

Today

 Moons of the solar system

Events

- Homework 5
 - Due next time

Round objects in the solar system with diameter < 10,000 km



Selected Moons of the Solar System, with Earth for Scale



Earth

Obvious Definition

• A moon is an object that orbits a planet

Sizes of Moons

- Small moons (< 300 km)
 No geological activity
- Medium-sized moons (300–1,500 km)
 - Geological activity in past
- Large moons (> 1,500 km)
 Ongoing geological activity

crudely speaking



Moons of the Gas Giants (Medium and Large)

- Enough self-gravity to be spherical
- Have substantial amounts of ice - as important as rock to overall composition
- Circular orbits mostly prograde (in the same direction as planet rotates)
- Formed in orbit around jovian planets

Rocks and metals condense, hydrogen compounds stay vaporized.

Hydrogen compounds, rocks, and metals condense.

Inside frost line: terrestrial planets

Beyond frost line: Gas giants, icy moons, dwarf planets, comets



Medium and Large Moons

- Density
 - -low
 - typically ~2 g/cc
 - more than Gas giants
 - less than Terrestrials
- Composition
 - rock
 - ice / subsurface water

Ice is just another common "rock" mineral in the outer solar system.

Small Moons



- Far more numerous than the medium and large moons
- Not enough gravity to be spherical: "potato-shaped"
- Often just captured asteroids

The moons of Jupiter



Io



• Io is the most volcanically active body in the solar system.

Volcanic activity discovered on Io during the Voyager fly-by

What're the odds?



Io's Volcanoes



• Volcanic eruptions continue to change lo's surface.

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optical

infrared



dark volcanic craters in the optical correspond to hot spots in the infrared

show interactive optical/IR image



lo's surface very young

Constantly re-covered in fresh lava & sulfur dioxide snow



Tidal Heating

small tidal bulges

Io is squished and stretched as it orbits Jupiter. larger tidal bulges when closer to Jupiter

Jupiter

Orbit is elliptical because of orbital resonances with other moons



show interactive figure



• Volcanic eruptions continue to change Io's surface.





Europa's interior also warmed by tidal heating.



Europa may have a 100-km-thick ocean under an icy crust.

Rising plumes of warm water may sometimes create lakes within the ice, causing the crust above to crack . . .

... explaining surface terrain that looks like a jumble of icebergs suspended in a place where liquid or slushy water froze.

Tidal stresses crack Europa's surface ice.





Europa

- Icy surface
 - cracks driven by tidally driven "geological" activity
 Liquid ocean
 - beneath?
 - popular spot to
 speculate about
 the potential for
 life

2001 Space Odyssey written when 2001 was far in the future



There are serious proposals to send a robotic submersible to Europa.

2001 warned against that because this thing will eat you. ϕ



Ganymede



- Largest moon in the solar system
- Clear evidence of geological activity
- Salty ocean under thick crust of ice
- Tidal heating plus heat from radioactive decay?

Ganymede Interior



The moons of the Jovian planets



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Callisto



"Classic" cratered
 iceball

 No tidal heating, no orbital resonances

Saturn

- Has one large moon - Titan
- a large number of medium-sized and small moons
- Rings composed of many tiny icy moonlets

Saturn's large moon Titan



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- Titan is the only moon in the solar system which has a thick atmosphere.
- It has a thick haze layer that obscures the surface at optical wavelengths.

Saturn's large moon Titan



- Atmospheric composition:
 - $-90\% N_2$
 - 5% Argon
 - 5% CH₄ (methane)
 - other hydrocarbons
- Hazy

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Saturn's large moon Titan



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- Relative to Earth:
 - 1.5 Atm pressure
 - 4x denser
 - comparable total
 mass (1.2x)
 - more extended
 - due to lower gravity
- Cold
 - -180° C

Titan is

- Big for a moon, and
- cold can retain an atmosphere



atmospheric haze in optical light

Underneath the atmosphere is terrain, including seas of liquid hydrocarbons solid surface

liquid methane Some transparent windows in the infrared. Reveals widespread lakes of liquid methane. Weather on Titan involves methane clouds and rain.

Titan's Surface



- The *Huygens* probe provided a first look at Titan's surface in early 2005.
- It had liquid methane, "rocks" made of ice. Huygens descent movie <u>https://www.youtube.com/watch?v=HtYDPj6eFLc</u>

https://www.youtube.com/watch?v=bS9wlVsFlzA



 Almost all of them show evidence of past volcanism and/or tectonics.



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 Mimas has a big crater that makes it look like the Death Star.

 Ice fountains of Enceladus suggest it may have a subsurface ocean.







 Iapetus is dark on one side & bright on the other. It seems to have collected a goo of space debris emitted by Phoebe on the leading (dark) side of its orbit.

 Iapetus has a curious ridge around much of its equator



Small moons can also be weird. Hyperion looks like a sponge.







Medium Moons of Uranus



 They have varying amounts of geological activity.

 Miranda has large tectonic features and few craters (possibly indicating an episode of tidal heating in past). https://www.youtube.com/watch?v=v6Bt7u-EQHM

Moons of Neptune



cantaloupe terrain

Triton

Neptune's Moon Triton



Triton's southern hemisphere as seen by Voyager 2.



This close-up shows lava-filled impact basins similar to the lunar maria, but the lava was water or slush rather than molten rock.

The occasional geyser, heated by sunlight, streaks the downwind terrain with dark material Similar to Pluto, but larger

- Evidence for past geological activity
- orbits retrograde
 - unique for such a large moon
 - may have been a binary partner of Pluto captured by Neptune





Why are small icy moons more geologically active than small rocky planets?



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Rocky Planets versus Icy Moons



- Rock melts at higher temperatures.
- Only large rocky planets have enough heat for activity.
- Ice melts at lower temperatures.
- Tidal heating can melt internal ice, driving activity.