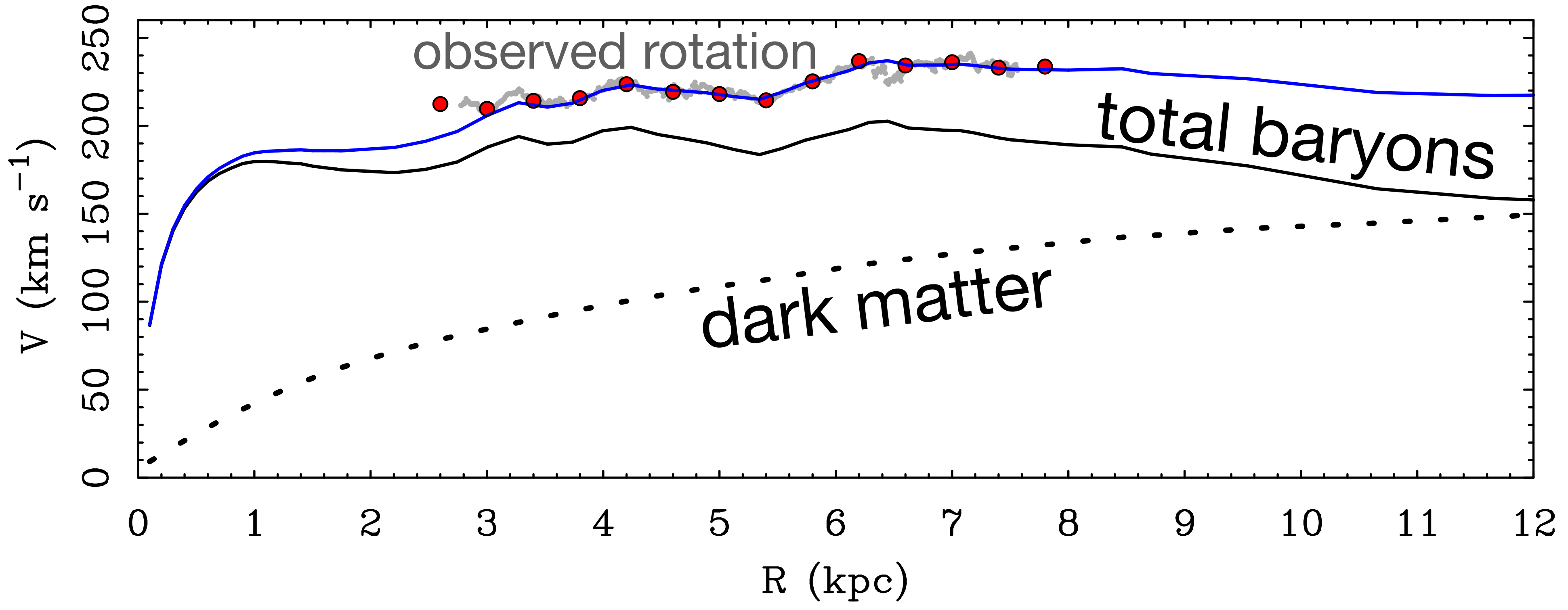
The main uncertainty is the mass-to-light ratio of the stars.



