Dark Matter & MOND: Up the Proverbial Creek

STACY MCGAUGH

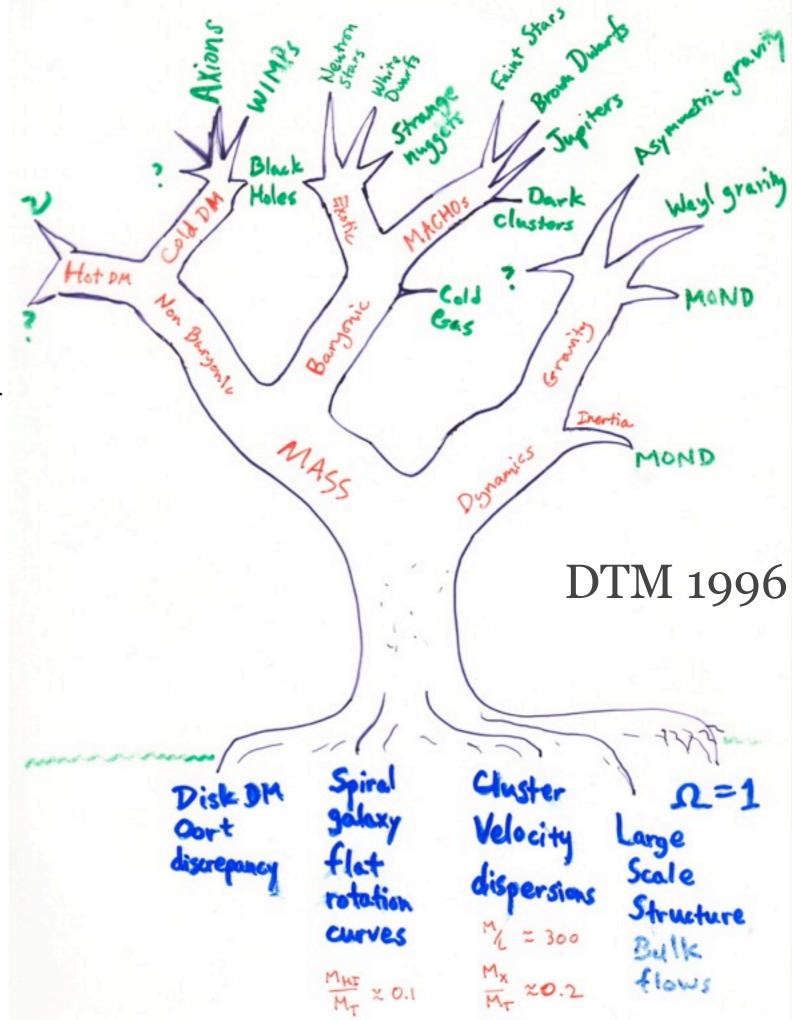


WITH COLLABORATORS
FEDERICO LELLI (CWRU)
MARCEL PAWLOWSKI (CWRU)
JIM SCHOMBERT (U. OREGON)

AND SUPPORT FROM

NSF

NASA
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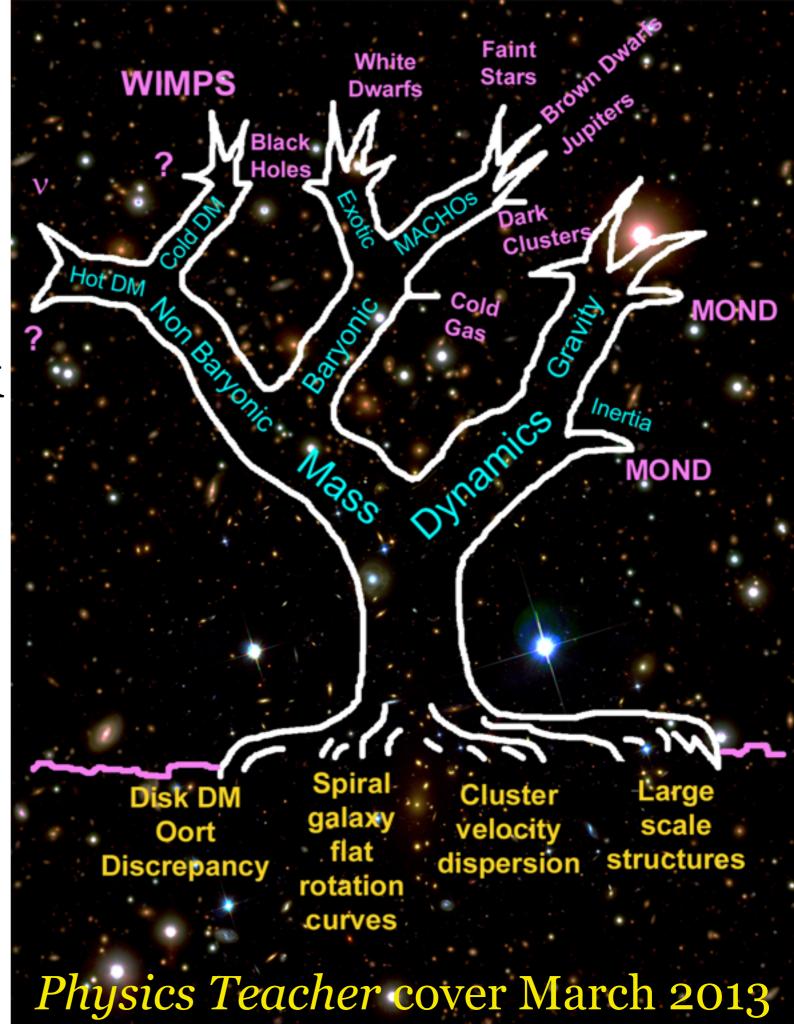


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A Field of Two Attitudes



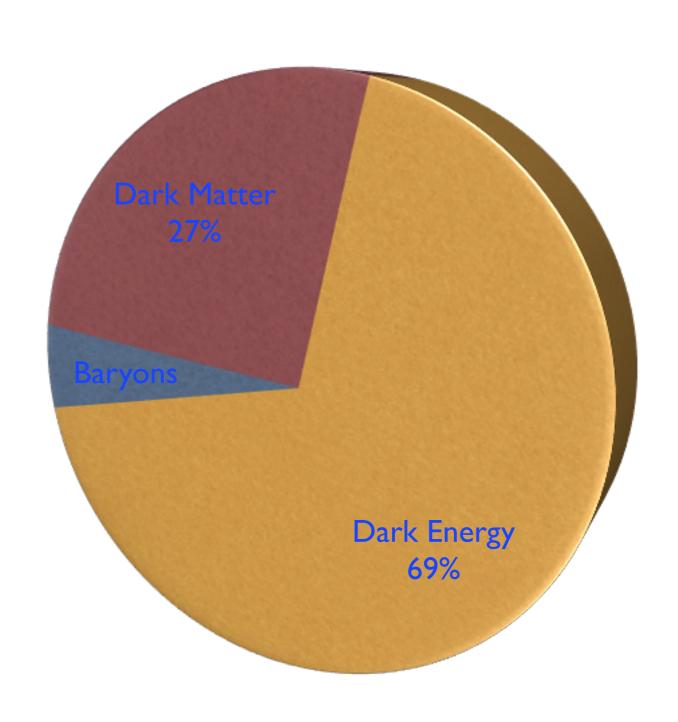
ACDM is obviously correct. Or not.

ACDM is the current cosmological paradigm.

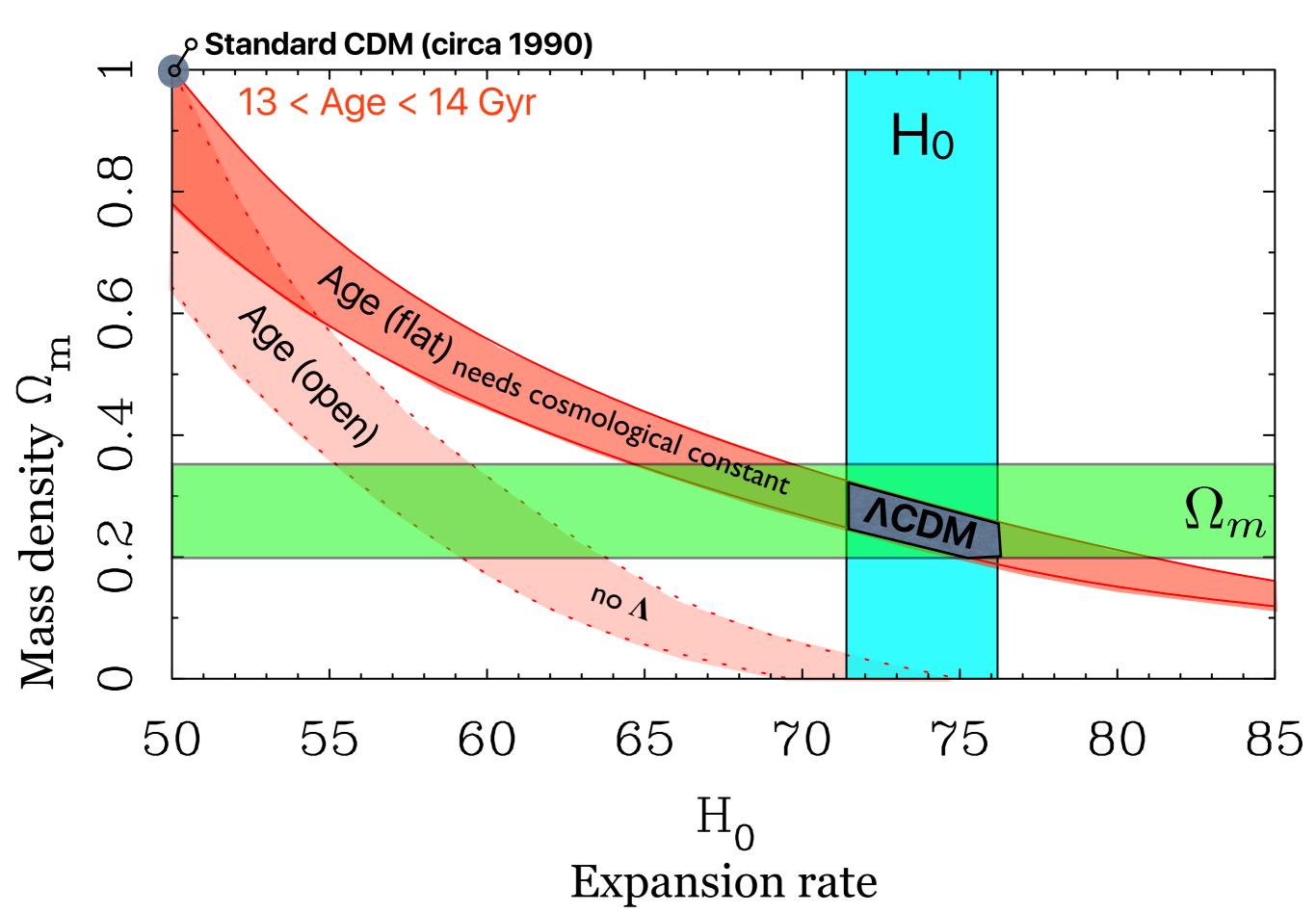
The product of decades of large scale measurements

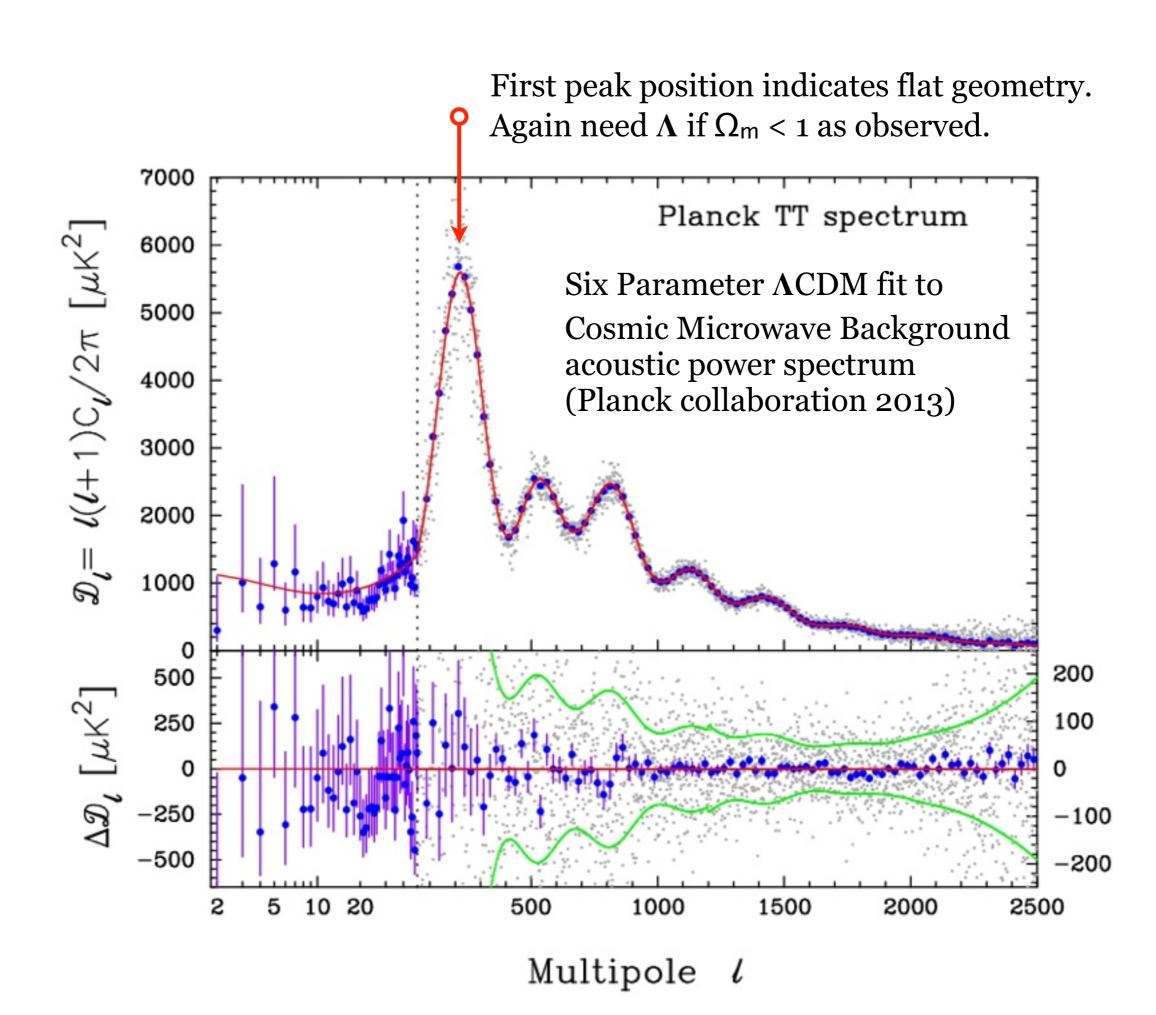
Modern cosmology nicely explains

- Expanding Universe
 - redshift-distance relation
 - geometry of space-time
- Finite Age (~ 14 Billion years)
- Early hot phase (Big Bang)
 - Nucleosynthesis of the light elements (H, He, Li)
 - Cosmic Microwave Background

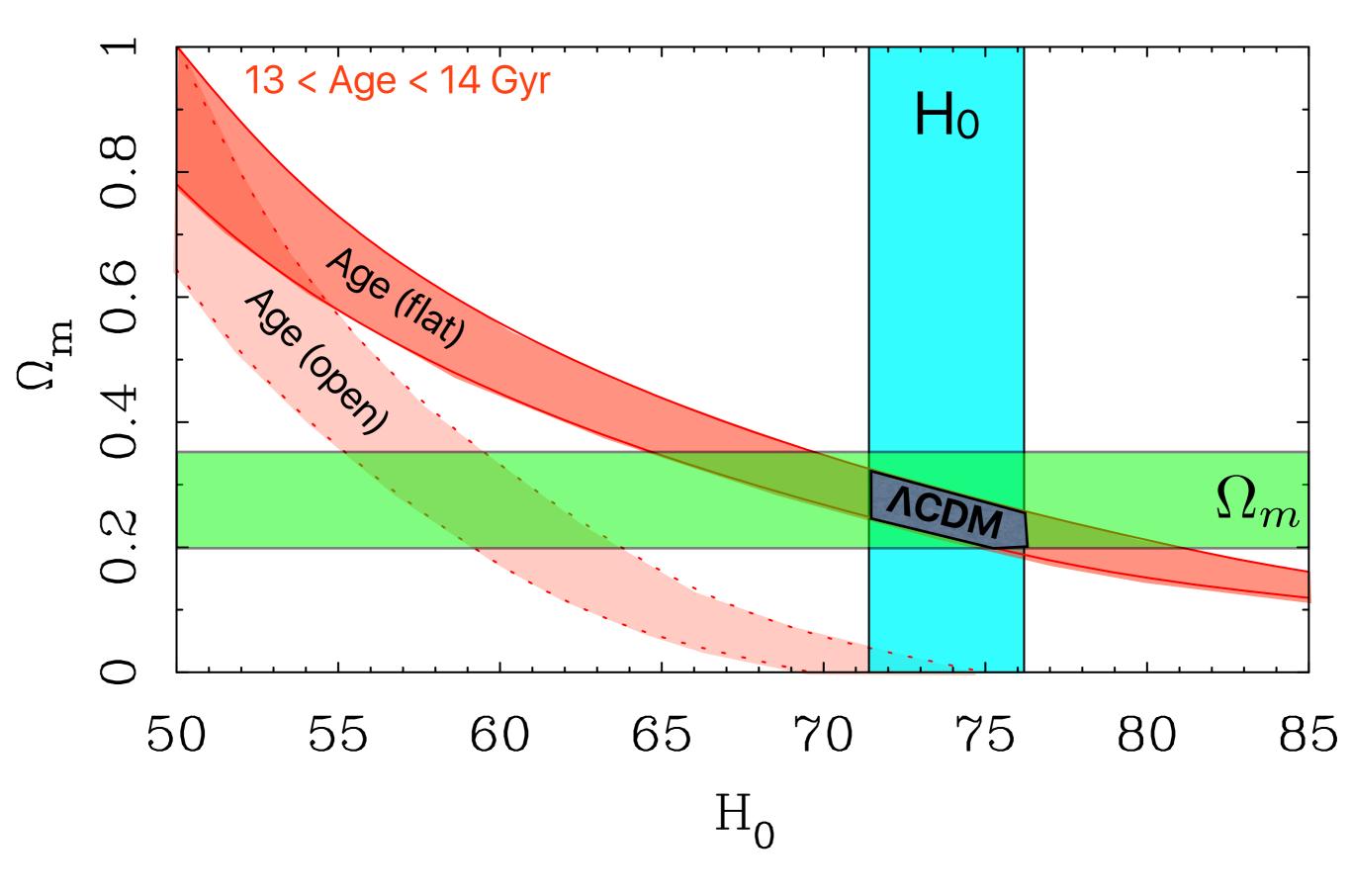


Cosmology predating SN, CMB (circa 1995)