Mass model

$$\frac{V^2}{R} = -\frac{\partial \Phi}{\partial R} = 2\pi G \Sigma(R)$$

$$V_{DM}^2 = V_{obs}^2 - V_{bar}^2$$



Mass model with DM halo

Total rotation decomposed into baryonic and dark components

Galactic Kinematics

Galactic constants

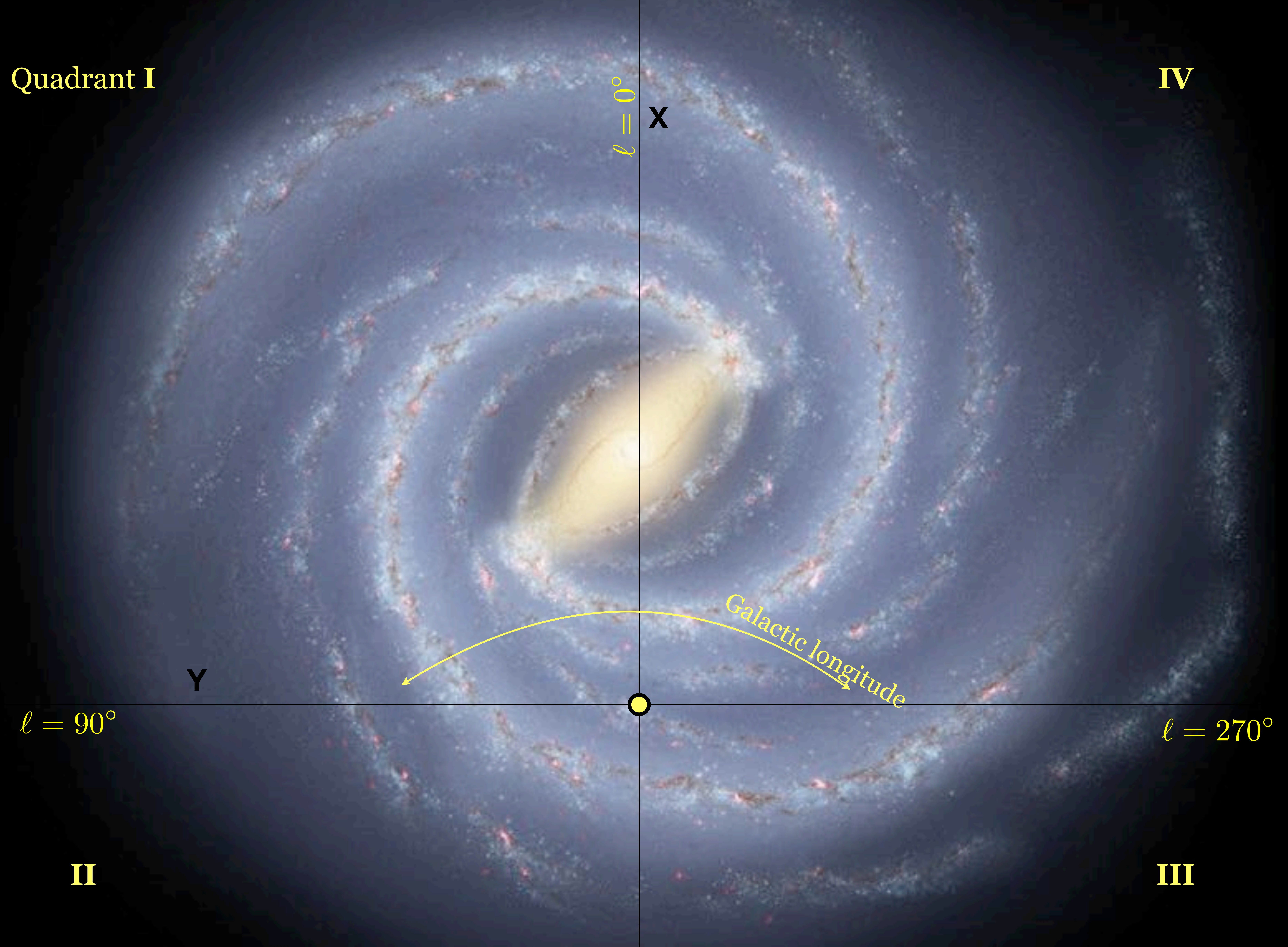
$$R_0 \; \Theta_0 \; A \; B$$

$$\Omega < \kappa < \nu_z$$

Local Standard of Rest

Epicycle approximation

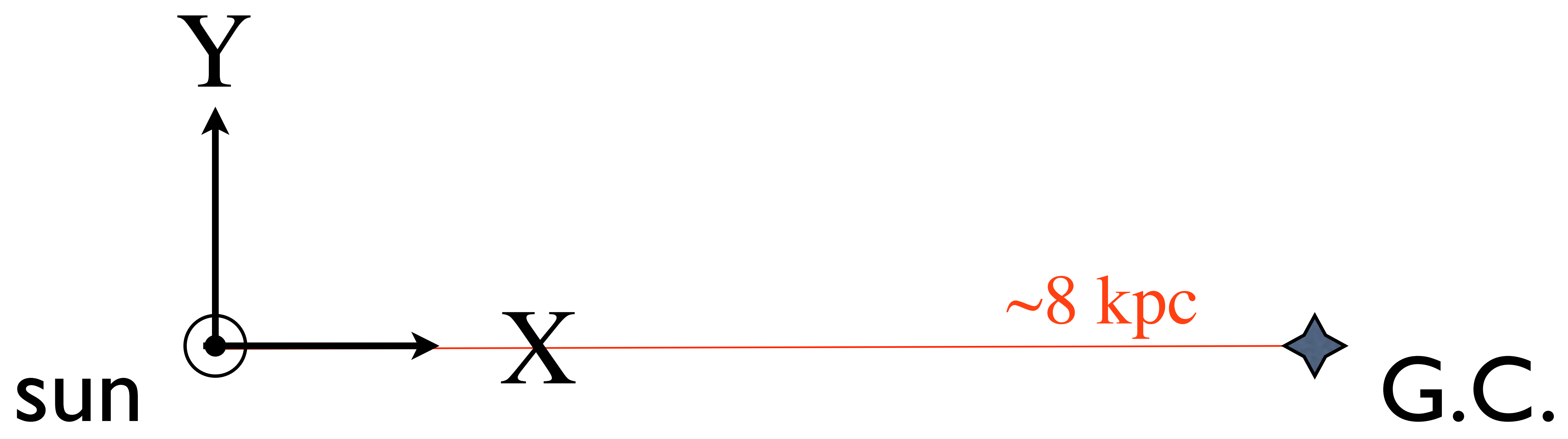
Galactic mass distribution: bulge/bar, stellar disk, gas disk, dark matter