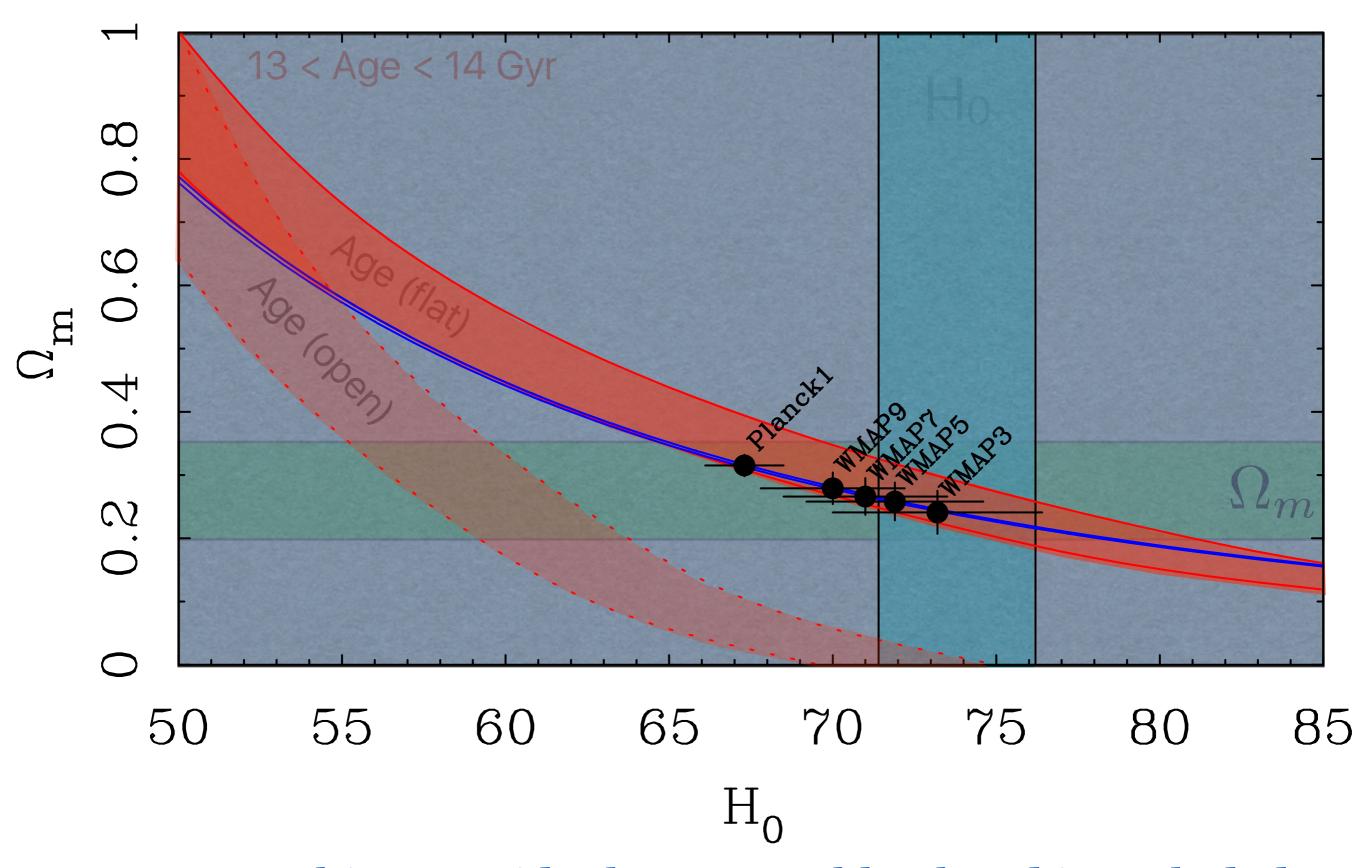




Cosmology predating SN, CMB (circa 1995)



2013 Planck constraint: $\Omega_m h^3 = 0.0959 \pm 0.0006$



Everything outside the narrow blue band is excluded

Modern cosmology only works with

- non-baryonic cold dark matter
 - whatever it is



We have direct knowledge of only a few % of the total mass-energy density of the universe

BLOOM COUNTY





by Berkeley Breathed



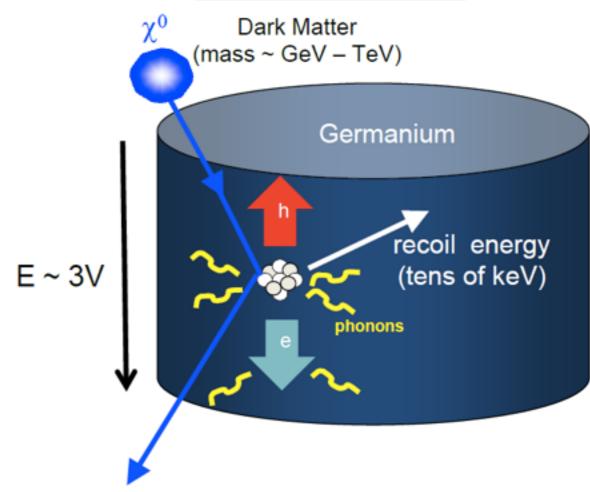


The most favored dark matter candidate is the

WIMP -

Weakly Interacting Massive Particle - thought to reside in the supersymmetric sector (which itself remains hypothetical)

WIMP detection



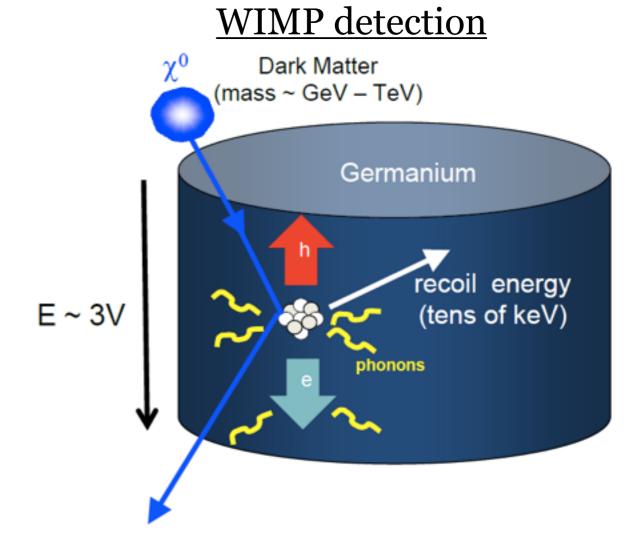
Experimental results to date (early 2016): nada

Direct detection: Many, many experiments

CDMS, LUX, XENON, DAMA, etc., etc.

Basic idea: WIMP passing through detector interacts via weak force; scatters off nucleus. Detect deposited energy of recoil. (analogous to neutrino detection).





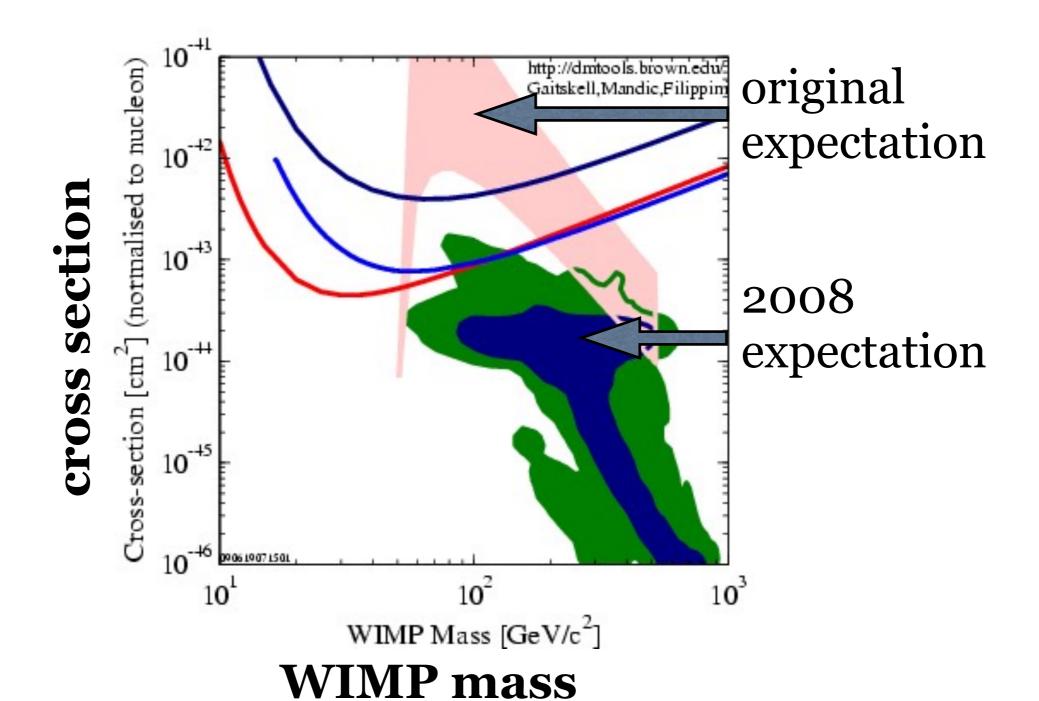
Searches for

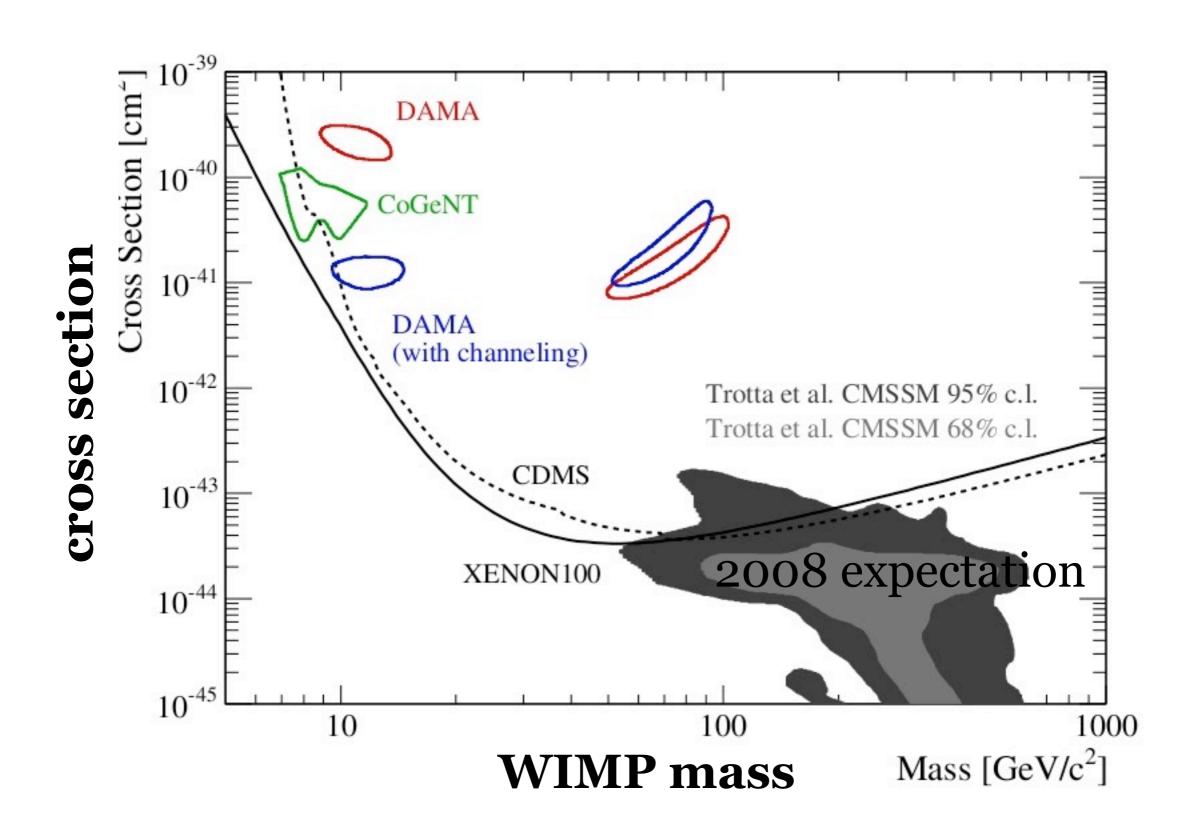
Cold Dark Matter

usually expressed as exclusion reguiobs in the WIMP mass-cross section diagram



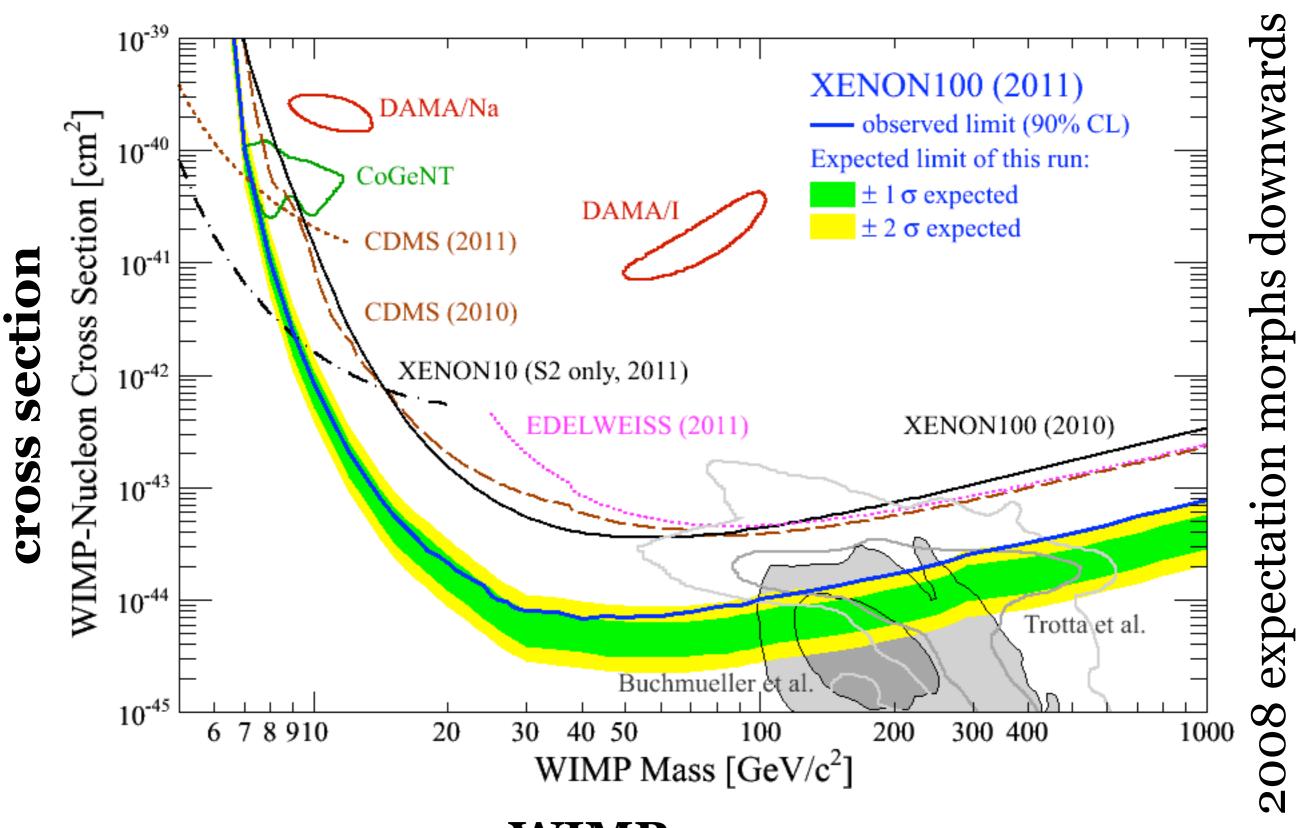
DATA listed top to bottom on plot CDMS (Soudan) 2004 Blind 53 raw kg-days Ge ZEPLIN III (Dec 2008) result XENONIO 2007 (Net 136 kg-d) Ellis et al., Spin dep. sigma in CMSSM Trotta et al 2008, CMSSM Bayesian: 68% contour Trotta et al 2008, CMSSM Bayesian: 95% contour