Local Galactic Coordinates

Cartesian coordinates centered on solar system

As opposed to Galactic Center. Beware sign conventions



X, Y, Z:

X points towards the Galactic Center

Y points in direction of the sun's orbital motion

Z is perpendicular to the Galactic Plane

U, V, W are velocities in these directions

Definitions of Galactic Quantities

R_0	distance to Galactic Center		8.12 kpc
Θ_0	orbital velocity of LSR		233 km/s
Ω_0	angular velocity of LSR	$\Omega = \dfrac{V}{R}$	$P = \dfrac{2\pi R}{V} = \dfrac{2\pi}{\Omega}$
A	Oort constant A	$A = \dfrac{1}{2}\left(\dfrac{V}{R} - \dfrac{dV}{dR}\right)_{R_0}$	shear
B	Oort constant B	$B = -\dfrac{1}{2}\left(\dfrac{V}{R} + \dfrac{dV}{dR}\right)_{R_0}$	vorticity
κ	epicyclic frequency	$\kappa^2 = -4B(A - B)$	

Frequencies often expressed in Galactic units: km/s/kpc

LSR - local standard of rest

The Local Standard of Rest (LSR) is the point coincident with the sun that is on a perfectly circular orbit.

The sun itself is not on a circular orbit, nor is its mean distance from the galactic center equal to our current location.

The net velocity of populations of stars is zero wrt the LSR; this is how we define it.

More generally, if the Galactic potential is not axis-symmetric (e.g., because of the Galactic bar), then the LSR orbit is oval.

Solar Motion

The residual solar motion wrt the average of local stars is

radial	$U_\odot = 10\,\mathrm{km\,s^{-1}}$
azimuthal	$V_\odot = 12\,\mathrm{km\,s^{-1}}$
vertical	$W_\odot = 7\,\mathrm{km\,s^{-1}}$

Some say V = 5 km/s,
some say 15 km/s!