2011 exclusion

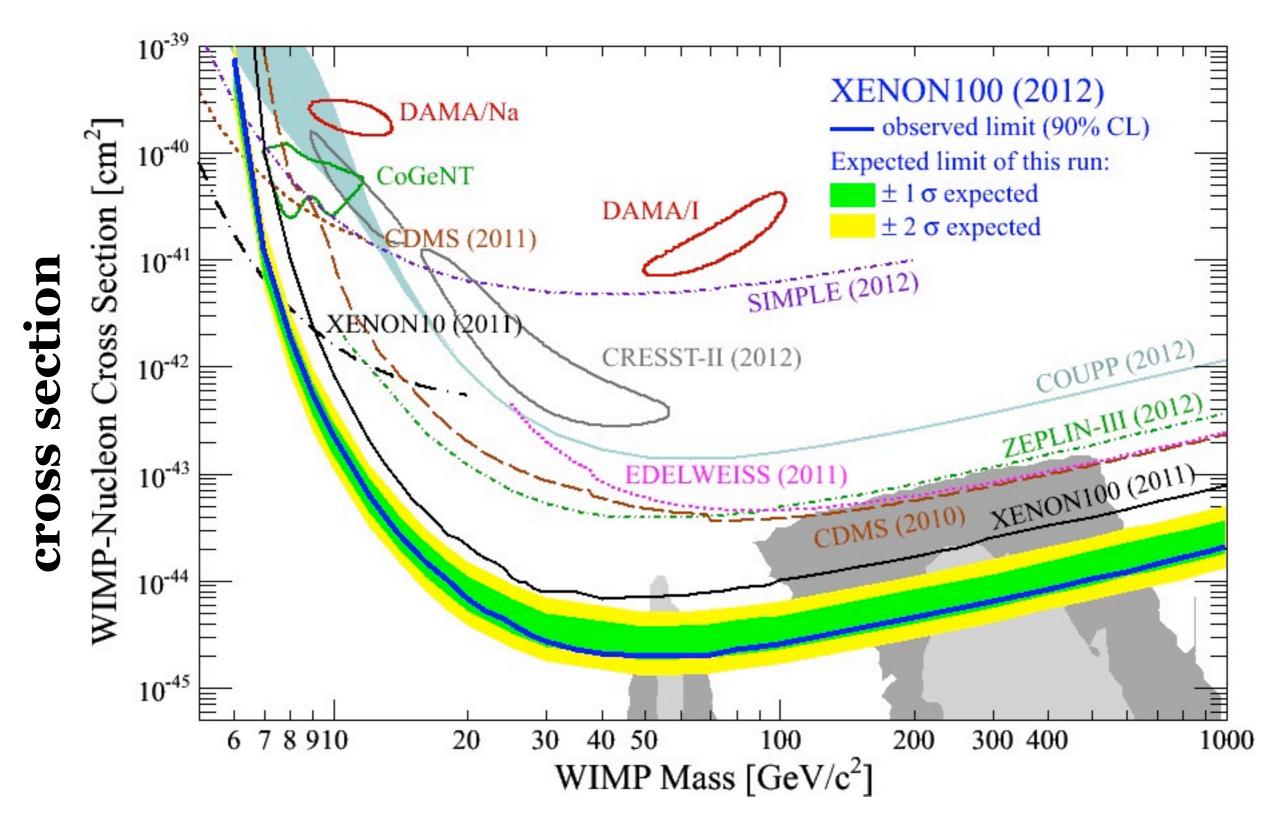
WIMP detection



WIMP mass

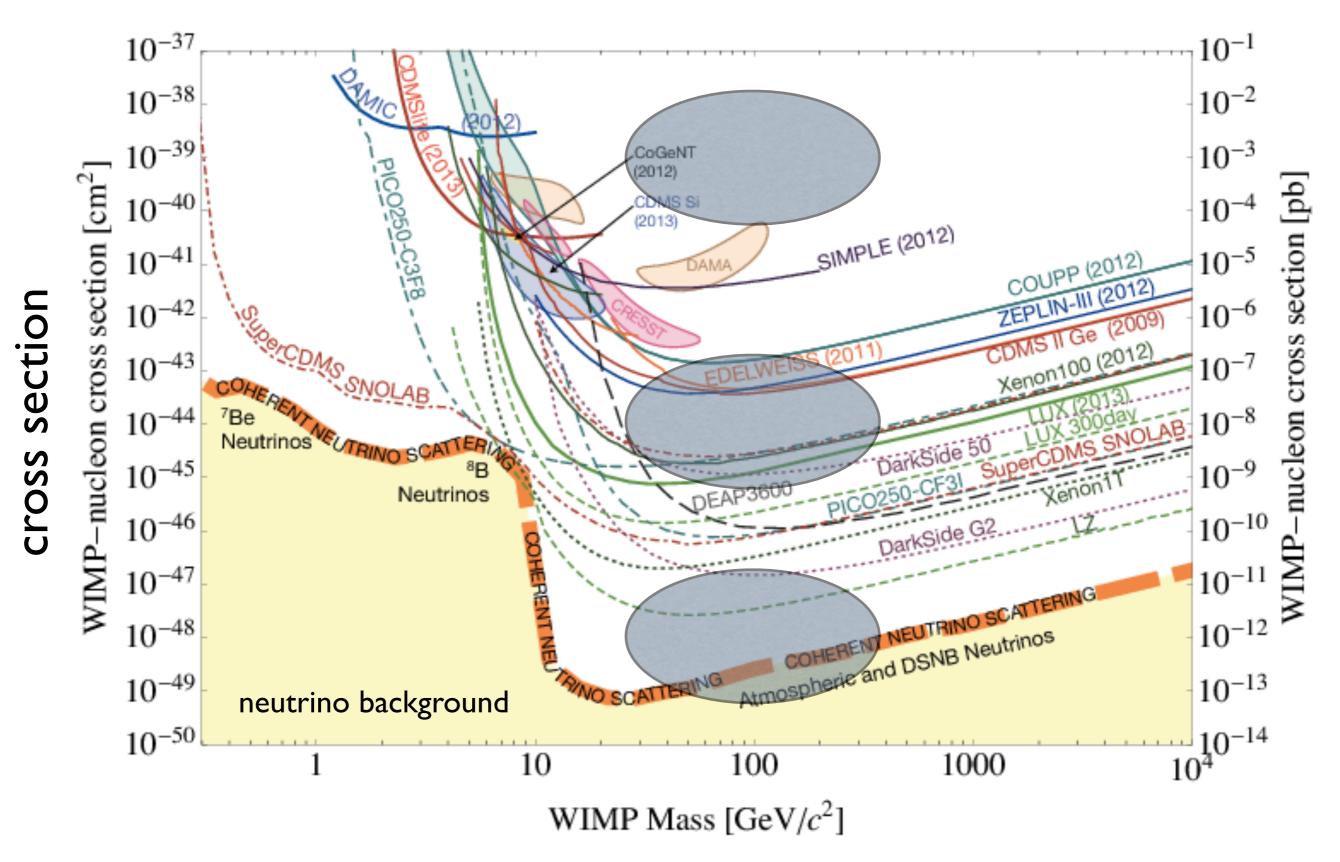
2012 exclusion

WIMP detection



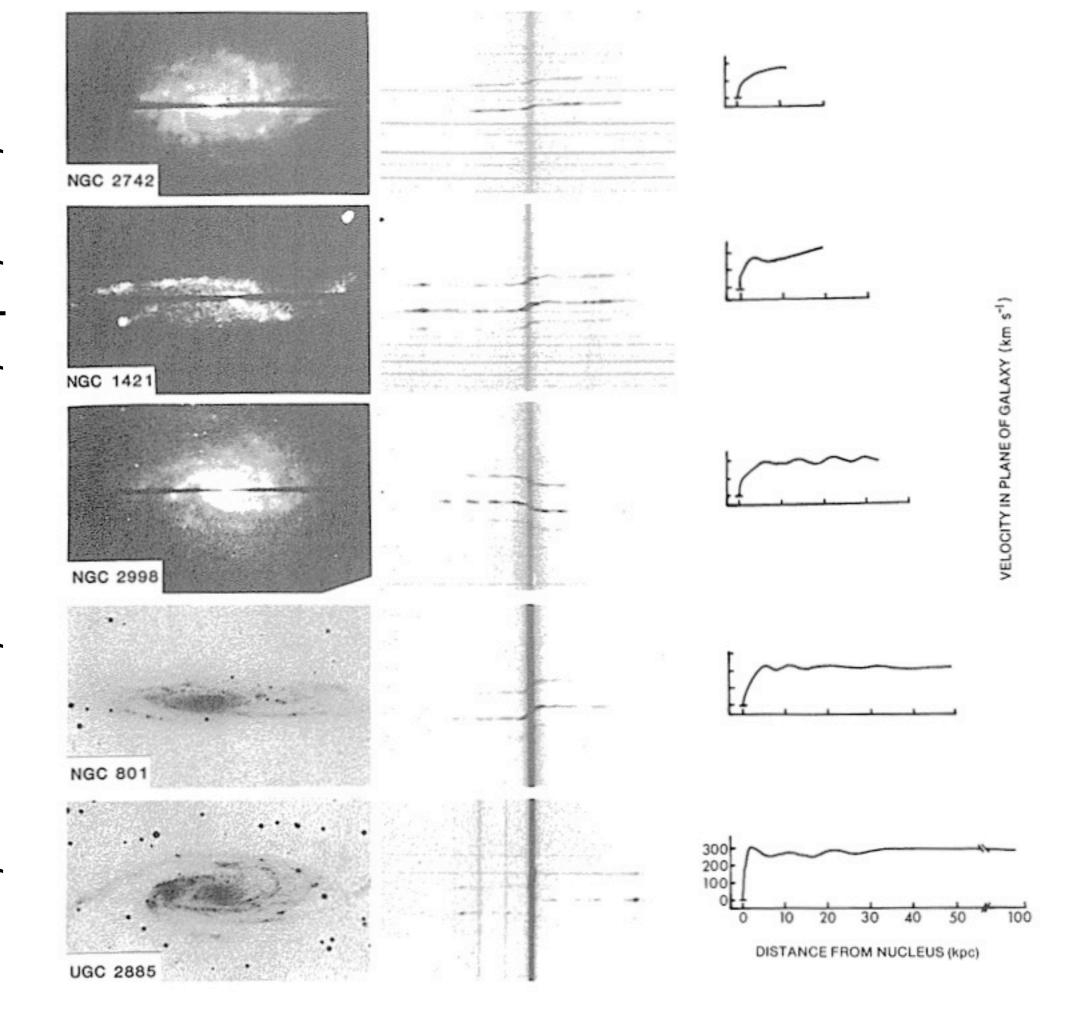
WIMP mass

cross section: 10⁻³⁹ 'natural'. Then 10⁻⁴⁴. Next: 10⁻⁴⁸



WIMP mass (~100 GeV natural)

A single galaxy might seem a little thing to those who consider only the immeasurable vastness of the universe, and not the minute precision to which all things therein are shaped.



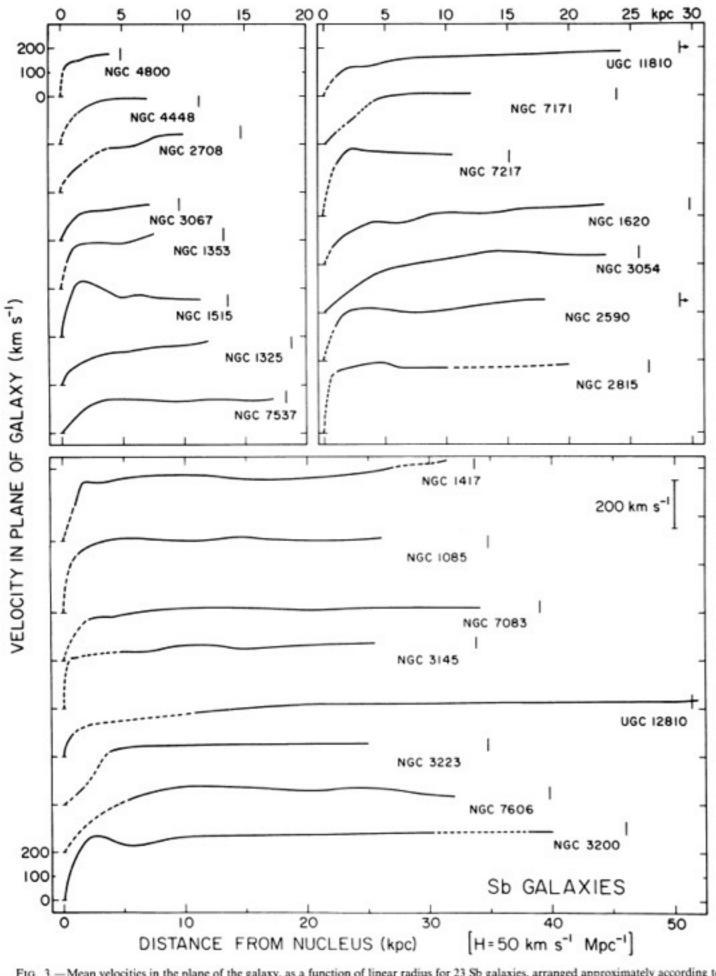


Fig. 3.—Mean velocities in the plane of the galaxy, as a function of linear radius for 23 Sb galaxies, arranged approximately according to increasing luminosity. Adopted curve is rotation curve formed from the mean of velocities on both sides of the major axis. Vertical bar marks the location of R₂₅, the isophote of 25 mag arcsec⁻², corrected for effects of internal extinction and inclination. Regions with no measured velocities are indicated by dashed lines.

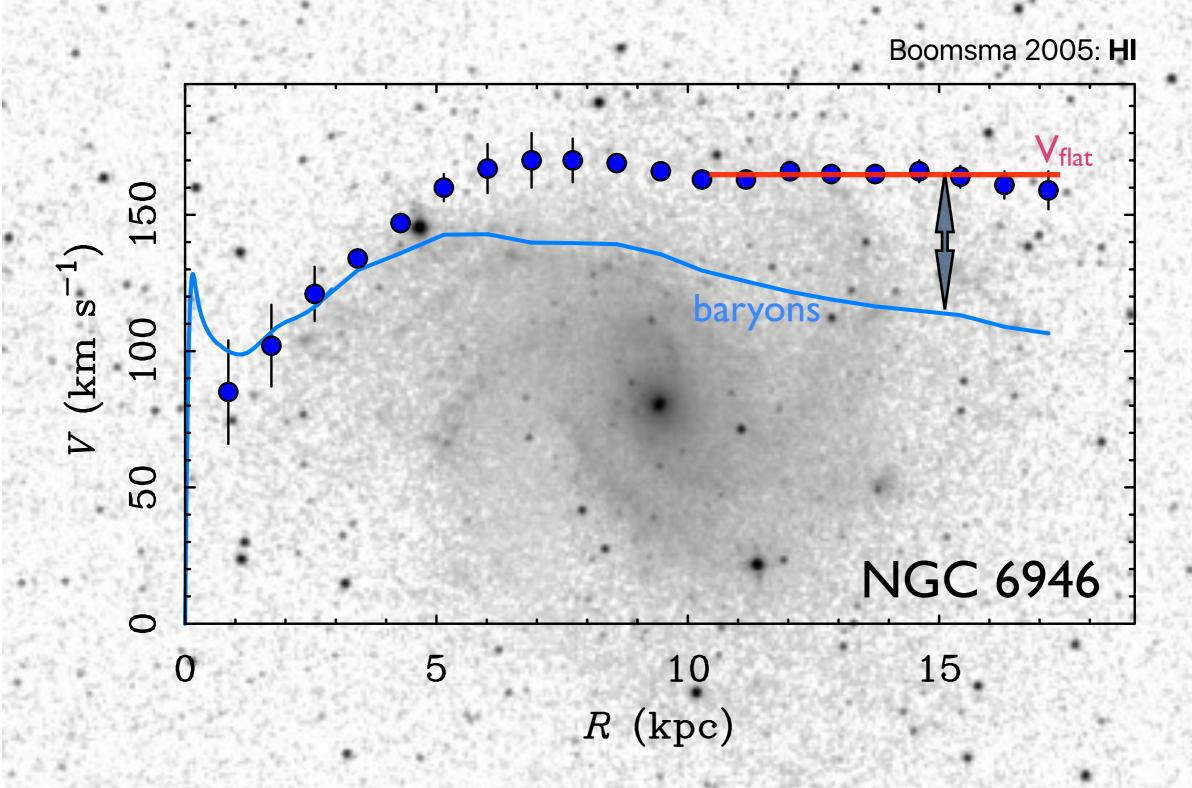
Rotation curves become flat at large radii

 $V \propto {\rm const}$

 $M \propto R$

 $\rho \propto R^{-2}$

Rubin, Thonnard, & Ford 1978, ApJ, 225, L107





Federico Lelli | Stacy McGaugh | James Schombert

SPARC is a sample of

- 175 disk galaxies with resolved observations of both stars and gas
- Spitzer [3.6] photometry tracing the stellar mass
- 21 cm maps of the atomic (HI) gas
- high-quality HI/Hlpha rotation curves tracing the gravitational potential

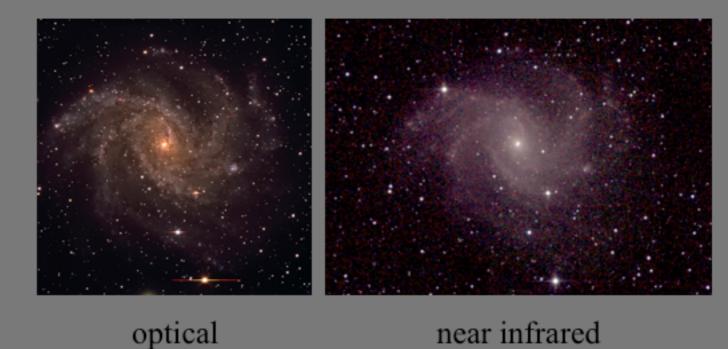
SPARC spans a wide range in

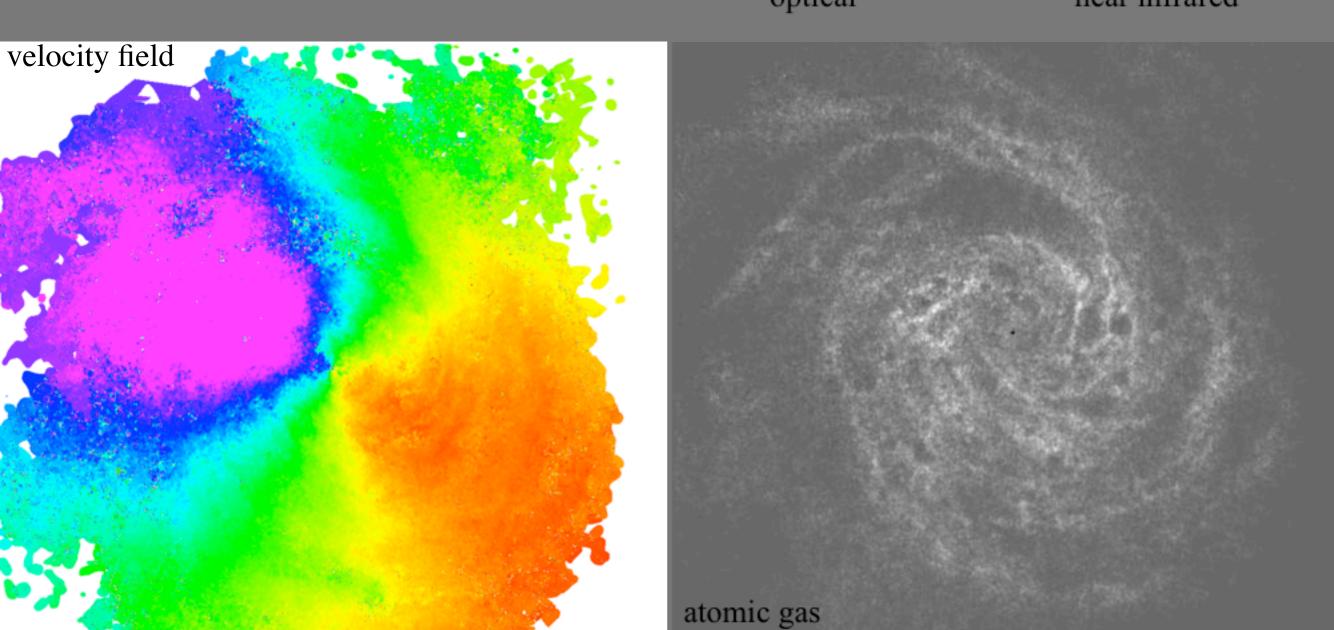
- morphological types (S0 to Irr basically everything that rotates)
- stellar masses (5 dex) $10^7 < M_* < 10^{12} \ \mathrm{M_{\odot}}$
- surface brightnesses (4 dex) $5 < \Sigma_0 < 10^4 \ {\rm M_{\odot} \ pc^{-2}}$
- all gas fractions $0.01 < f_g < 0.99$

NGC 6946

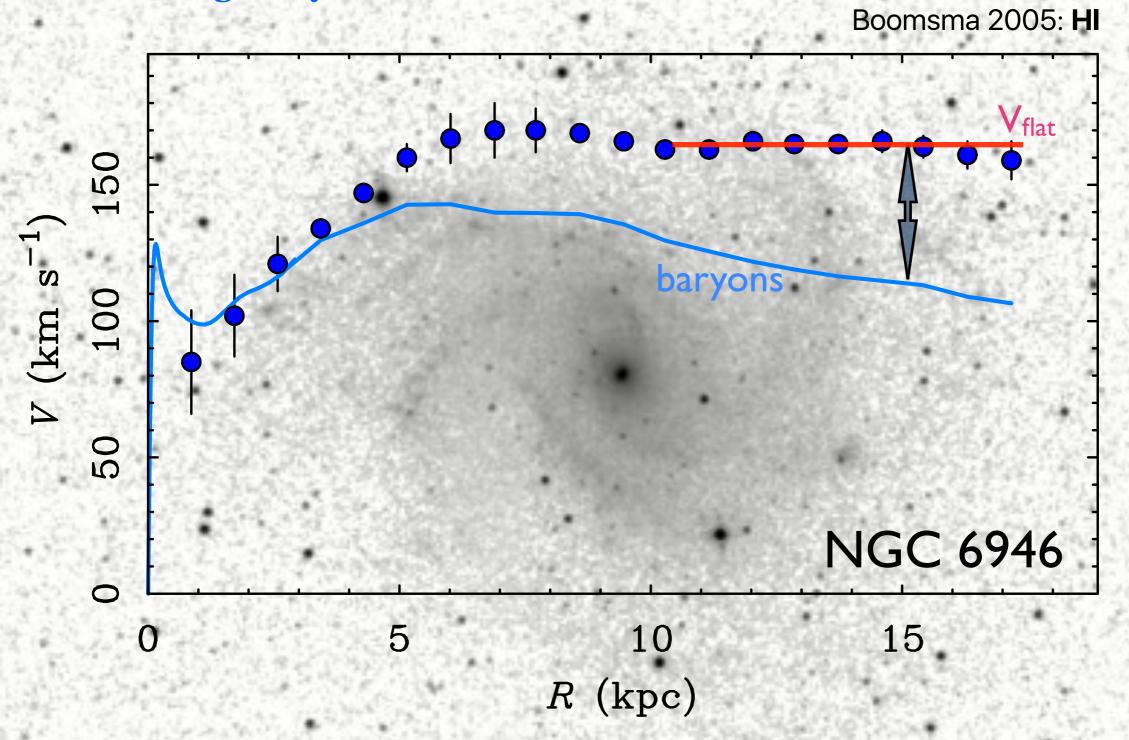
For each galaxy, have at least

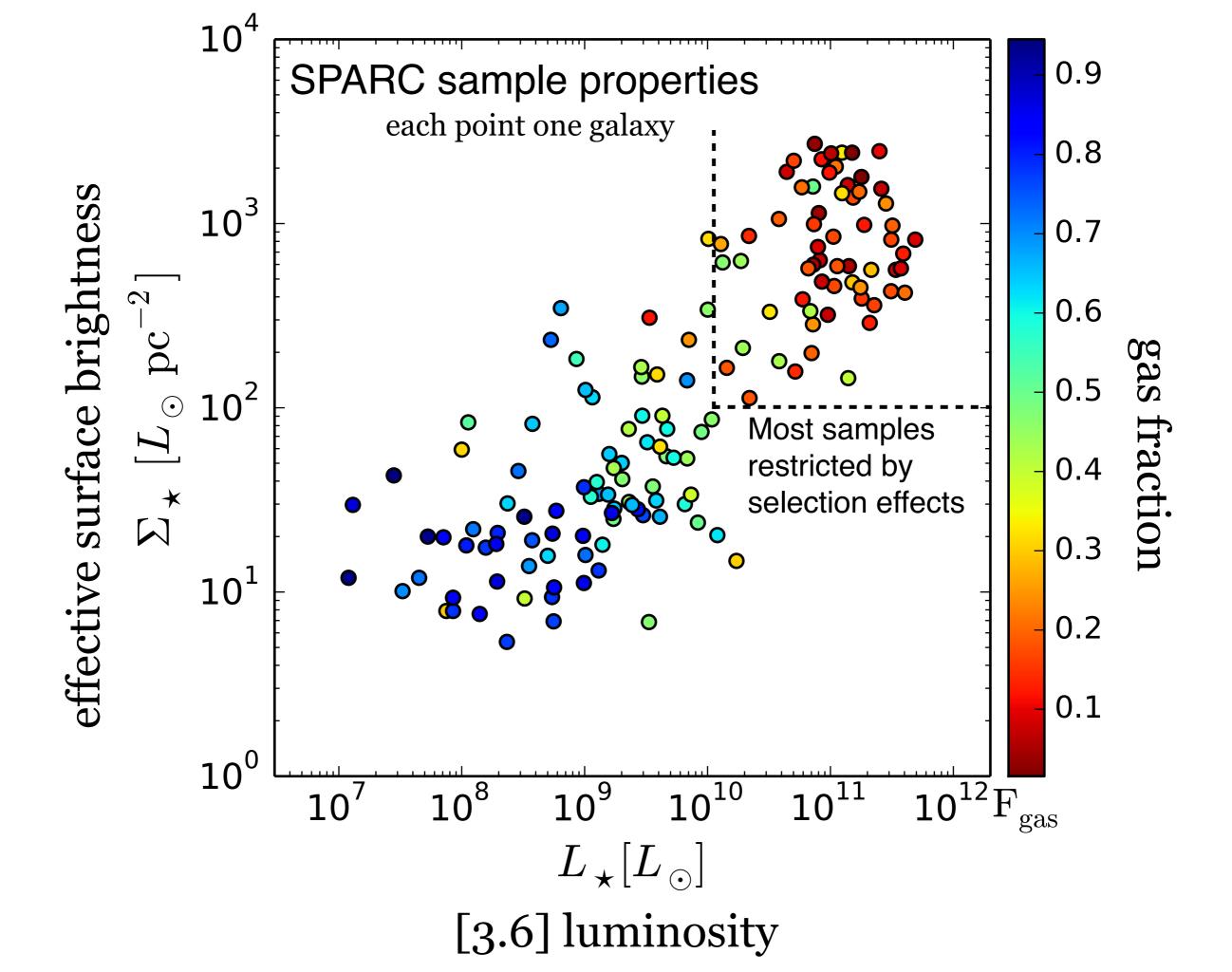
- near-IR [3.6] image (stars)
- HI map (atomic gas)
- HI velocity field