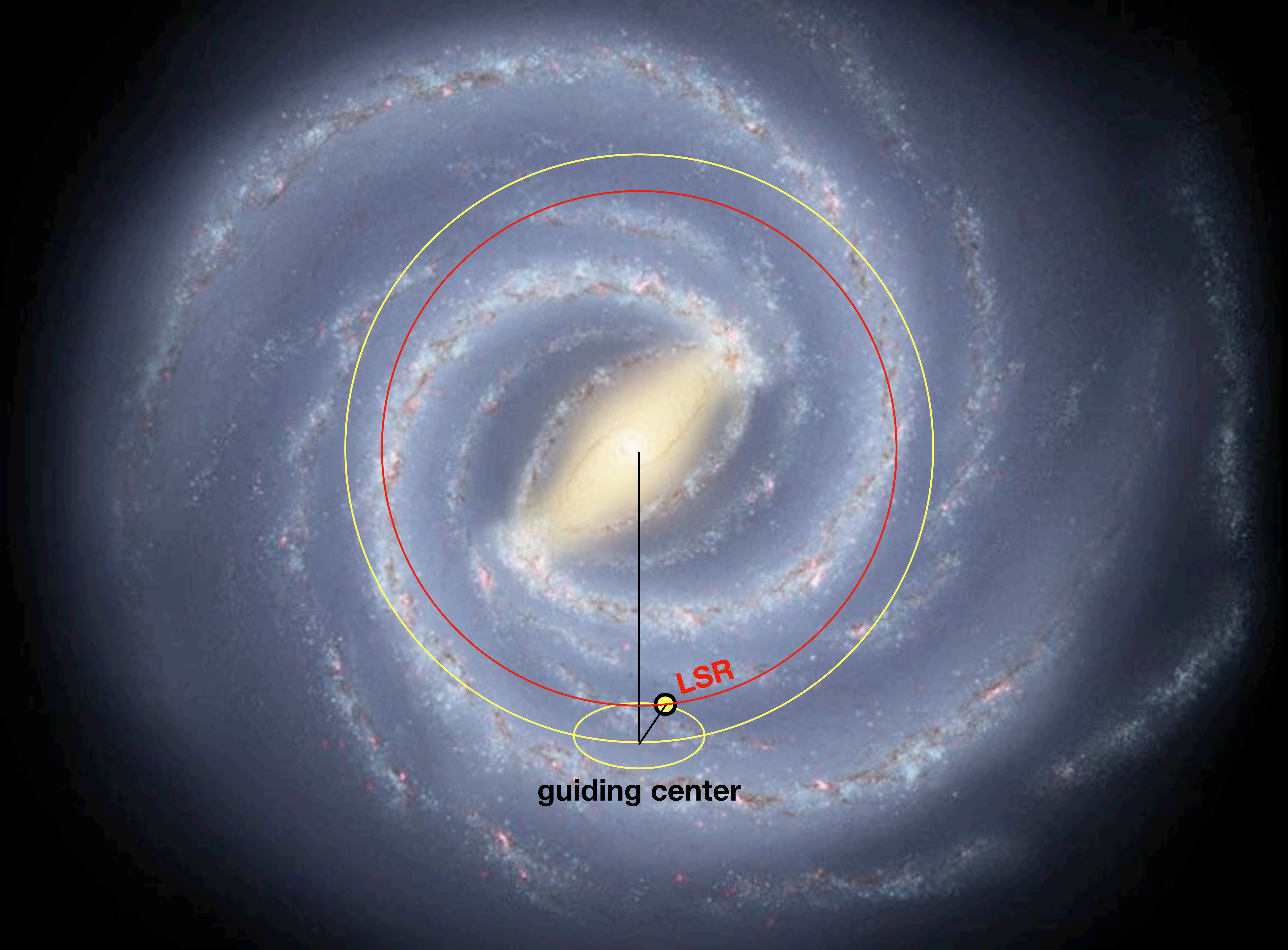
The Sun is moving

- a bit towards the galactic center
- faster than the LSR
- northward out of the galactic plane

Currently we are near the mid-plane

(Remember this doesn't account for the rotation of the disk!)