Excess rotation indicates new physics: dark matter or modified gravity



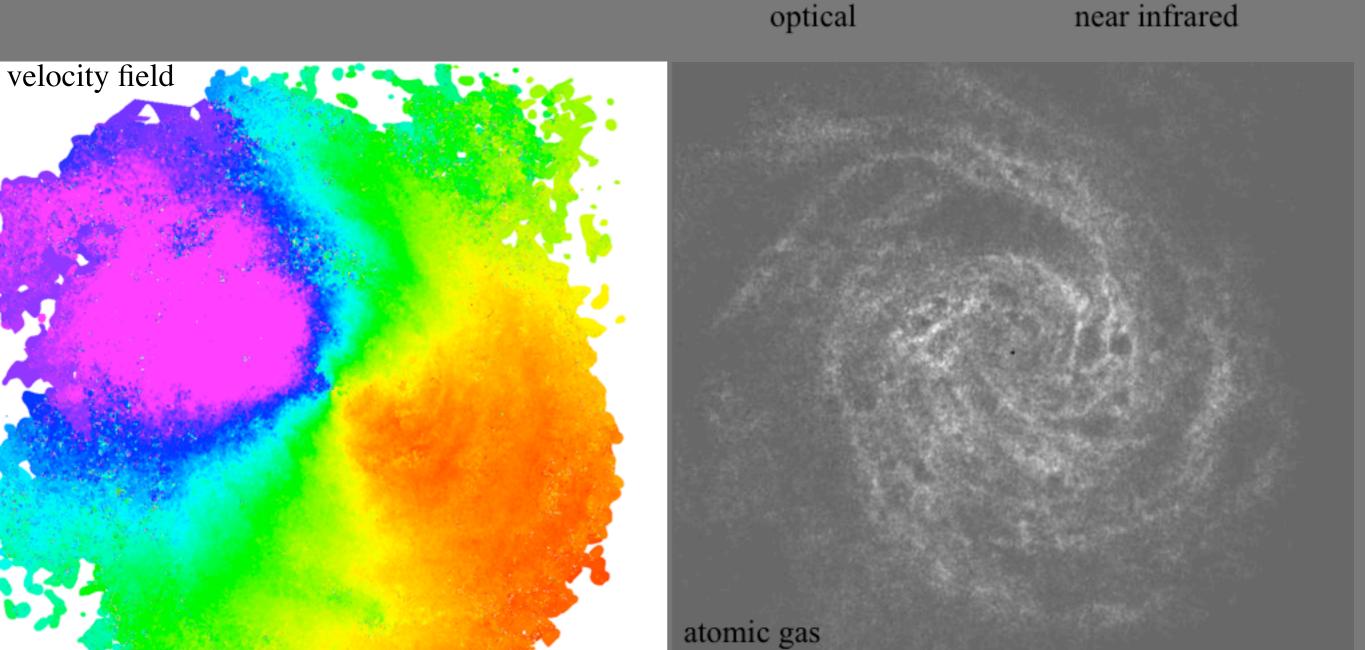


NGC 6946

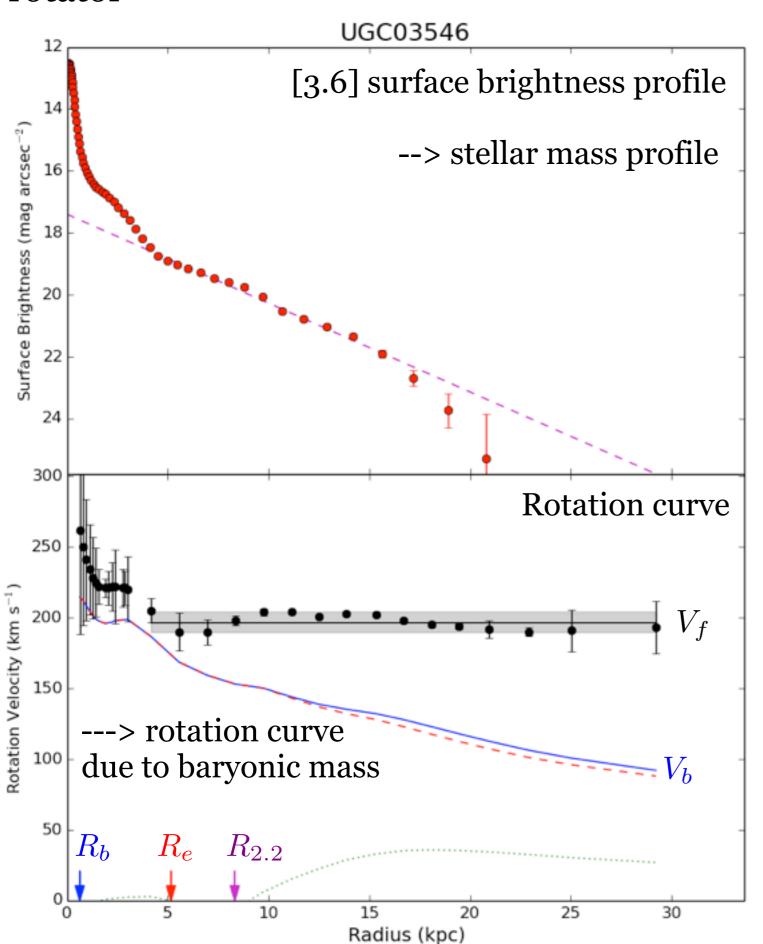
For each galaxy, construct mass models

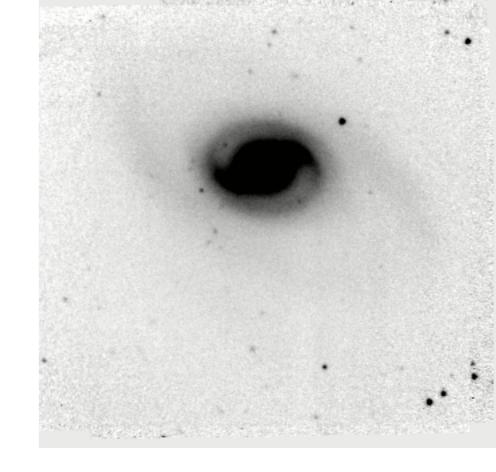
- [3.6] image -> stellar mass
- HI map -> gas mass velocity field -> rotation curve





Fast rotator

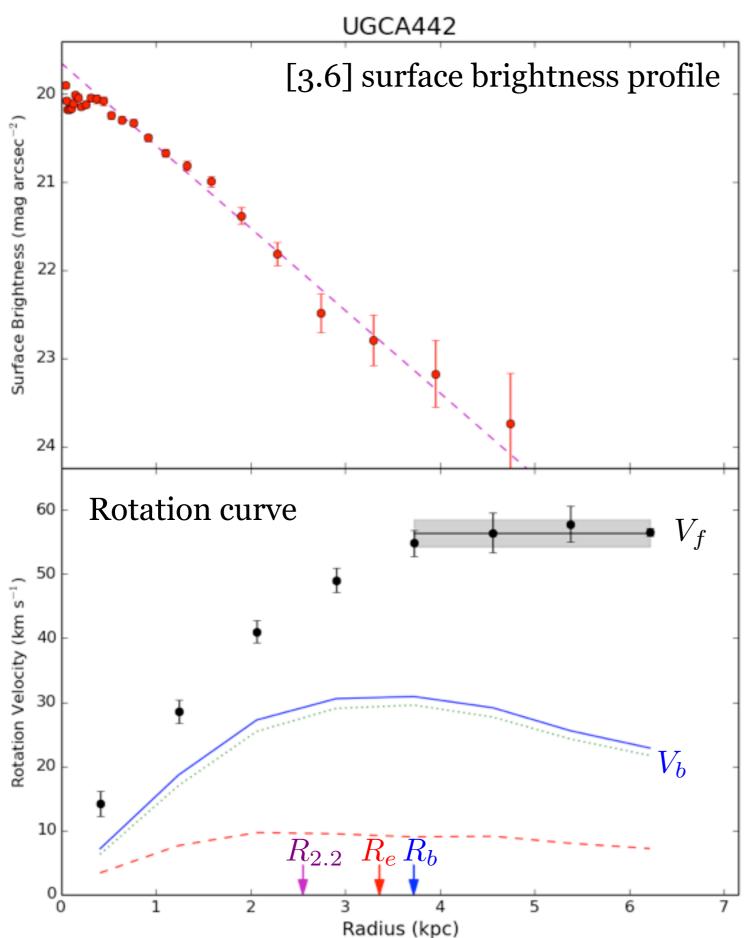


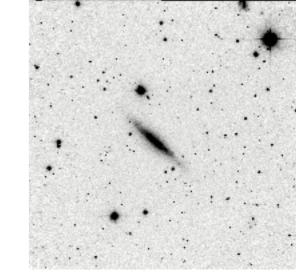


Observed mass scales with rotation velocity: Tully-Fisher Relation

Fast rotators are high mass

Slow rotator

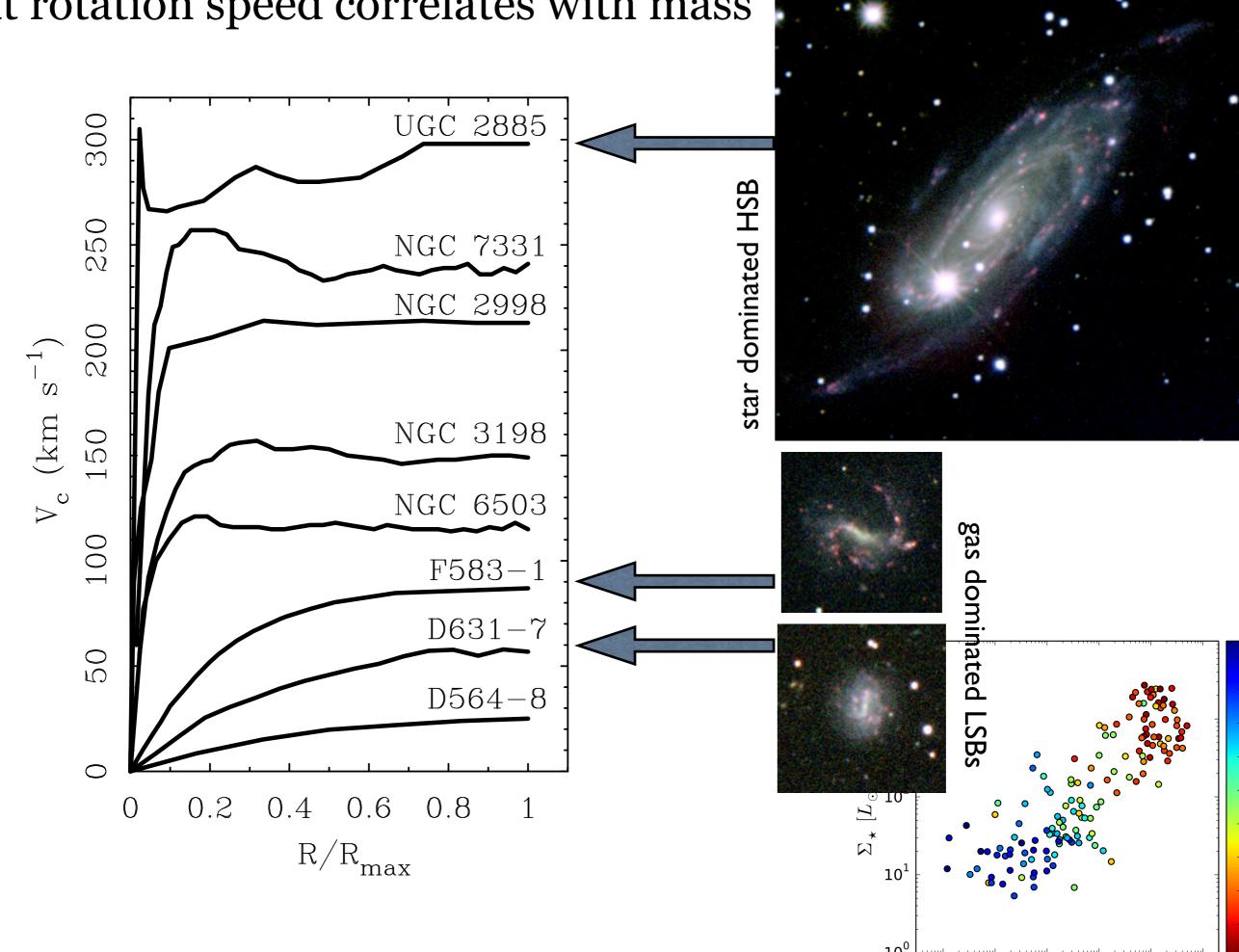




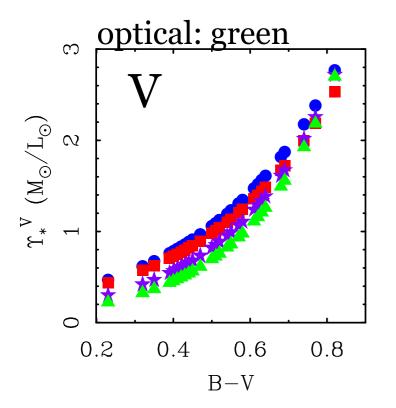
Observed mass scales with rotation velocity: Tully-Fisher Relation

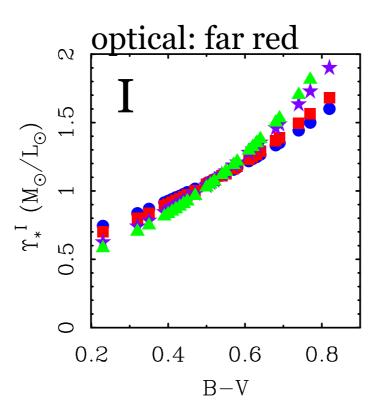
Slow rotators are low mass

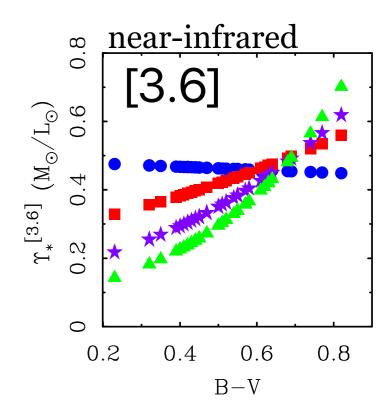
Flat rotation speed correlates with mass



Stellar Mass







Stellar population synthesis models modified for self-consistency Band dependent mass-to-light ratios correlate with color

Table 7 Revised CMLR

Model	a_V	b_V	α_I	β_I	$lpha_{[3.6]}$	$eta_{[3.6]}$	$\Upsilon^V_{0.6}$	$\Upsilon^I_{0.6}$	$\Upsilon^K_{0.6}$	$\Upsilon_{0.6}^{[3.6]}$
Bell et al. (2003) Portinari et al. (2004) Zibetti et al. (2009) Into & Portinari (2013)	-0.654 -1.075	$1.290 \\ 1.837$	$-0.321 \\ -0.477$	$0.701 \\ 1.004$	$-0.594 \\ -1.147$	$0.467 \\ 1.289$	$1.32 \\ 1.07$	$\frac{1.26}{1.33}$	$0.63 \\ 0.54$	$0.49 \\ 0.42$

Note. — Stellar mass-to-light ratios in the V, I, and K-bands given by the formula $\log \Upsilon^j_* = \alpha_j + \beta_j (B - V)$. For each model, the V-band is identical to that in Table 2, but the I and [3.6] bands have been revised to attain self-consistency with the V-band, and further corrected for V - I as a second color term: the Δ of Table 6 have been incorporated to produce these revised CMLR. The resulting lines are fits to the data in Fig. 8, providing our best estimate of the CMLR. For reference, the mass-to-light ratio at B - V = 0.6 is also given. For the K_s -band, we assume $\Upsilon^K_* = 1.29 \Upsilon^{[3.6]}_*$.