Orbits of individual stars: the epicycle approximation



Orbits for 4 individual stars

orbital eccentricity correlates with age

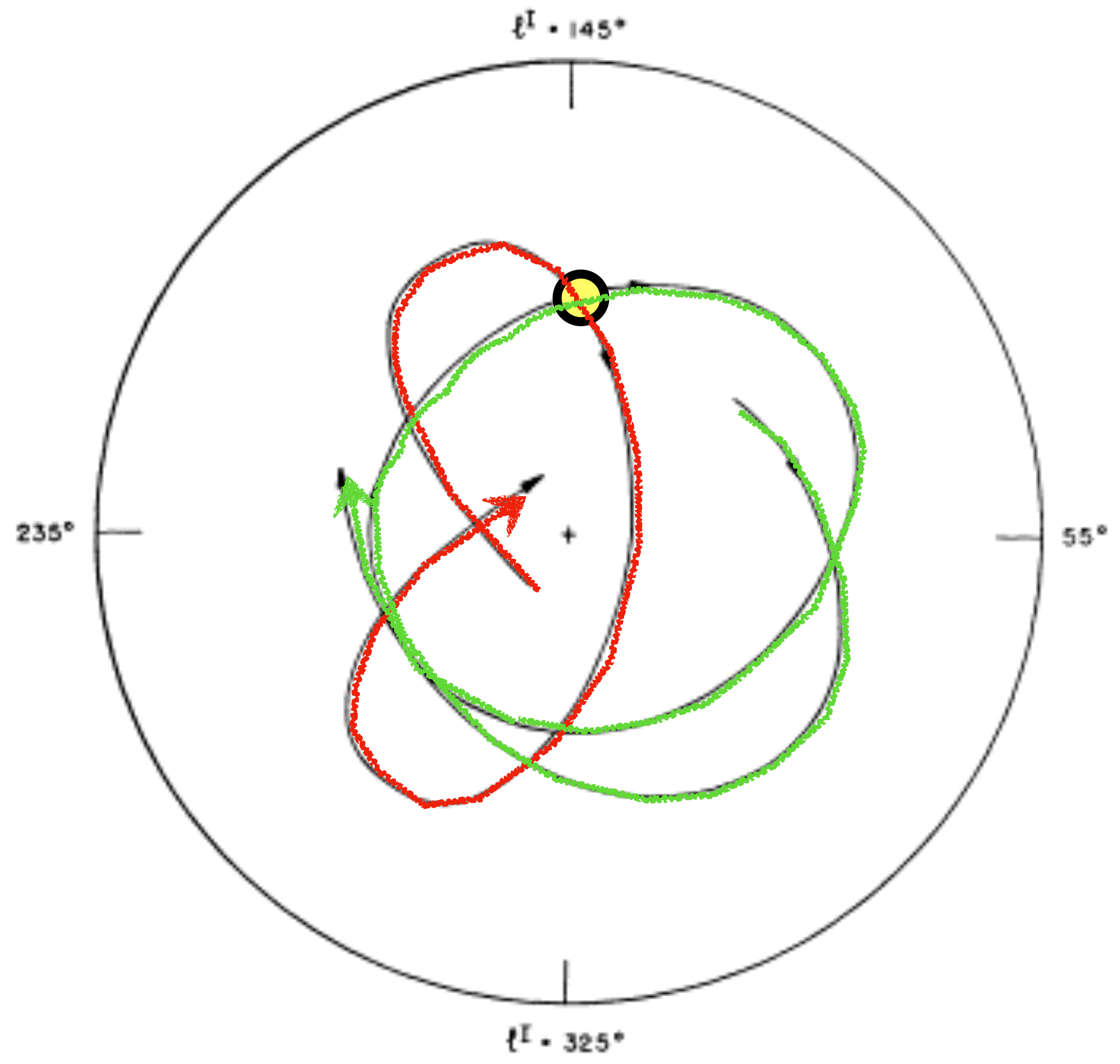


FIG. 2.—Segments of the galactic orbits for two of the program stars. The more circular orbit is for HD 117635 with an ultraviolet excess of $\delta = +0^{m}05$ The more elliptical orbit is for HD 11980 with $\delta = +0^{m}17$. Both orbits pass through the solar neighborhood, which is designated by a circle on the $l = 145°$ axis at a distance of 10 kpc from the galactic center The galactic center is shown as a cross The outer circle has a radius of 20 kpc.

from Eggen, Lynden-Bell, & Sandage (1962)

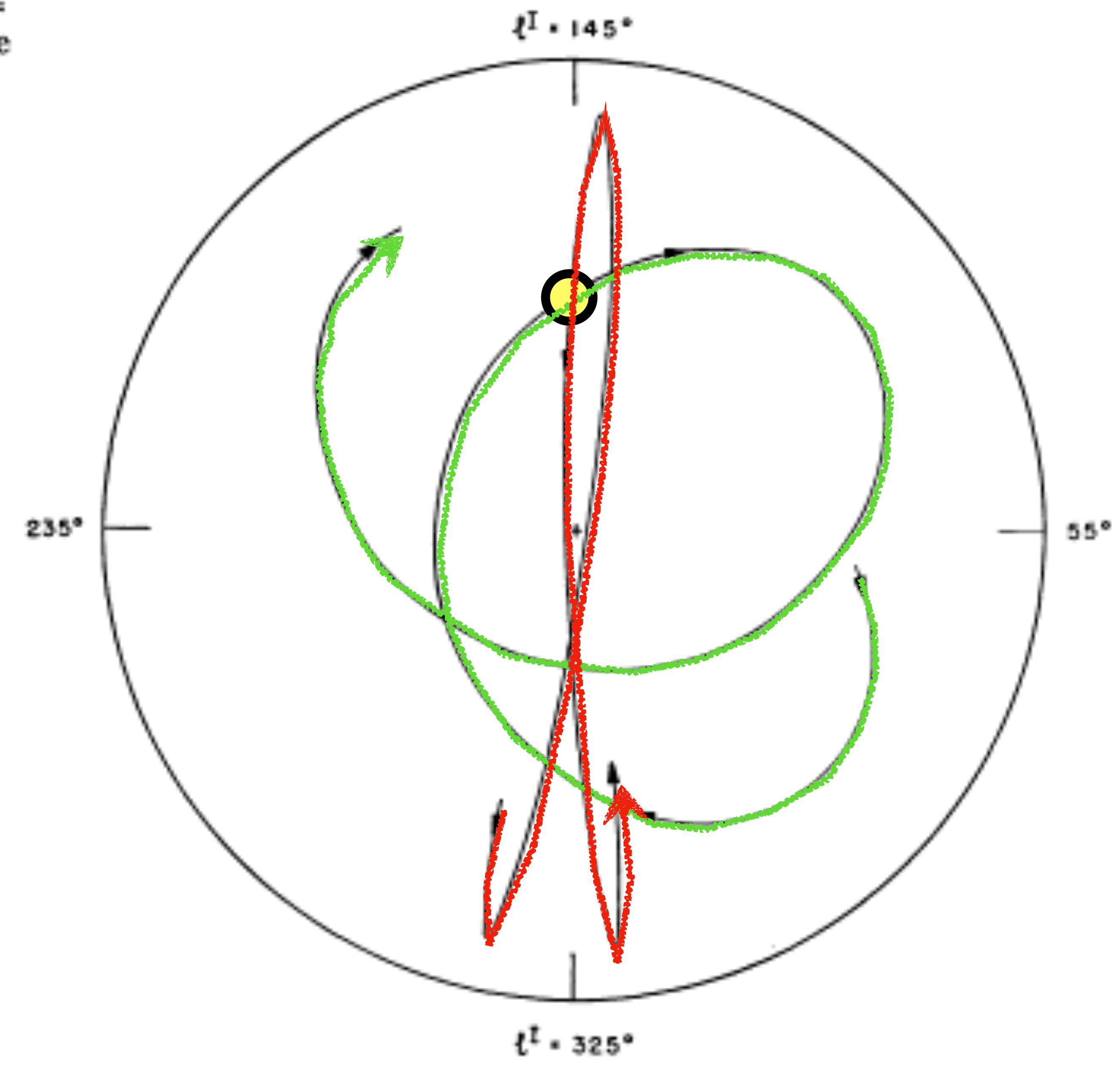


FIG 3.—Same as Fig. 2 The more circular orbit is for HD 29587 with $\delta = +0^{m}13$. The more elliptical orbit is for Ross 106 with $\delta = +0^{m}26$. The orbit for Ross 106 is retrograde.