Situation at [3.6] improved by proper treatment of metallicity distribution

Net result:
to a good approximation,
M*/L constant in the
near-IR

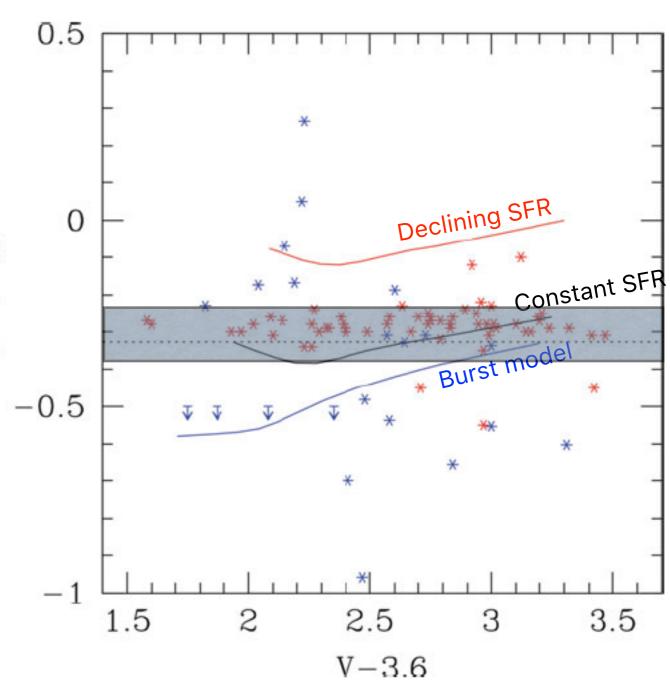
For disk component

$$\Upsilon_*^{[3.6]} = 0.5 \ \mathrm{M_{\odot}/L_{\odot}}$$
 to +/- 20%

For bulge component

$$\Upsilon_*^{[3.6]} = 0.7~{\rm M}_{\odot}/{\rm L}_{\odot}$$
 only 31 galaxies have bulges

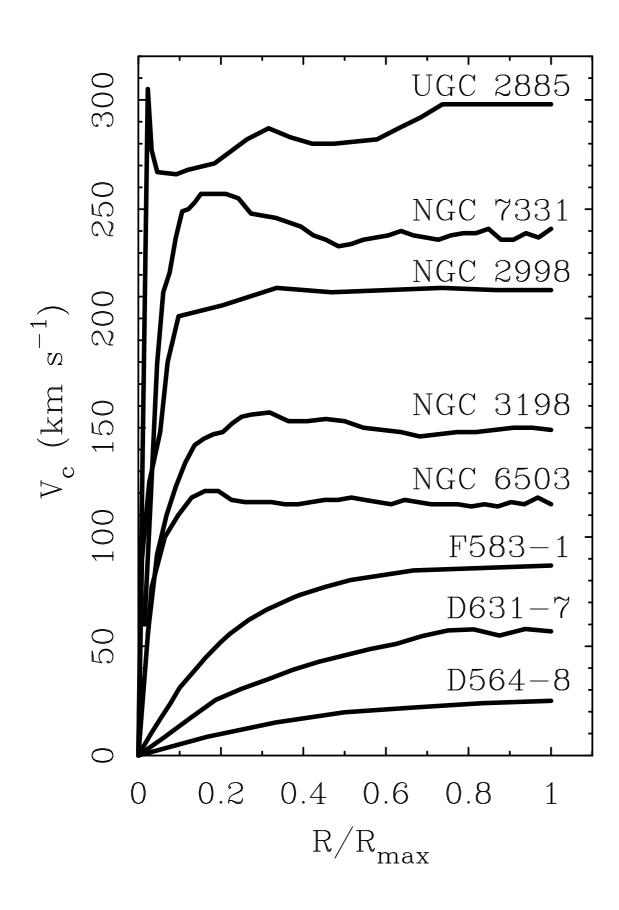
The metallicity distribution matters as well as the age distribution



Models with realistic metallicity distributions predict flatter CMLR in near-IR

There should always be some scatter in the CMLR

Flat rotation speed correlates with mass



Measures of mass

Stellar mass

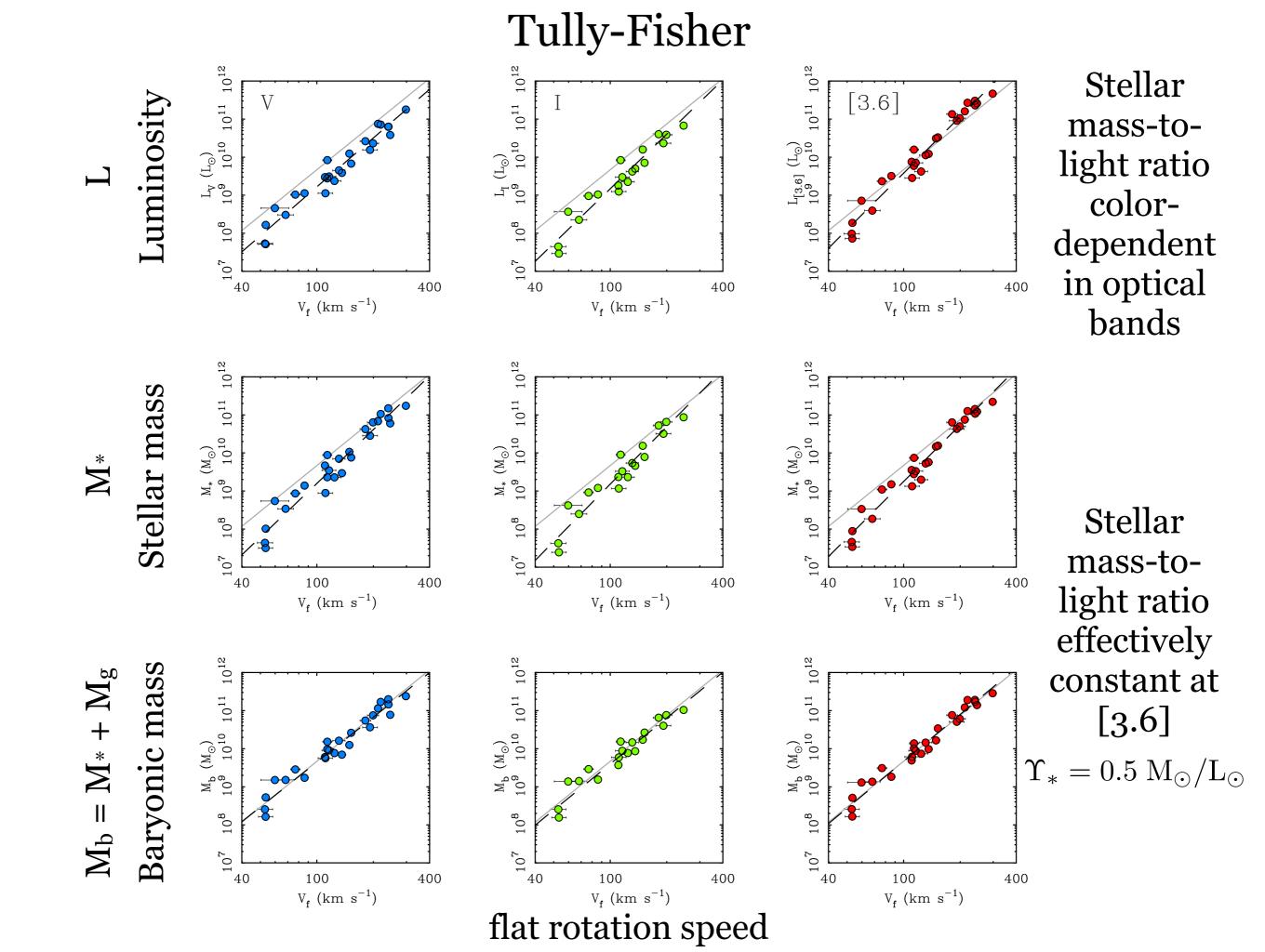
$$\Upsilon^i_*=lpha_i+eta_i(B-V)$$
 $M_*=\Upsilon^i_*L_i$ e.g., $\Upsilon^{[3.6]}_*=0.5~{
m M}_\odot/{
m L}_\odot$

Gas mass

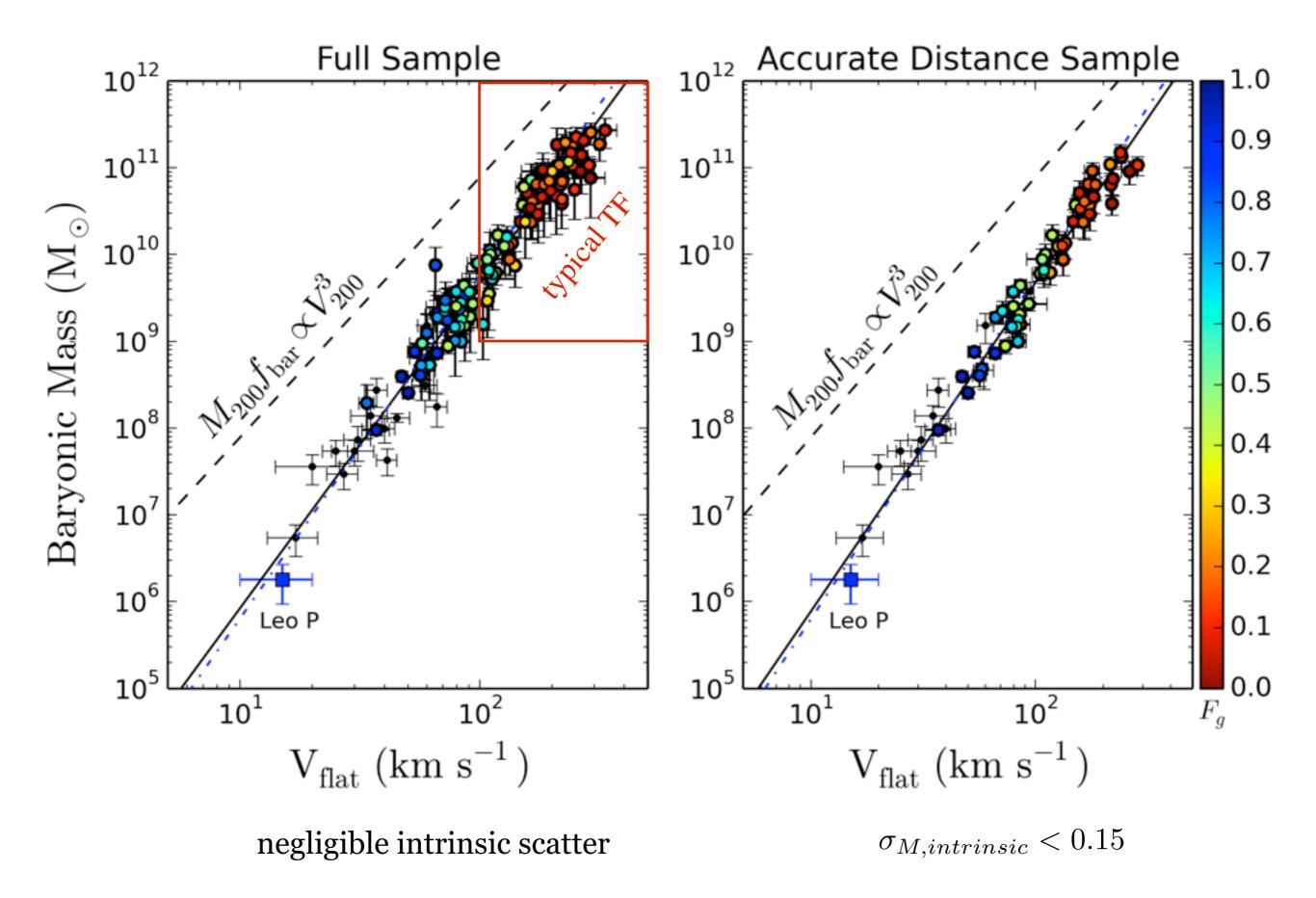
Direct from 21 cm flux corrected for helium abundance

Baryonic mass

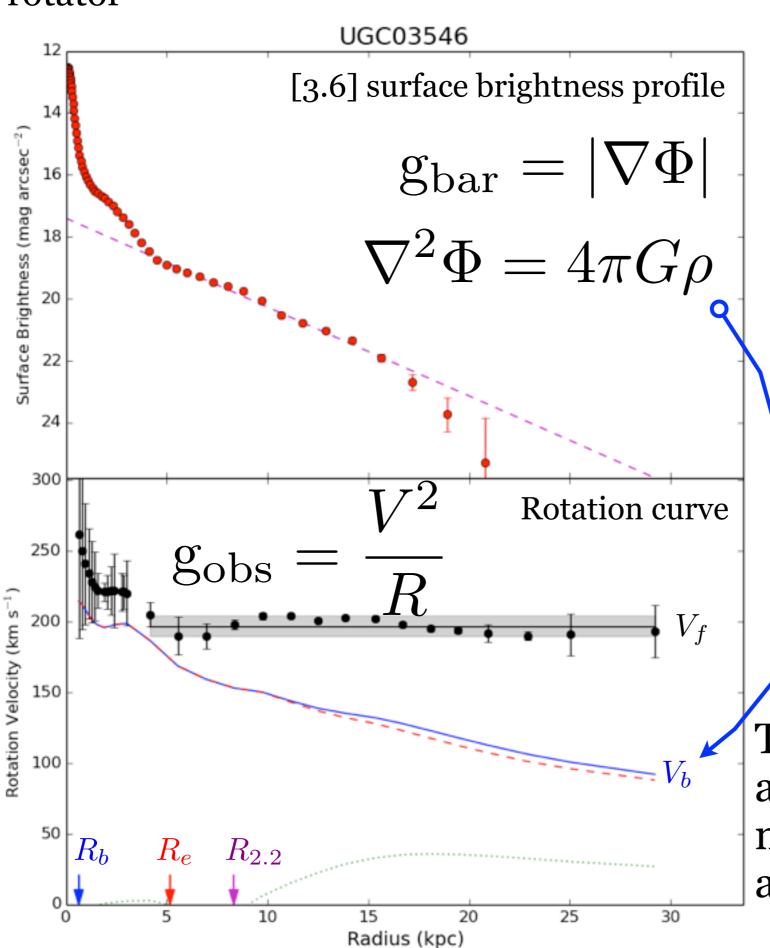
$$M_b = M_* + M_g$$

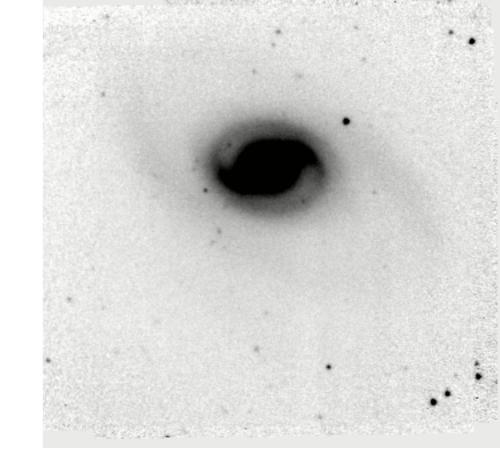


Baryonic Tully-Fisher relation: $M_b = 47 \text{ V}^4$



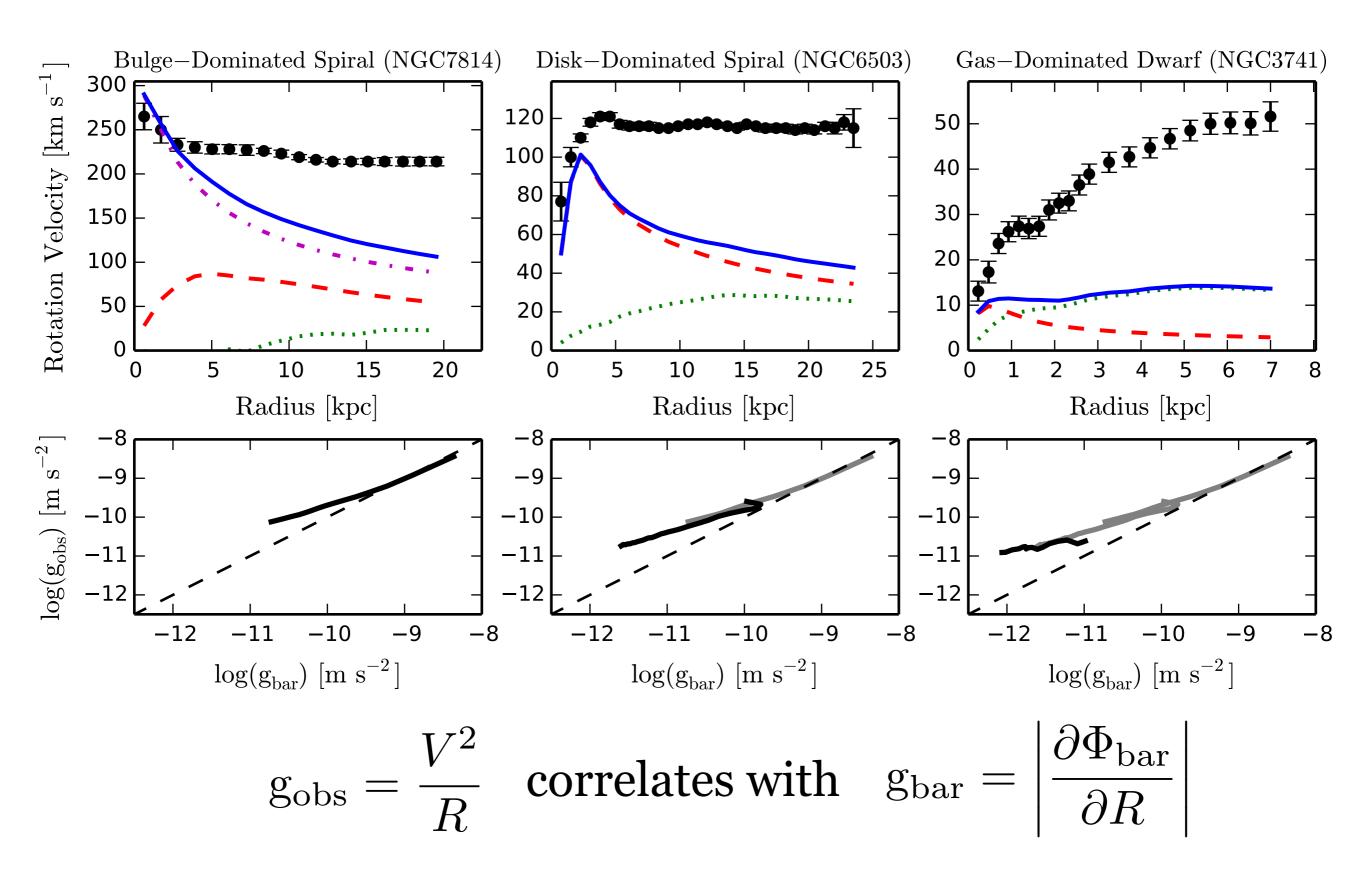
Fast rotator





The Tully-Fisher Relation is a global coupling between mass and light. There is also a local relation.

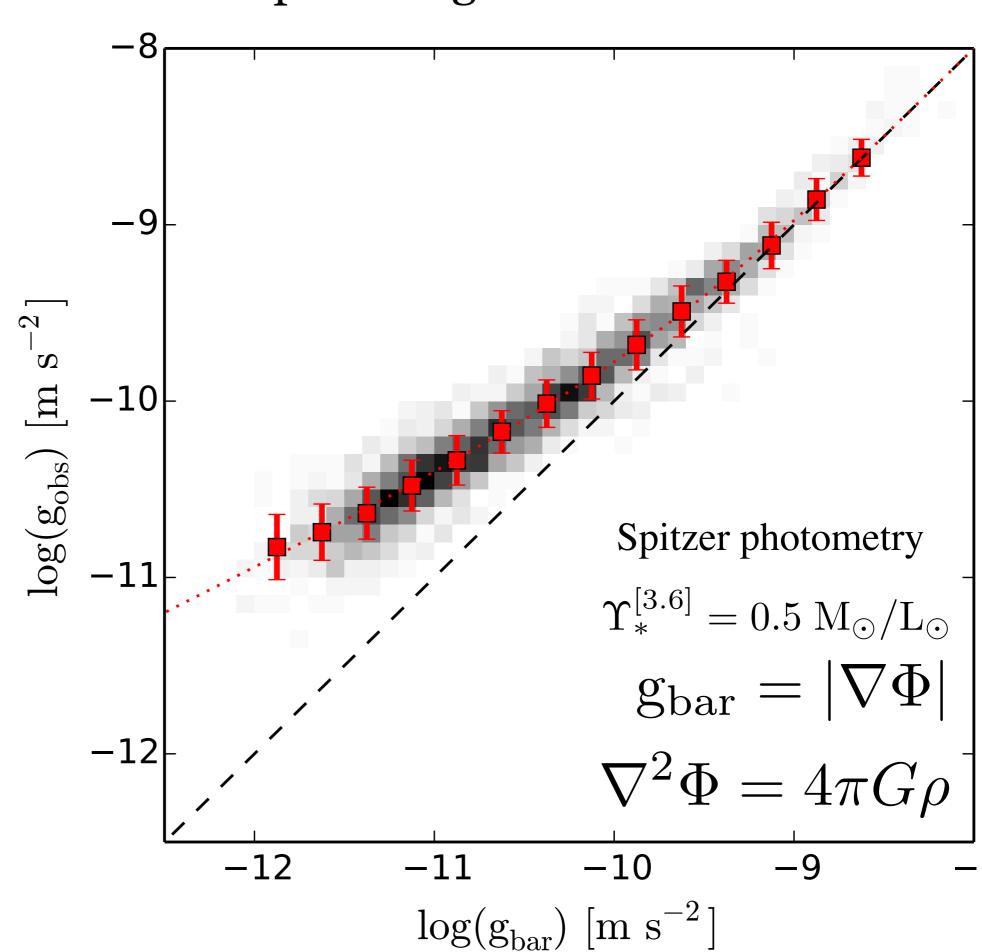
Use the same [3.6] mass-to-light ratio for all galaxies



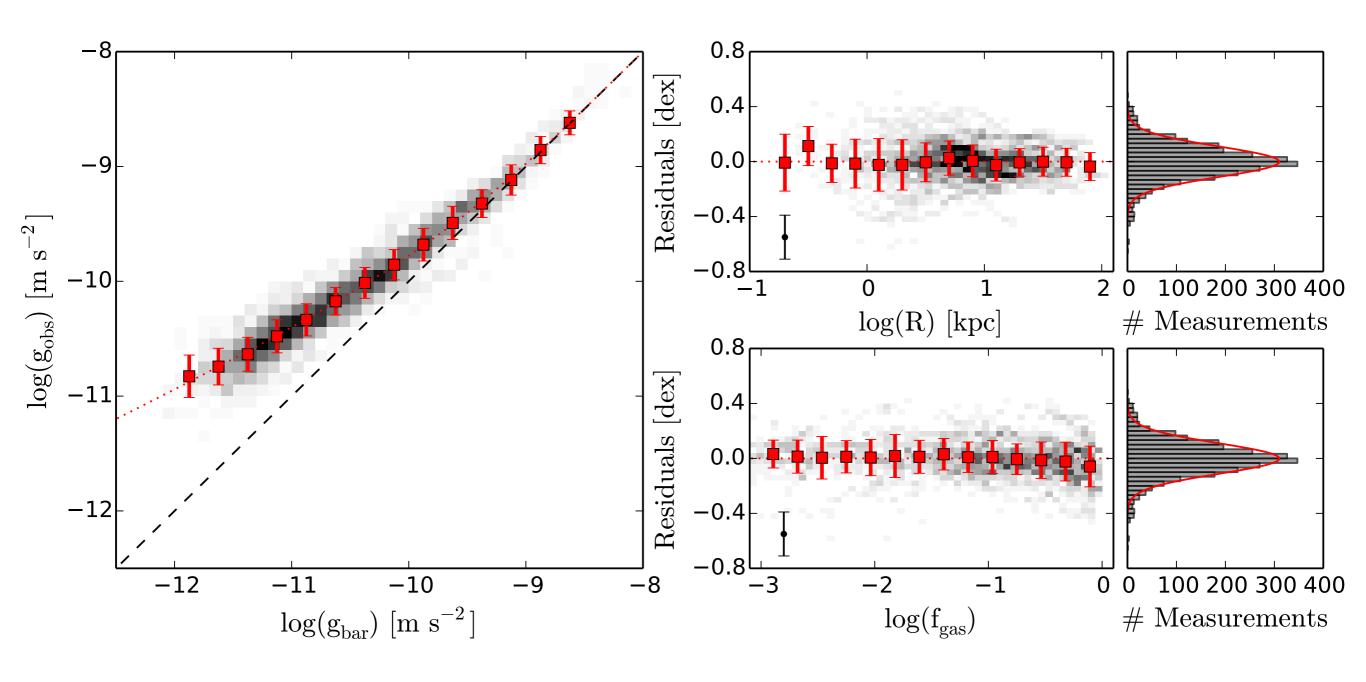
Mass couples to light

Rotation curves
~ 2700 points in
153 galaxies
(disturbed gals &
i < 30° excluded)

$$g_{\rm obs} = \frac{V^2}{R}$$



Mass couples to light



$$g_{\text{obs}} = \frac{g_{\text{bar}}}{1 - e^{-\sqrt{\frac{g_{\text{bar}}}{g_{\dagger}}}}}$$

$$g_{\dagger} = 1.2 \times 10^{-10} \text{ m s}^{-2}$$

 ± 0.02 (random) ± 0.24 (systematic)

3 Laws of Galactic Rotation

- 1. Rotation curves tend towards asymptotic flatness $V_f \rightarrow {\rm constant}$
- 2. Baryonic mass scales as the fourth power of rotation velocity $M_b \propto V_f^4$ (Baryonic Tully-Fisher Relation)

Just the facts, mam.
Just the facts.



3. Gravitational force couples to baryon distribution

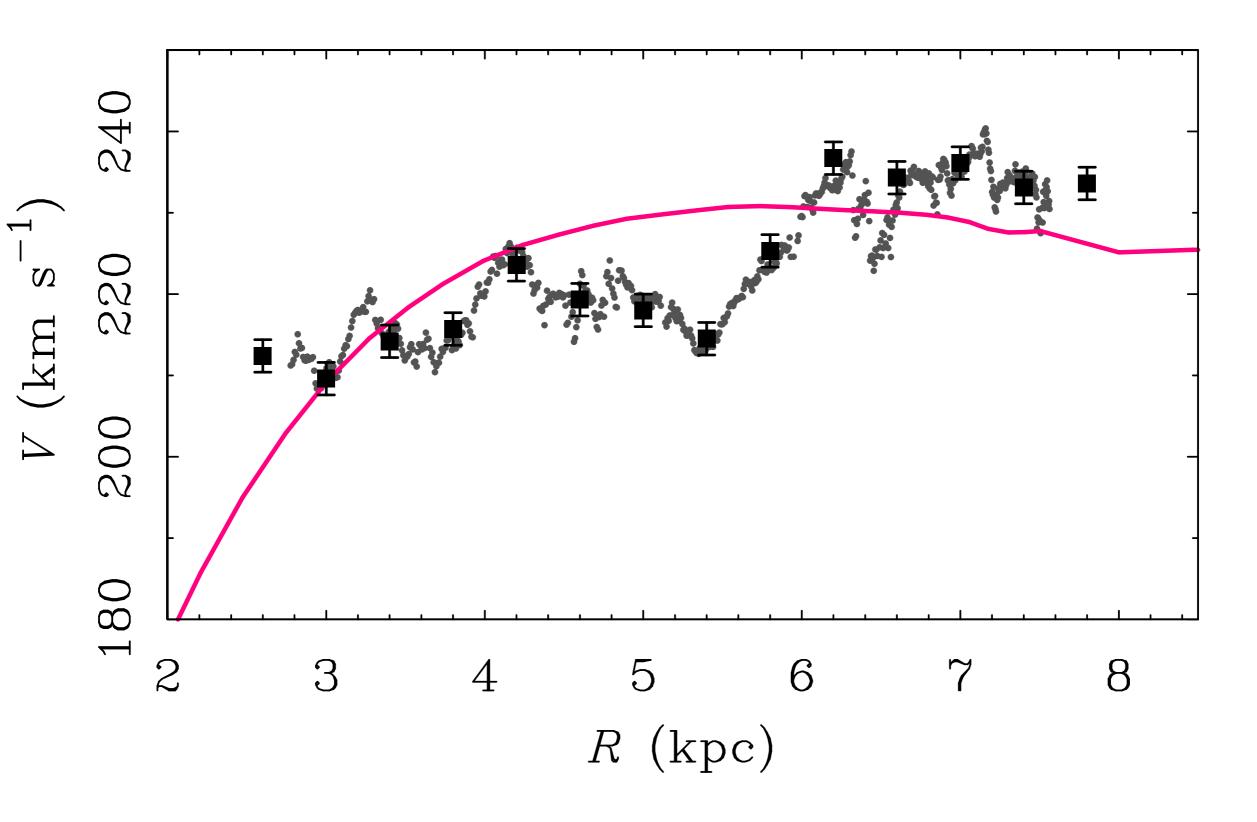
$$g_{\rm bar} = \left| \frac{\partial \Phi_{\rm bar}}{\partial R} \right|$$

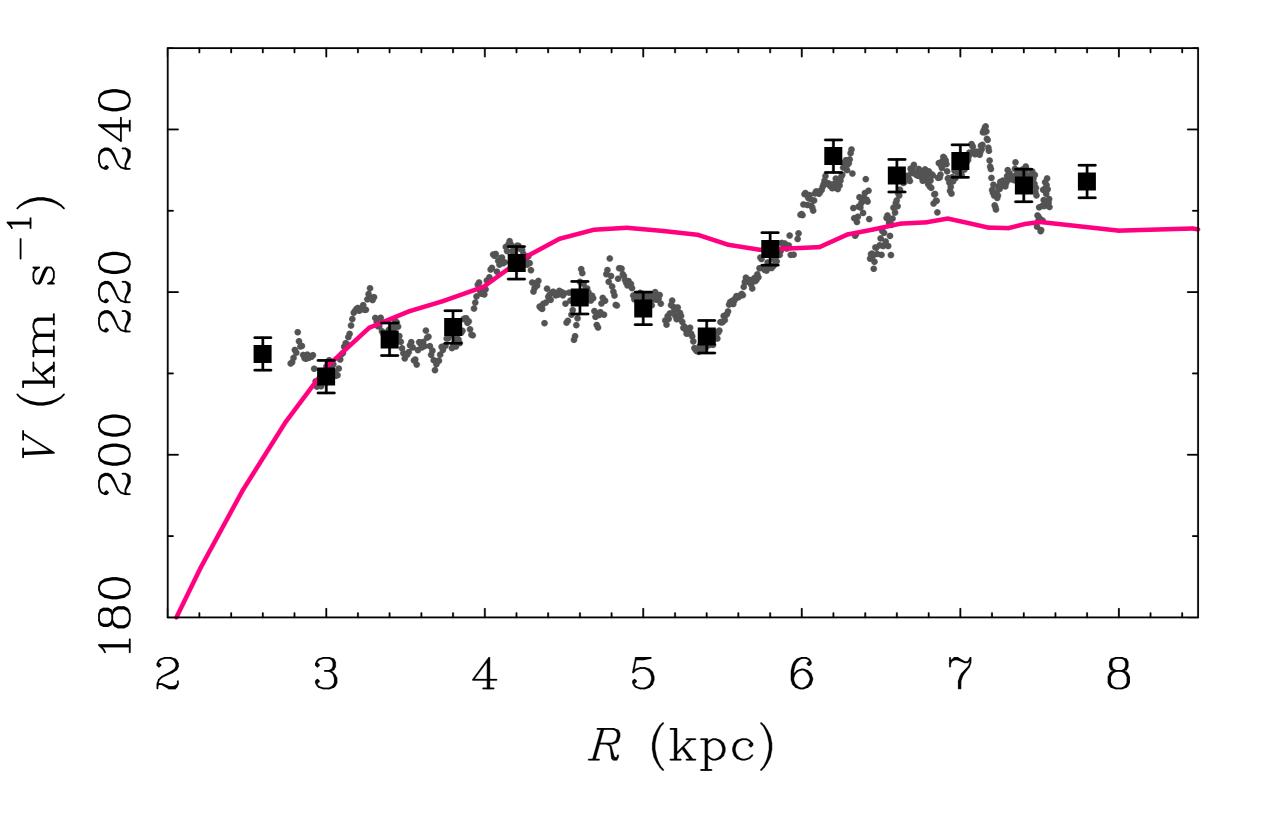
predictive of *total* gravitational potential

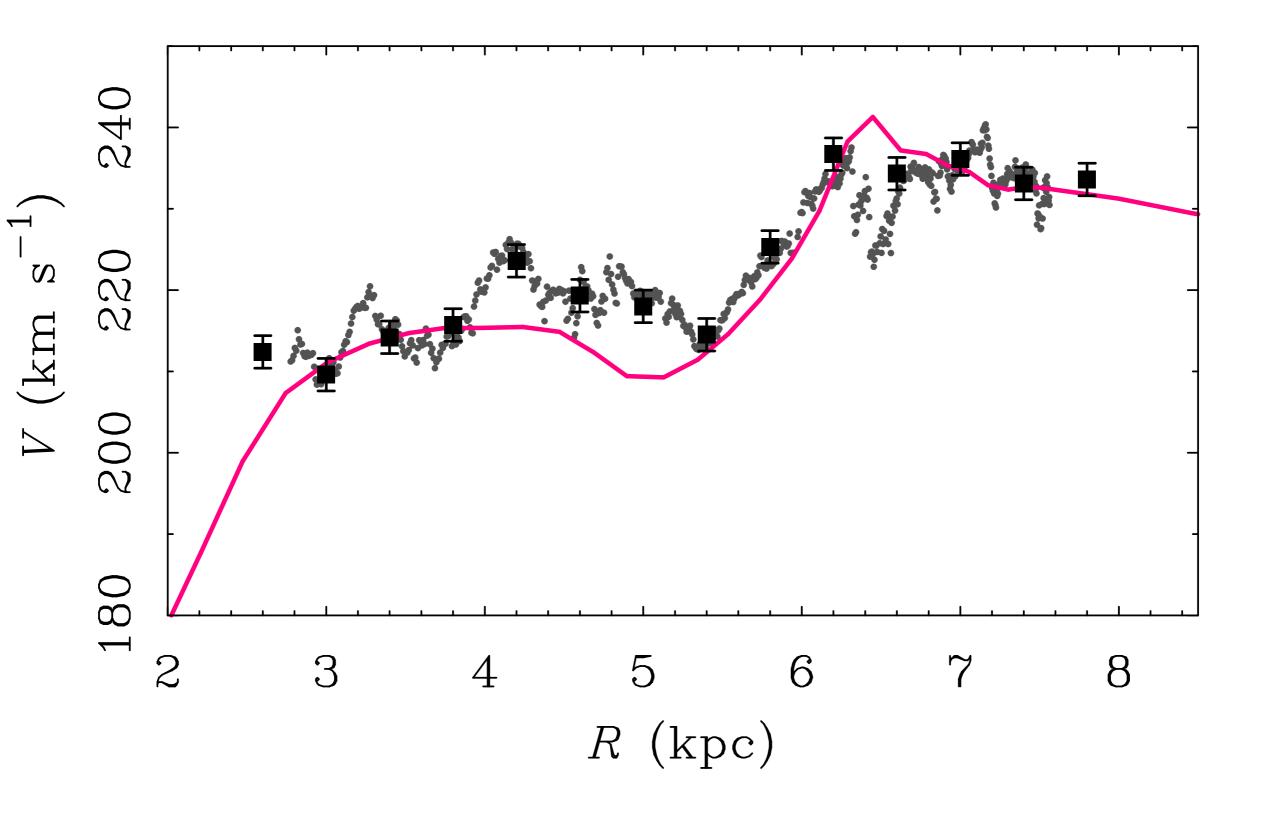
Example application: our own Galaxy

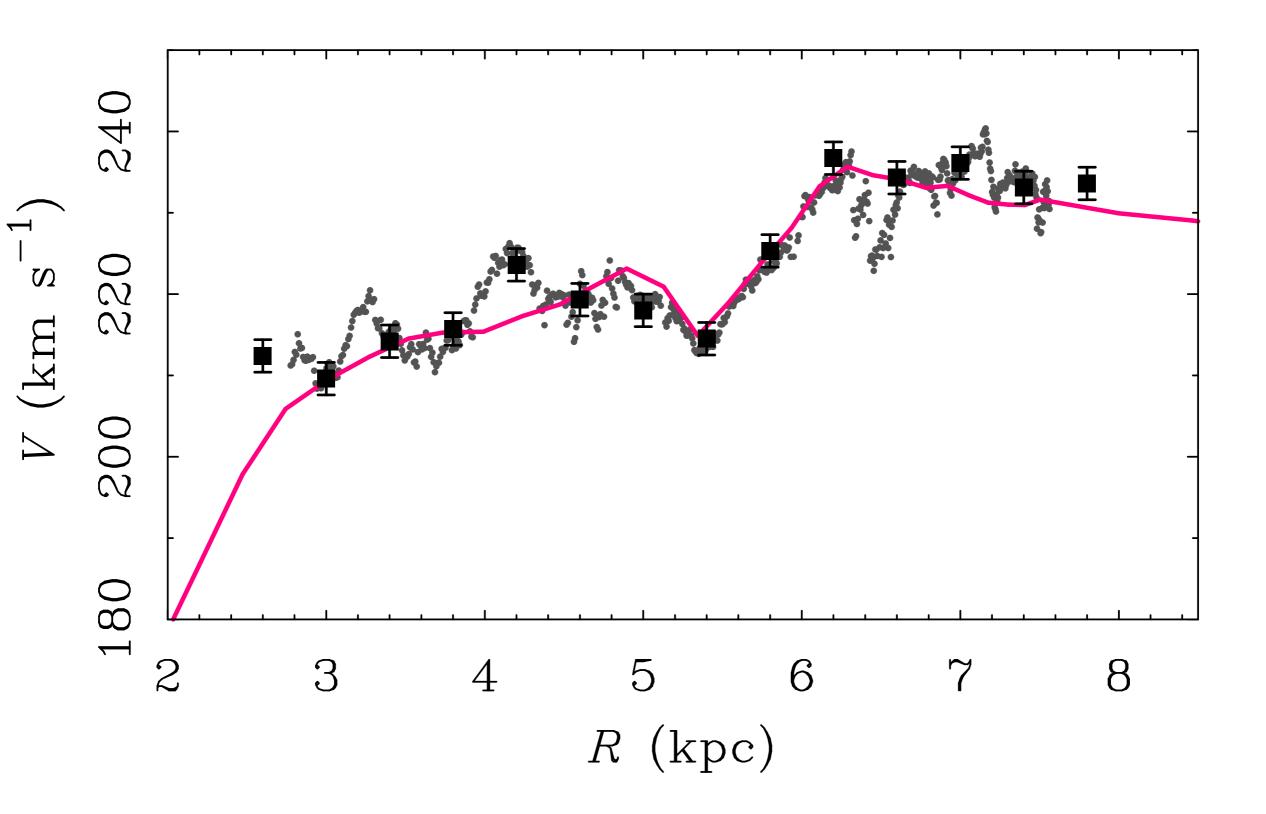


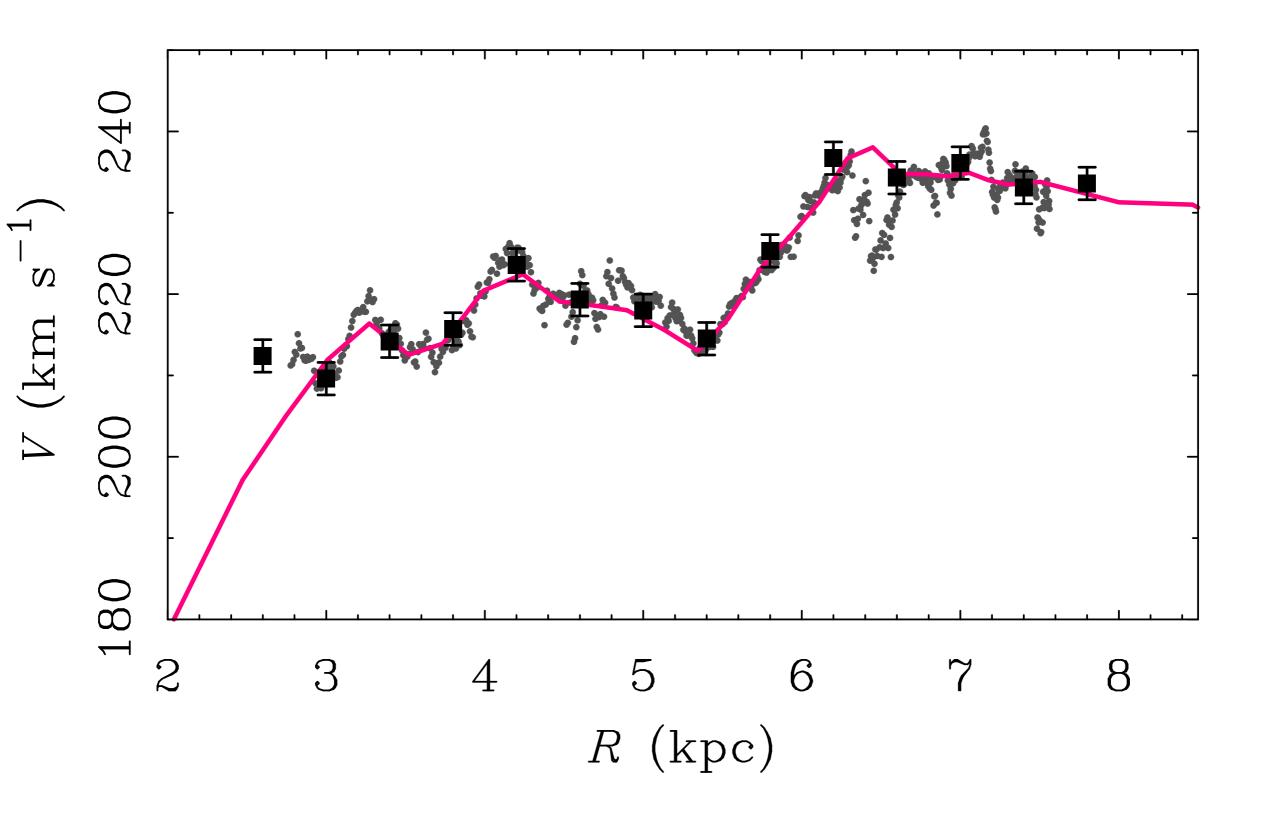
Fitting the details of the terminal velocity curve: infer the stellar mass distribution from V(R)