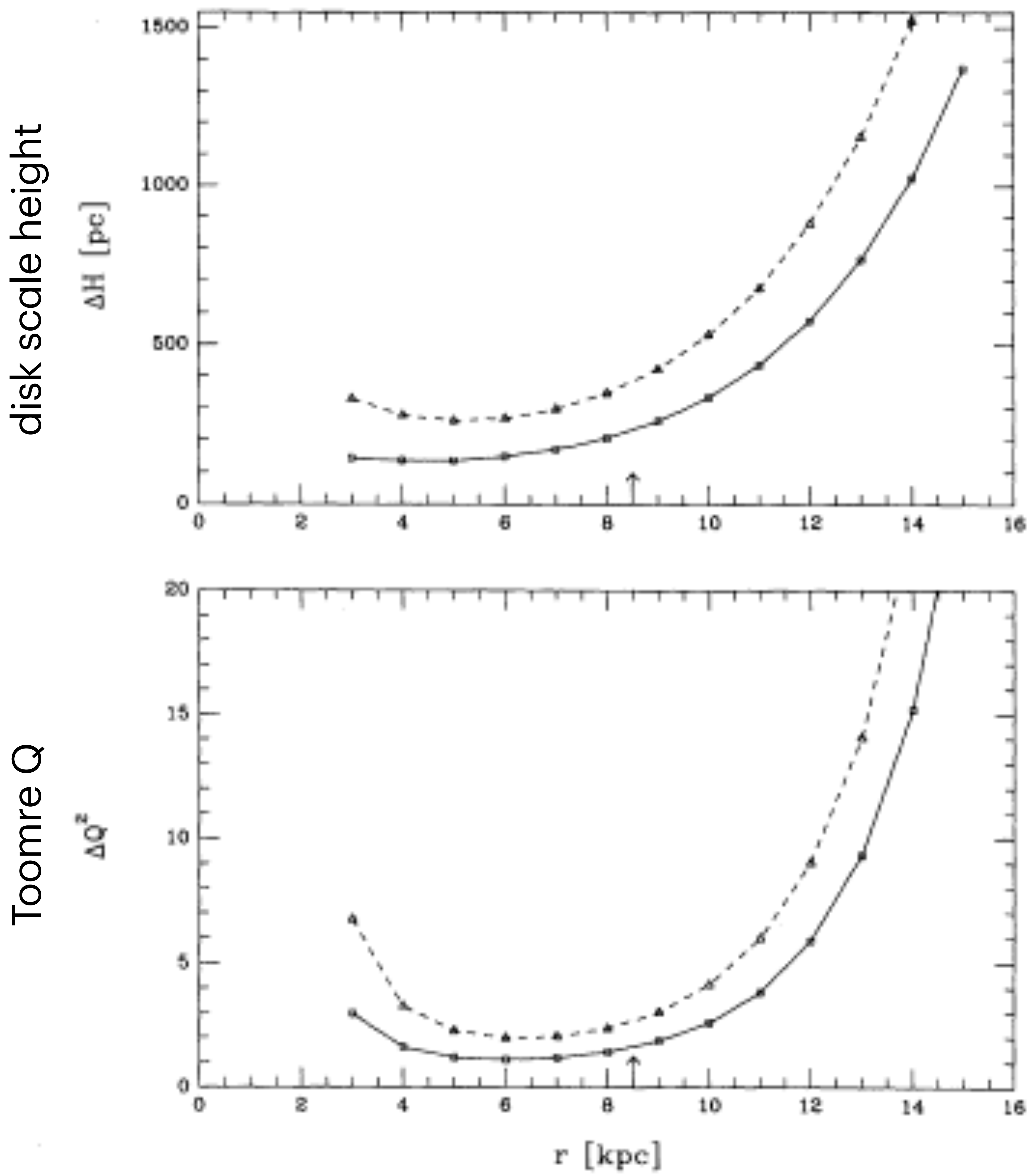
Disk Heating

Limits on mergers & such

- Spiral disks are dynamically cold
- Mergers heat and thicken disks
  - like dropping stones in a pond
- Places limit on mergers
- Milky Way has not experienced a substantial merger (mass ratio > 1:10) for a long time
  - Thick disk ~8 Gyr old

Also constrains nature of dark matter. A DM halo made of million solar mass black holes would stir up the disk too much.

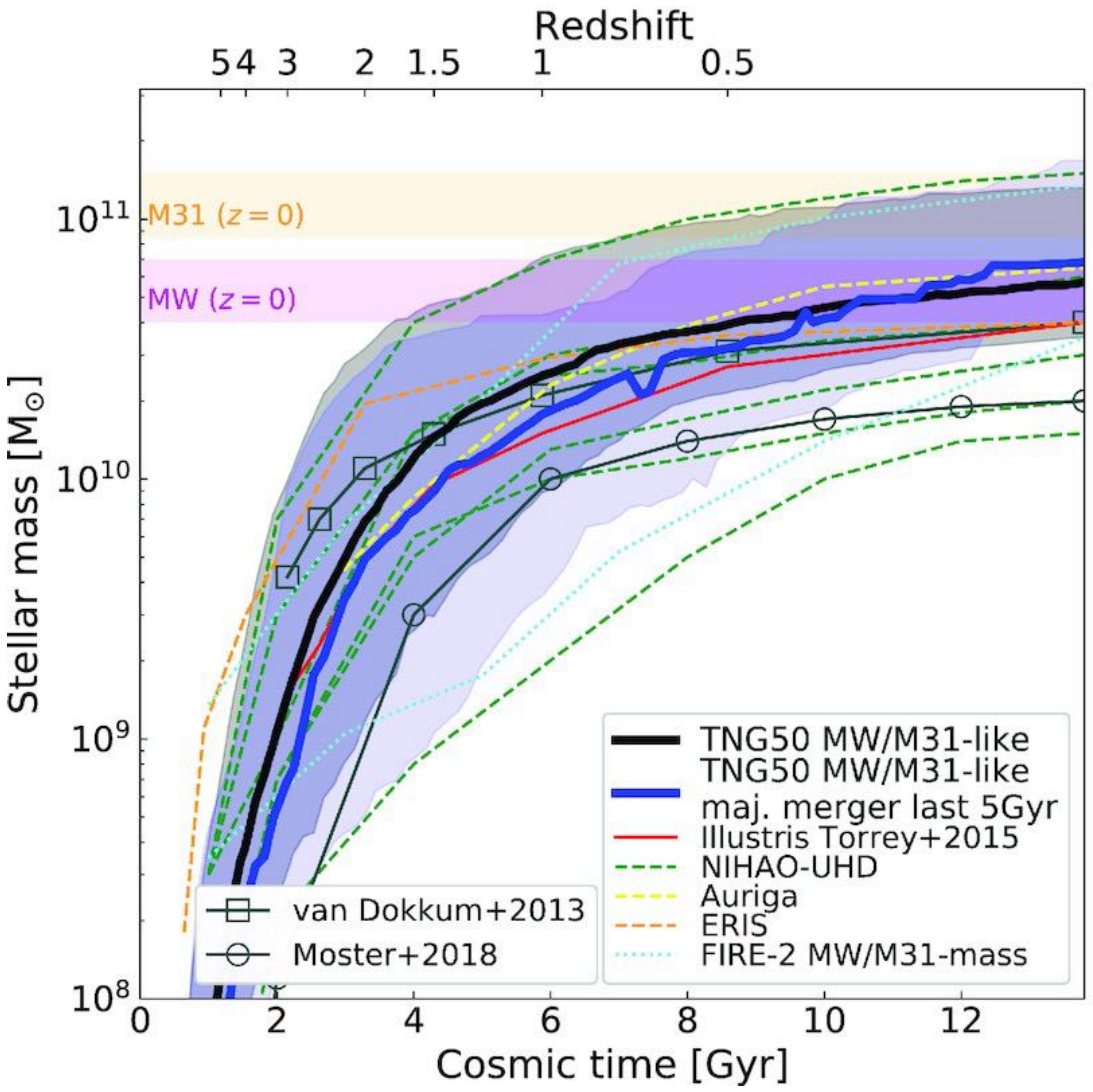
Toth & Ostriker (1992)