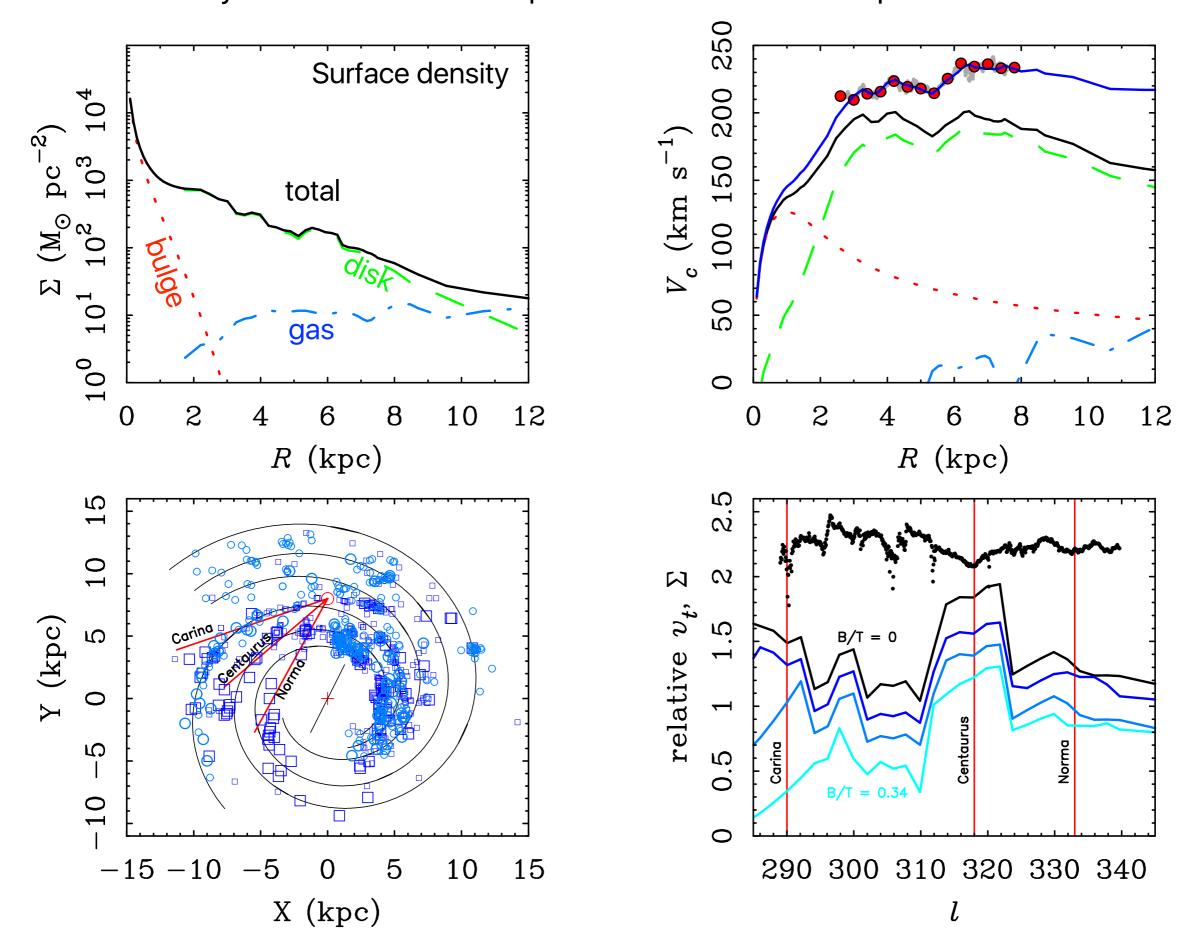


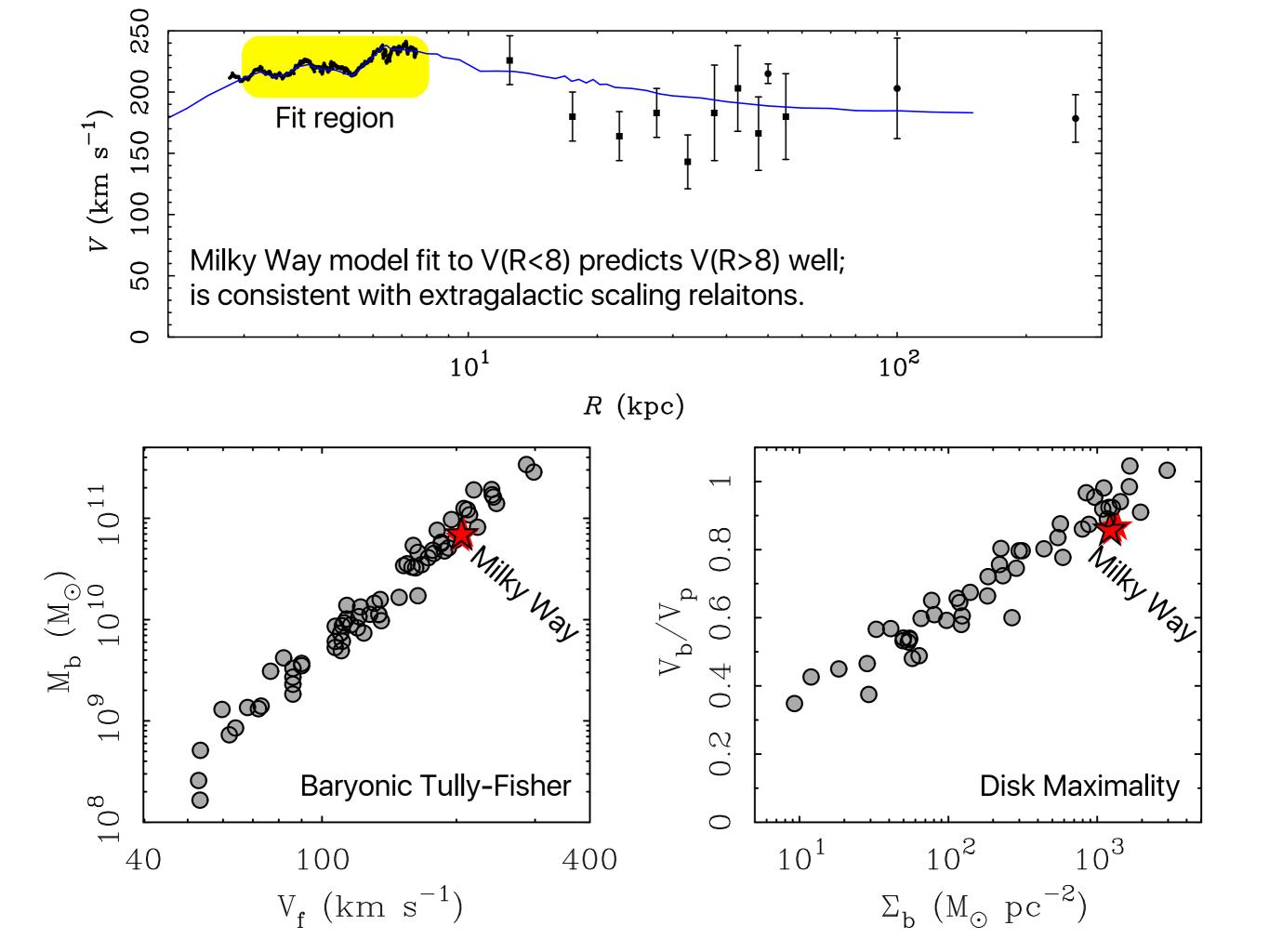






Fitting the details of the terminal velocity curve uncovers details of Milky Way structure: the inferred density enhancement corresponds to the Centaurus spiral arm.





3 Laws of Galactic Rotation

- 1. Rotation curves tend towards asymptotic flatness $V_f \to {\rm constant}$
- 2. Baryonic mass scales as the fourth power of $M_b \propto V_f^4$ rotation velocity (Baryonic Tully-Fisher Relation)
- 3. Gravitational force couples to baryon distribution

$$\left| rac{\partial \Phi_{\mathrm{bar}}}{\partial \varphi_{\mathrm{obs}}}
ight|
ightarrow g_{\mathrm{obs}}$$

These are empirical results that must be explained by a successful theory.

Requires considerable fine-tuning to interpret in terms of dark matter.

Follows more naturally from a universal force law.

$$\left| \frac{\partial \Phi_{\text{bar}}}{\partial R} \right| \rightarrow g_{\text{obs}} = \left| \frac{\partial \Phi_{\text{tot}}}{\partial R} \right|$$



MOND

Modify gravity at an acceleration scale

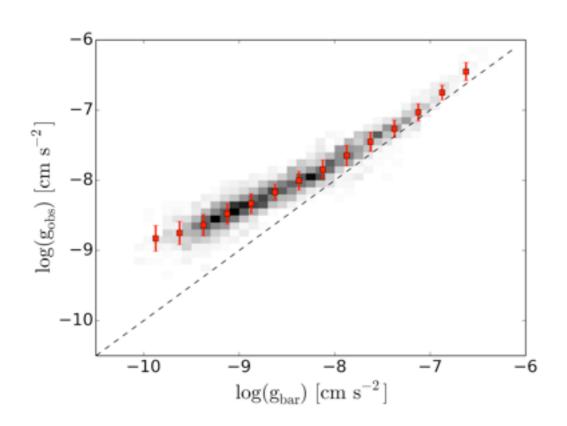
$$a_0 \approx 10^{-10} \text{ m s}^{-2} \sim cH_0 \sim c\Lambda^{1/2}$$

$$a \gg a_0$$
 $a \to g_N$

$$a \ll a_0$$
 $a \to \sqrt{g_N a_o}$

Hypothesized by Milgrom (1983)

The 3 Laws are a natural consequence of MOND, which already has a long string of successful predictions

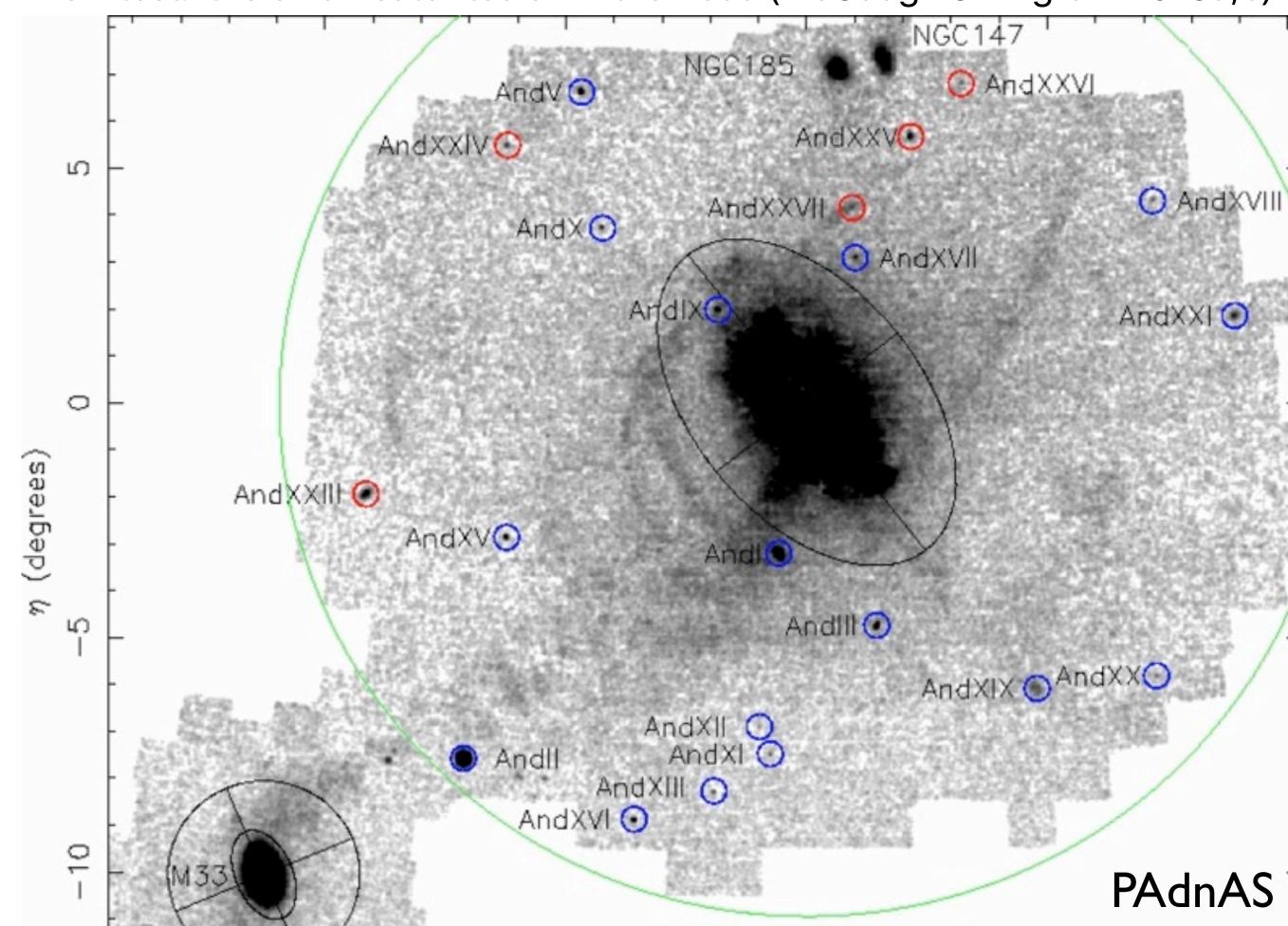


MOND predictions

The Tully-Fisher Relation

- Normalization = $1/(a_0G)$
- Fundamentally a relation between Disk Mass and V_{flat}
- No Dependence on Surface Brightness
- Dependence of conventional M/L on radius and surface brightness
- Rotation Curve Shapes
- Surface Density ~ Surface Brightness
- Detailed Rotation Curve Fits
- Stellar Population Mass-to-Light Ratios

A new test: the dwarf satellites of Andromeda (McGaugh & Milgrom 2013a,b)



Newtonian regime

$$g_{in} > a_0$$

$$M = \frac{RV^2}{G}$$



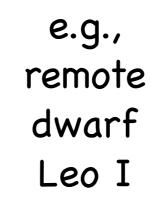
e.g., surface of the Earth

ISO

MOND regime

$$g_{in} < a_0$$

$$M = \frac{V^4}{a_0 G}$$





External Field dominant Newtonian regime

$$g_{in} < a_0 < g_{ex}$$



$$M = \frac{RV^2}{G}$$

e.g.,
Eotvos-type
experiment on
the surface of
the Earth

EFE

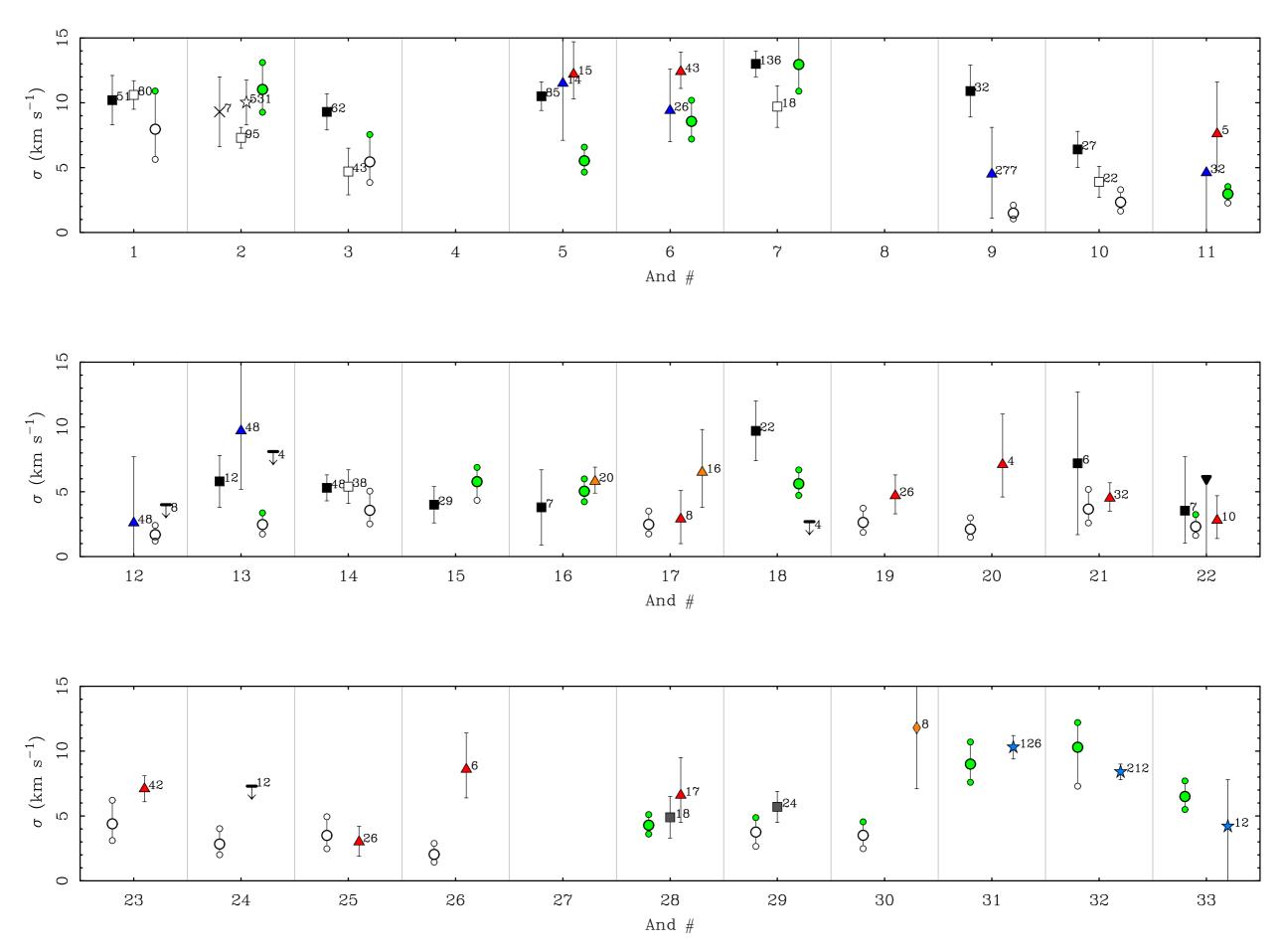
External Field dominant quasi-Newtonian regime

$$g_{in} < g_{ex} < a_0$$

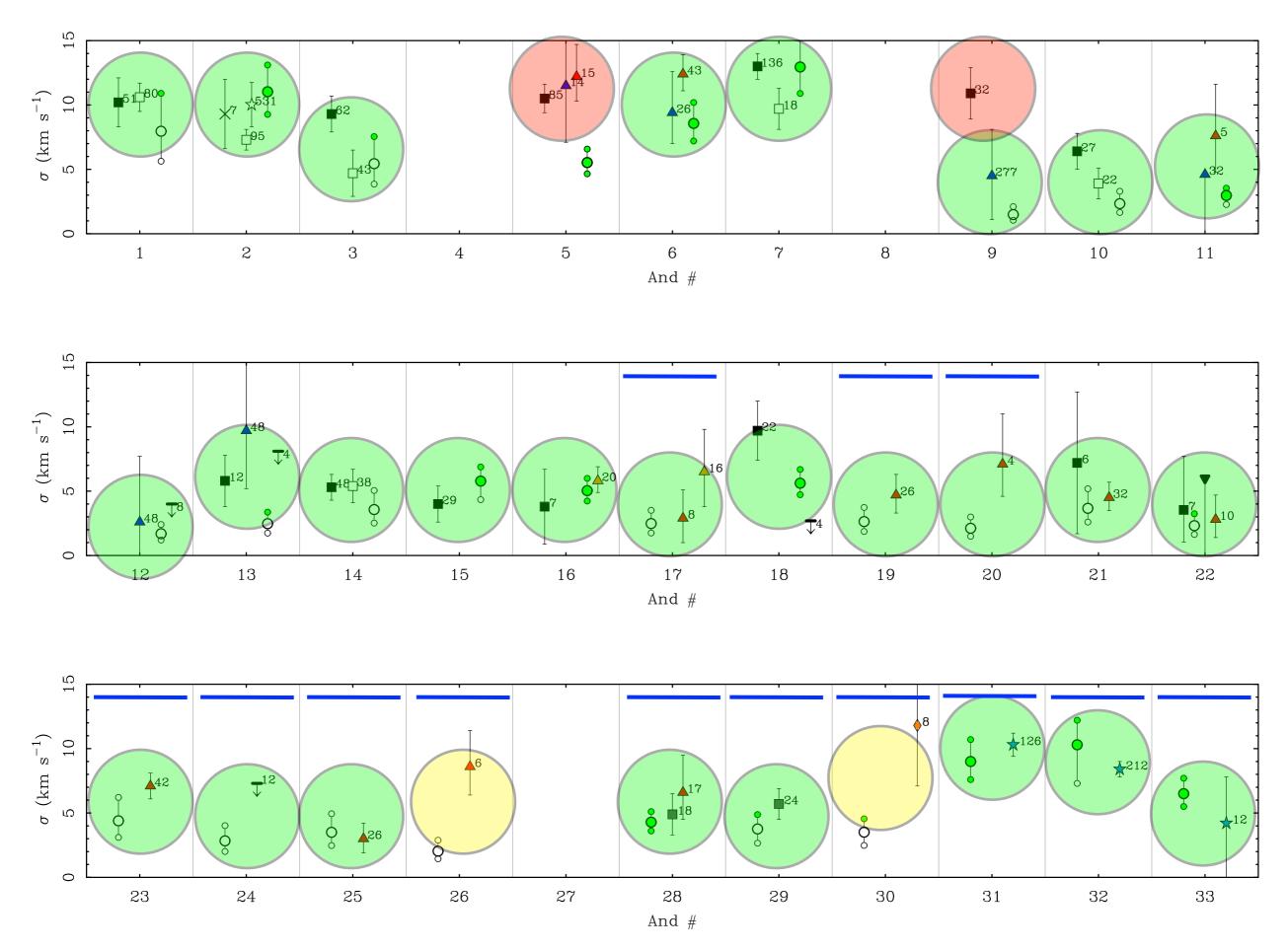
$$M = \frac{g_{ex}}{a_0} \frac{RV^2}{G}$$

e.g., nearby dwarf Segue 1

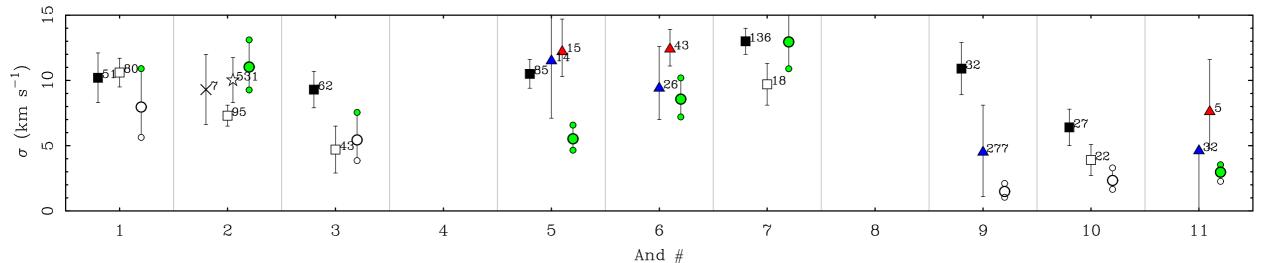




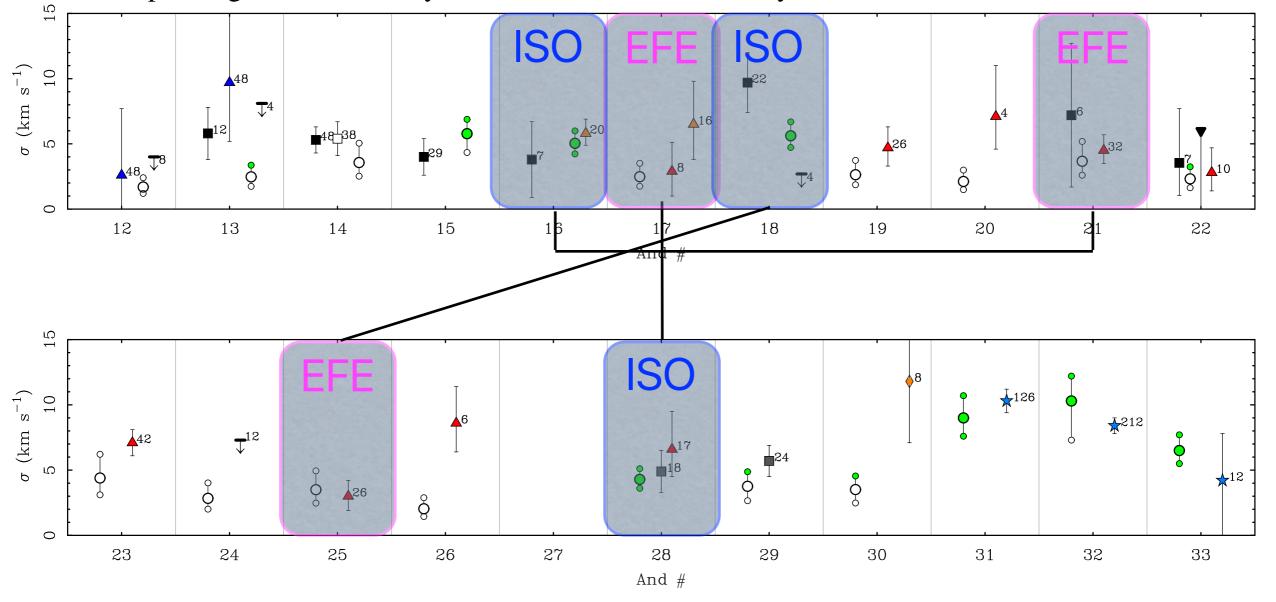
Velocity dispersions of M31 dwarfs correctly predicted (a priori in many cases) by MOND.



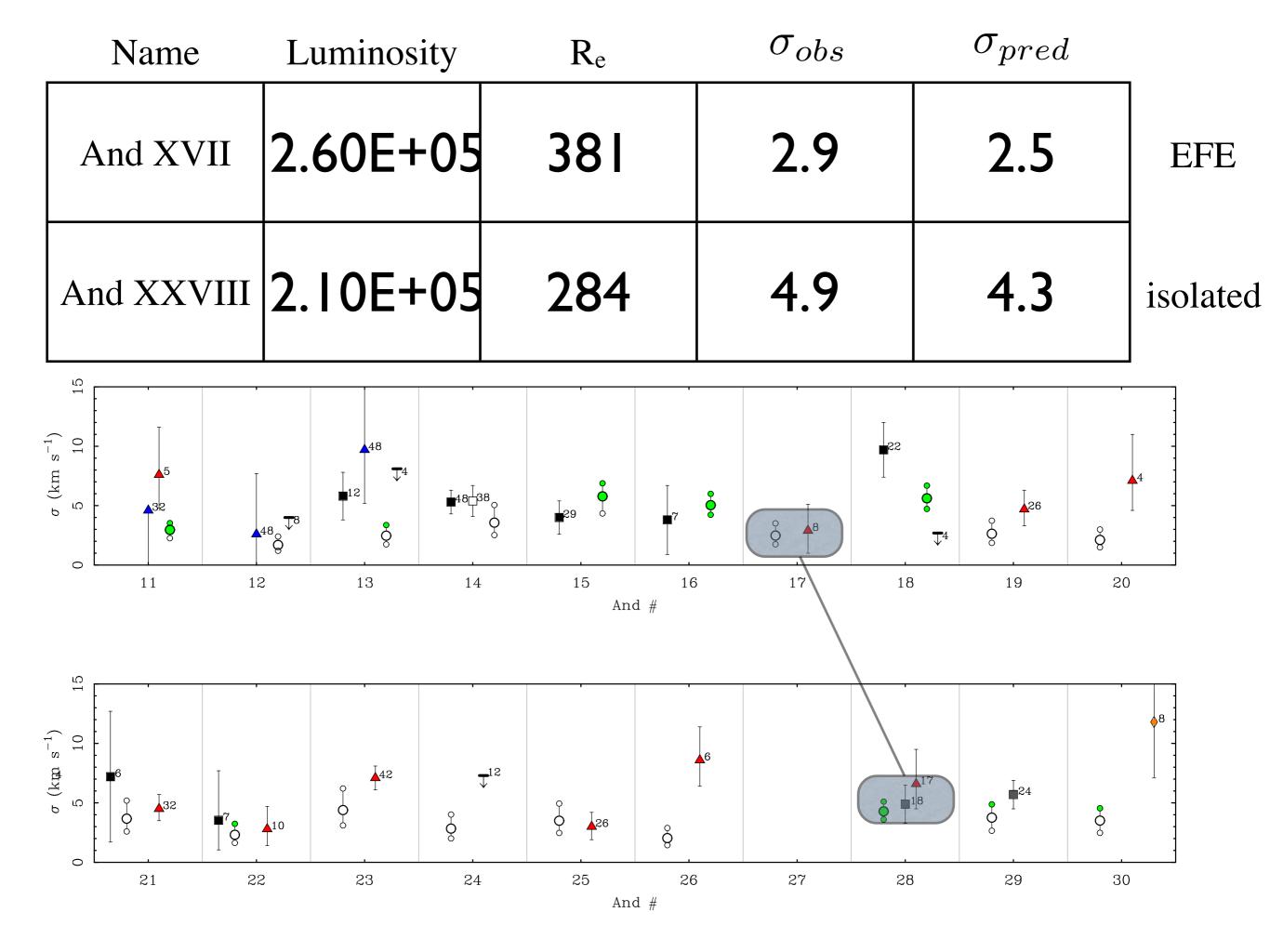
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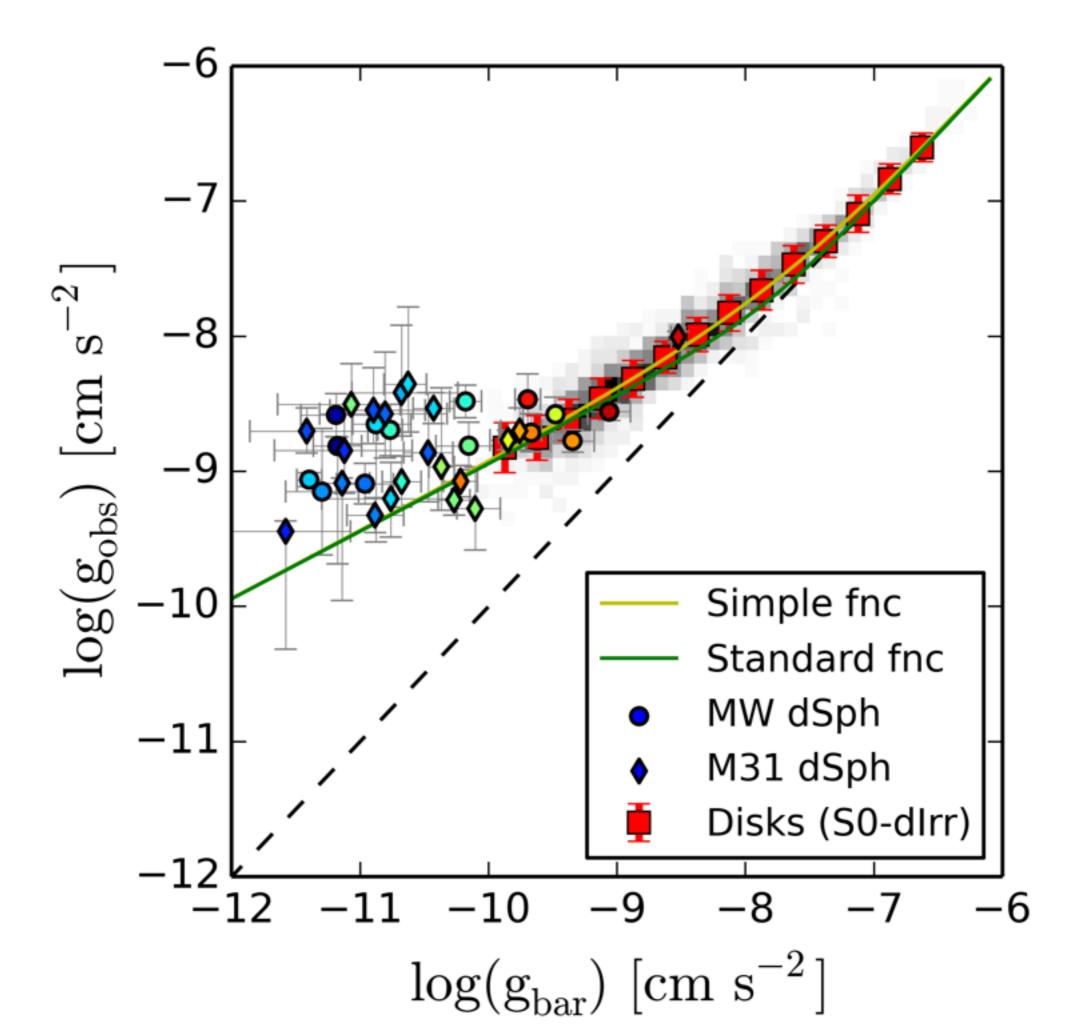


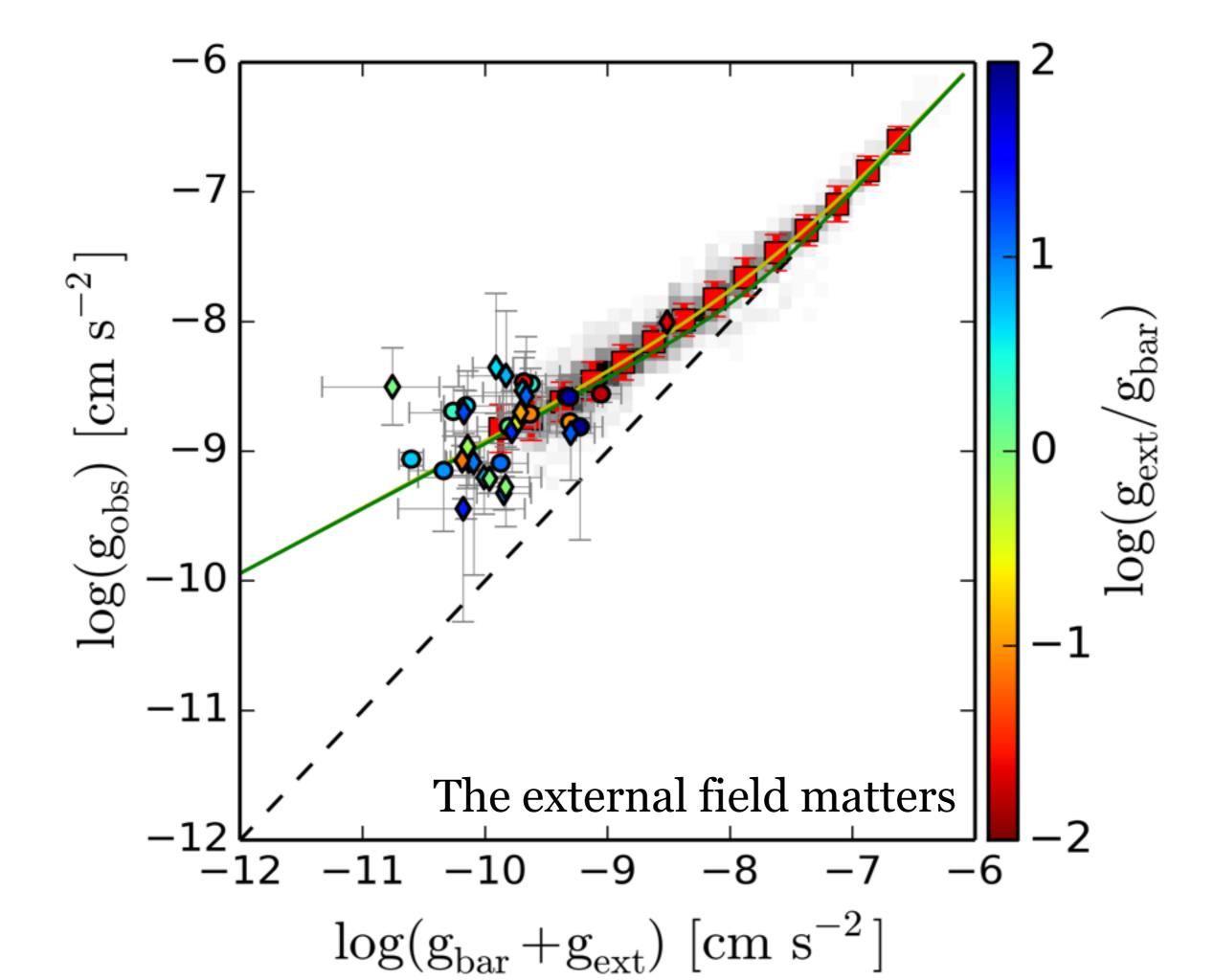
Pairs of photometrically identical dwarfs should have different velocity dispersion depending on whether they are isolated are dominated by the external field effect.



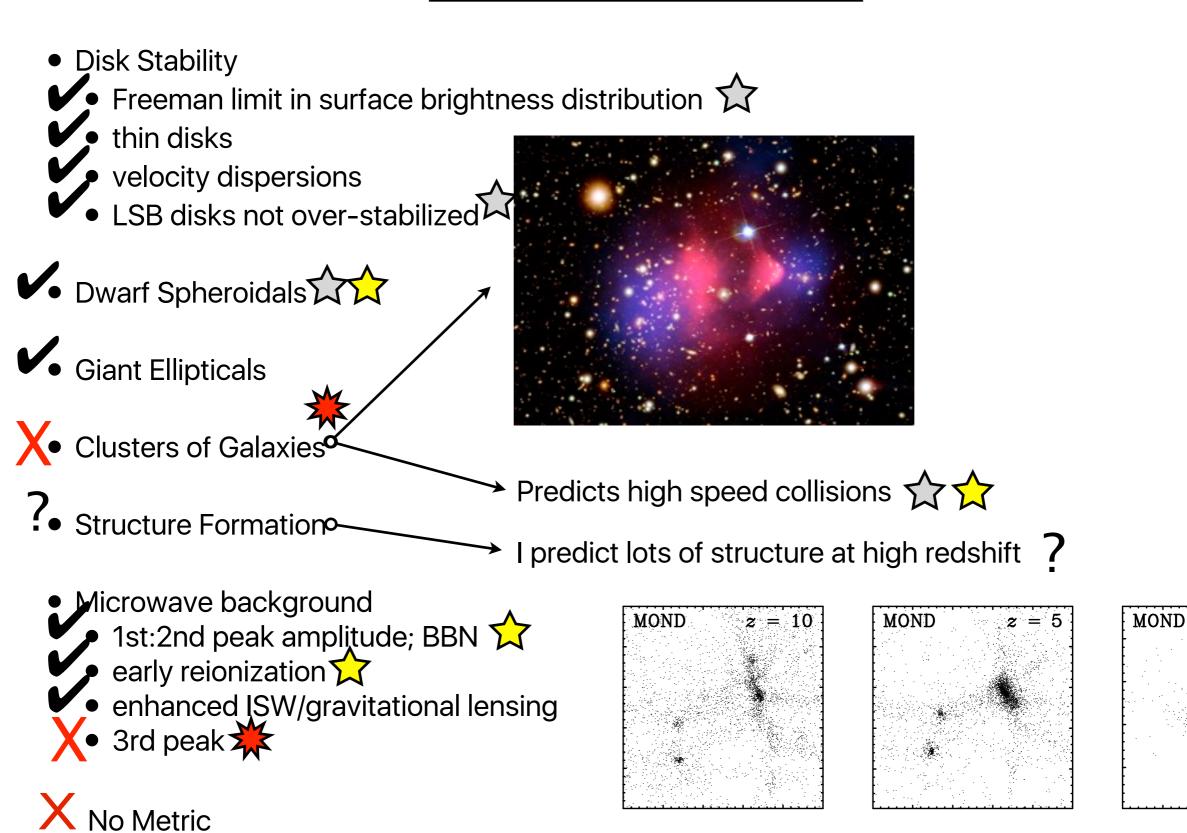
There is no EFE in dark matter - this is a unique signature of MOND.







Other MOND tests



Don't know expansion history



the seeming novelty and singularity of this opinion can be no sufficient reason to prove it erroneous