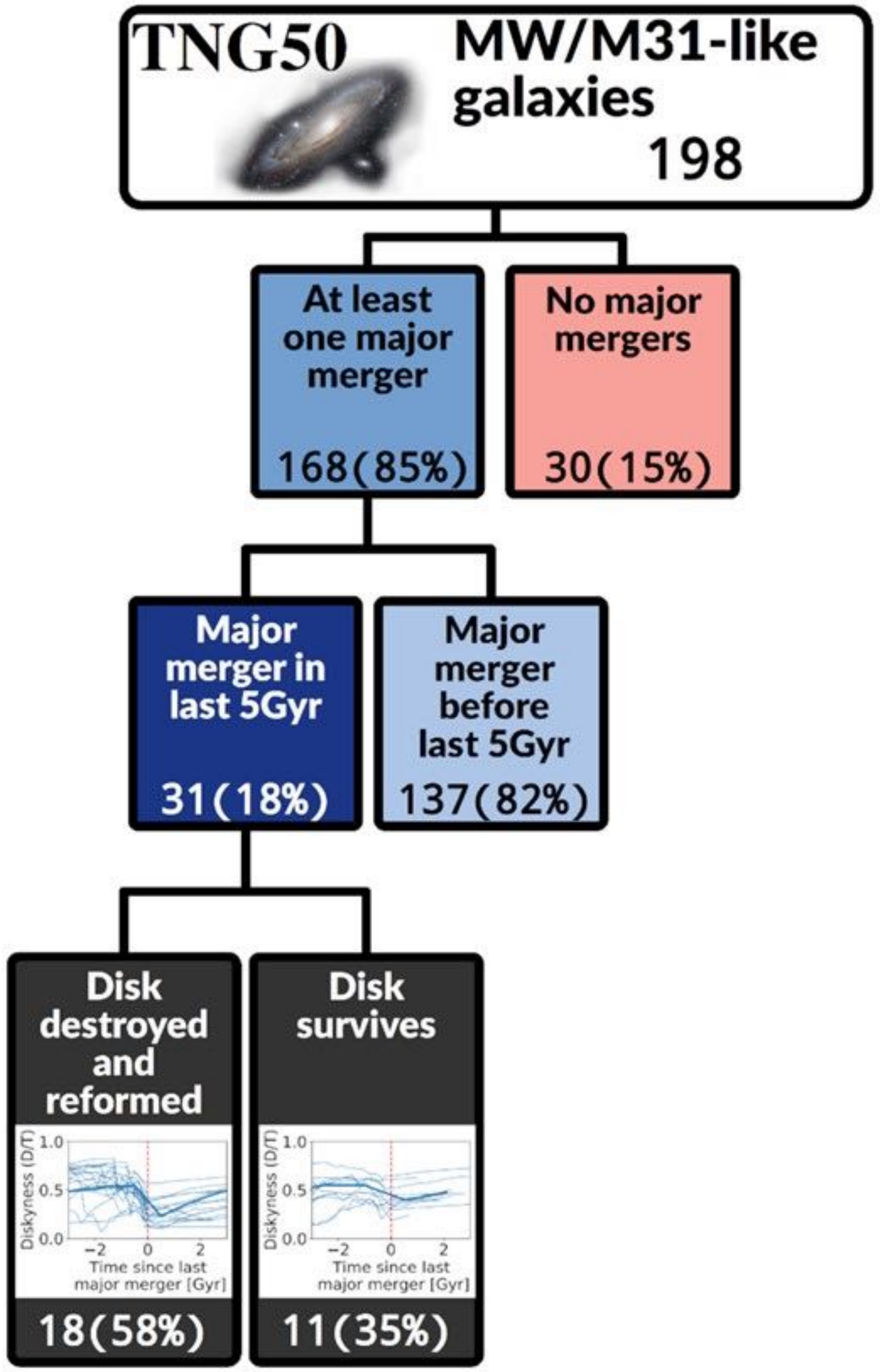


Disk thickening (top) and heating (bottom) due to 1:10 mass ratio merger. Also depends on the concentration of the merging object (solid line fluffy, dashed line dense).

Sotillo-Ramos+ (2022)

What do we expect in LCDM?





Makes a prediction for the ages and orbits of stars -
Disks lacking stars older than 5 Gyr should be common